

Article I: Reaction of  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  ions with  $\text{H}_2$   
at low temperatures-The pathway to ammonia  
production in the interstellar medium

Rednyk, S., Roučka, Š., Kovalenko, A., Tran, T. D., Dohnal, P.,  
Plašil, R., & Glosík, J.

2019

*Astronomy & Astrophysics*, 625(A74), 1-8

# Reaction of $\text{NH}^+$ , $\text{NH}_2^+$ , and $\text{NH}_3^+$ ions with $\text{H}_2$ at low temperatures

## The pathway to ammonia production in the interstellar medium

S. Rednyk, Š. Roučka, A. Kovalenko, T. D. Tran, P. Dohnal, R. Plašil, and J. Glošík

Department of Surface and Plasma Science, Faculty of Mathematics and Physics, Charles University, V Holešovičkách 2, 180 00 Prague, Czech Republic  
e-mail: stepan.roucka@mff.cuni.cz

Received 28 August 2018 / Accepted 23 March 2019

### ABSTRACT

**Aims.** We present an experimental investigation of the exothermic reactions of  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  ions with  $\text{H}_2$  at temperatures relevant for interstellar clouds.

**Methods.** The reactions were studied using a variable-temperature 22-pole radio frequency ion trap instrument.

**Results.** The temperature dependences of rate coefficients of these reactions have been obtained at temperatures from 15 up to 300 K. The reaction of  $\text{NH}^+$  with  $\text{H}_2$  has two channels, which lead to  $\text{NH}_2^+$  (~97%) and  $\text{H}_3^+$  (~3%) with nearly constant reaction rate coefficients ( $k_{\text{NH}^+}^a(17\text{ K}) = 1.0 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  and  $k_{\text{NH}^+}^b(17\text{ K}) = 4.0 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$ , respectively). The reaction of  $\text{NH}_2^+$  with  $\text{H}_2$  produces only  $\text{NH}_3^+$  ions. The measured rate coefficient monotonically decreases with increasing temperature from  $k_{\text{NH}_2^+}(17\text{ K}) = 6 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  to  $k_{\text{NH}_2^+}(300\text{ K}) = 2 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ . The measured rate coefficient of the reaction of  $\text{NH}_3^+$  with  $\text{H}_2$ , producing  $\text{NH}_4^+$ , increases with decreasing temperature from 80 K down to 15 K, confirming that the reaction proceeds by tunnelling through a potential barrier.

**Key words.** astrochemistry – molecular data – molecular processes – methods: laboratory: molecular – ISM: molecules

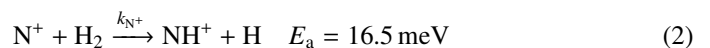
## 1. Introduction

Nitrogen is among the six most abundant elements in the Universe. Owing to the dominance of hydrogen in the Universe, it is not surprising that nitrogen hydrides are present in the interstellar medium (ISM). Ammonia was the first polyatomic molecule discovered in interstellar space (Cheung et al. 1968). Later, other nitrogen hydrides were detected: NH by Meyer & Roth (1991),  $\text{NH}_2$  by van Dishoeck et al. (1993), and  $\text{NH}_3\text{D}^+$  by Cernicharo et al. (2013). Atomic and molecular nitrogen were also observed in the ISM (Nieva & Przybilla 2012; Knauth et al. 2004). For information on recent observations of nitrogen hydrides, see the results from the *Herschel*/HIFI instrument (Persson et al. 2010, 2012; Caselli et al. 2017), and from the SOFIA instrument (Wyrowski et al. 2016). For a detailed description of interstellar chemistry of nitrogen hydrides, see for example Rist et al. (2013) and Harju et al. (2017) and the reviews by Le Gal et al. (2014), Gerin et al. (2016), and Acharyya & Herbst (2015). Despite some differences between observed values and the modelled predictions of populations of nitrogen hydrides (see e.g. Le Gal et al. 2014; Persson et al. 2012; Novotný et al. 2014 and references therein), it is generally accepted that the main pathway to gas-phase formation of ammonia in the ISM is a chain of hydrogen abstraction reactions followed by the dissociative recombination of  $\text{NH}_4^+$  (Herbst & Klemperer 1973; Le Gal et al. 2014; Gerin et al. 2016). The suggested pathway of the gas-phase formation of the  $\text{NH}_4^+$  in the ISM starting from  $\text{N}^+$  is (Le Gal et al. 2014)



In current understanding of nitrogen chemistry, the production of  $\text{N}^+$  in the low-temperature (10 K) dark clouds results from dissociative ionization of  $\text{N}_2$  in reaction with  $\text{He}^+$  (Hily-Blant et al. 2013; Le Gal et al. 2014). Alternatively, the chain may also be initiated by  $\text{N} + \text{H}_3^+ \rightarrow \text{NH}_2^+ + \text{H}$ , although this reaction has a high activation energy (Herbst et al. 1987; Scott et al. 1997; Le Gal et al. 2014). It is expected that neutral  $\text{NH}_3$  molecules are consequently formed in the dissociative recombination of  $\text{NH}_4^+$  ions with electrons. Other hydrides can also be formed by the recombination of ions from this sequence with electrons (Florescu-Mitchell & Mitchell 2006). To model the production of ammonia in the ISM it is important to know the rate coefficients of all the reactions in the chain (1) for temperatures down to 10 K.

The first binary ion-molecule reaction of the chain is



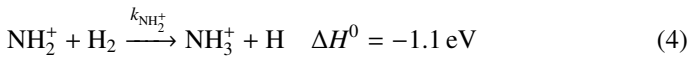
with a corresponding reaction rate coefficient denoted as  $k_{\text{N}^+}$ . Since the enthalpy of this reaction is not yet known with sufficient precision, we provide the activation energy  $E_a$  obtained by Zymak et al. (2013). The enthalpies of the other hydrogen abstraction reactions were taken from Rist et al. (2013). This endothermic reaction of the  $\text{N}^+$  ion with molecular hydrogen has been studied using several well-established experimental techniques (e.g. Adams & Smith 1985; SIFDT, Marquette et al. 1988; CRESU, and Gerlich 1993; 22-pole ion trap). It has also been studied in our laboratory using the 22-pole ion trap instrument with consideration of para- and ortho-spin configurations of the reacting hydrogen molecule (Zymak et al. 2013; Plašil et al. 2014).

The next reaction of the chain (1) is



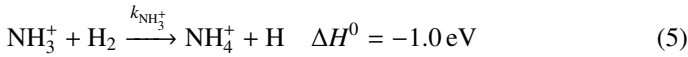
where  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  are reaction rate coefficients of channels with production of  $\text{NH}_2^+$  and  $\text{H}_3^+$ , respectively. We also define the overall rate coefficient of the reaction of  $\text{NH}^+$  ion with  $\text{H}_2$  as  $k_{\text{NH}^+} = (k_{\text{NH}^+}^a + k_{\text{NH}^+}^b)$ . The enthalpy of the proton transfer channel (3b) was calculated from the corresponding proton affinities (Hunter & Lias 1998). Reaction (3) was studied at 300 K by Fehsenfeld et al. (1967; FA), Kim et al. (1975; ICR), and Adams et al. (1980; SIFT). At 15 K, it was studied by Gerlich (1993; ion trap). The production of  $\text{H}_3^+$  ion (reaction (3b)) was only reported by Adams et al. (1980).

The produced  $\text{NH}_2^+$  ion further reacts with  $\text{H}_2$  in reaction



with a reaction rate coefficient  $k_{\text{NH}_2^+}$ . This reaction was studied by Fehsenfeld et al. (1967; FA), Kim et al. (1975; ICR), and Adams et al. (1980; SIFT) at 300 K. Using an ion trap instrument, the rate coefficient of the reaction (4) was measured at 15 K by Gerlich (1993).

The final reaction of chain (1) is



with a rate coefficient  $k_{\text{NH}_3^+}$ . This reaction was studied experimentally over broad range of temperatures by Fehsenfeld et al. (1975), Kim et al. (1975), Smith & Adams (1981), Luine & Dunn (1985), Böhringer (1985), Barlow & Dunn (1987), Adams & Smith (1984), and Gerlich (1993). It was also studied theoretically by Herbst et al. (1991) and recently by Álvarez-Barcia et al. (2016). At temperatures above 300 K, the measured temperature dependence of the rate coefficient of the exothermic reaction (5) exhibits dependence typical for endoergic reactions. If the temperature dependence of  $k_{\text{NH}_3^+}$  measured at temperatures above 300 K is extrapolated towards lower temperatures relevant for interstellar clouds using Arrhenius dependence (Fehsenfeld et al. 1975), then the values will be far too low to explain the observed  $\text{NH}_3$  abundances. This problem was solved when measurements at temperatures below 100 K indicated that the temperature dependence of the reaction rate coefficient has a local minimum, and then slowly increases with temperature decreasing below 50 K. These temperature dependences of rate coefficients of ion-molecule reactions with minimum have been observed several times (e.g. Smith & Adams 1981) and they are typical of exothermic reactions proceeding by tunnelling through a potential barrier (Ng et al. 1994). The observed increase in the reaction rate coefficient  $k_{\text{NH}_3^+}$  at low temperatures due to the tunnelling is sufficient for reaction (5) to play an important role in interstellar chemistry.

The recent observations of nitrogen hydrides in many areas of the ISM have led to the modelling of their production and destruction in the corresponding environments. This requires the knowledge of the rate coefficients of reactions playing a role in the production and destruction of nitrogen hydrides at temperatures down to 10 K. The present contribution reports the results of the studies of the reactions of ions  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  with  $\text{H}_2$  at temperatures from 15 to 300 K using a 22-pole radio frequency (RF) ion trap. After a brief description of the instrument and typical measuring procedures, new data including measured

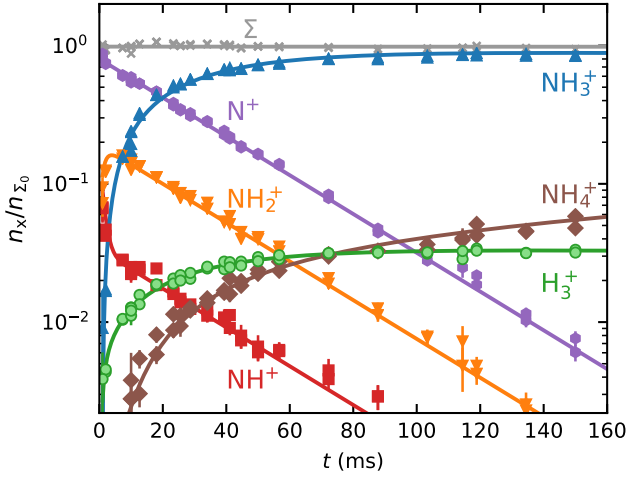
temperature dependences of the reaction rate coefficients ( $k_{\text{NH}^+}^a$ ,  $k_{\text{NH}^+}^b$ ,  $k_{\text{NH}_2^+}$ , and  $k_{\text{NH}_3^+}$ ) are presented. The new results are compared with values from previous experiments and with available theoretical predictions.

## 2. Experiment and data analysis

The experiments were carried out by means of a linear RF 22-pole ion trap instrument operating at temperatures ( $T_{22\text{PT}}$ ) from 10 to 300 K. The basics of storing ions in RF fields have been described thoroughly by Gerlich (1992, 1995). Only a very short description of the apparatus and its operation are given here (for details, see e.g. Gerlich 1992; Gerlich et al. 2011; Plašil et al. 2011; Zymak et al. 2013). The ion trap is surrounded by a copper box, which is mounted onto a cold head of a closed-cycle helium refrigerator. Helium and hydrogen are introduced into the trap via leak valves and optionally in short pulses ( $\sim 10$  ms) via a piezo valve (Gerlich 2008).

The primary reactant ions are produced by electron bombardment of the precursor gas in a storage ion source (SIS). In the present experiments, a mixture of  $\text{N}_2$  and  $\text{H}_2$  was used as a source gas. The produced ions were periodically extracted from the ion source and mass selected with a quadrupole mass filter. The mass-selected (primary) ions were transferred into the 22-pole ion trap, filled with a mixture of He buffer gas and  $\text{H}_2$  reactant gas. The helium number density used in the experiments was in the range of  $10^{13}$ – $10^{14}$   $\text{cm}^{-3}$  during the measurements and it was temporarily increased to  $10^{15}$   $\text{cm}^{-3}$  during the injection of ions into the trap by adding He via the piezo valve. In the present experiments, normal hydrogen was used as a reactant gas with number densities in the trap varying from  $10^{10}$  up to  $10^{12}$   $\text{cm}^{-3}$ . In normal  $\text{H}_2$ , the para/ortho ratio is 1/3, corresponding to the thermal equilibrium at 300 K (for discussion, see Zymak et al. 2013; Hejduk et al. 2012). As the rotational excitation in hydrogen gas is thermalized only within para and ortho manifolds,  $\text{H}_2$  gas flowing through the gas inlet system into the trap volume is not thermalized at trap temperatures below 200 K (Zymak et al. 2013; Hejduk et al. 2015). The gas number density inside the ion trap is determined using a spinning rotor gauge and a calibrated ionization gauge with estimated uncertainty of 20%. This constant systematic uncertainty is not included in the error bars of our figures, which indicate the relative uncertainties.

At the number densities of He buffer gas and  $\text{H}_2$  reactant gas in the present experiment, the kinetic energy of the injected ion was cooled by hundreds of collisions with He atoms prior to colliding with  $\text{H}_2$ . After various trapping times, the ions were extracted from the trap and after passing through a second quadrupole mass filter, they were detected with an MCP detector. The standard measuring procedure is based on filling the ion trap at a fixed frequency with a well-defined number of primary ions and by analysing the content of the ions in the ion trap after different trapping (reaction) times. The data are analysed under the assumption that the numbers of detected ions are proportional to the numbers of ions in the ion trap. In the following text these relative numbers of different ions of particular mass detected (counted) after trapping time  $t$  are denoted  $n_x(t)$ , where the index  $x$  refers to the various ions in the ion trap. For easier comparison of the experimental results, the data plotted in the figures were normalized by dividing by the total number of detected ions  $n_{\Sigma 0} = \sum n_x(t_0)$ , where  $t_0$  is the shortest trapping time. The symbol  $\Sigma(t)$  in the figures indicates the normalized total number of ions in the trap,  $\Sigma(t) = \sum n_x(t)/n_{\Sigma 0}$ . The mass discrimination of the detection system is considered in the

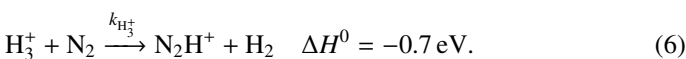


**Fig. 1.** Measured time evolutions of normalized numbers ( $n_x(t)/n_{x0}$ ) of  $\text{N}^+$ ,  $\text{NH}^+$ ,  $\text{NH}_2^+$ ,  $\text{NH}_3^+$ ,  $\text{NH}_4^+$ , and  $\text{H}_3^+$  ions in the trap filled with  $\text{He}/\text{H}_2$  gas mixture after the injection of  $\text{N}^+$  ions from the SIS. The collisional temperature is  $T = 16$  K, hydrogen number density is  $[\text{H}_2] = 5.7 \times 10^{10} \text{ cm}^{-3}$  and helium number density is  $[\text{He}] = 5.6 \times 10^{13} \text{ cm}^{-3}$ . The fitted solution of the corresponding system of rate equations is indicated with the solid lines. The crosses ( $x$ ) indicate the sum of the normalized numbers ( $\Sigma(t)$ ) of all ions in the ion trap.

data analysis (see e.g. Kovalenko et al. 2018, for details) and for some ions ( $\text{H}_3^+$  in comparison with  $\text{NH}^+$ ), it was calibrated using a suitable chemical reaction (see Sect. 3.1).

In recent experiments with our apparatus (Zymak et al. 2013; Plašil et al. 2012; Roučka et al. 2018) and in other 22-pole ion trap experiments (Hauser et al. 2015; Endres et al. 2017) it has been confirmed that the collisional temperature (here denoted simply  $T$ ) is slightly higher than the temperature of the copper box surrounding the ion trap (nominal ion trap temperature,  $T_{22\text{PT}}$ ). At the present experimental conditions, we can safely assume that the collisional temperature in the interaction of ions with  $\text{H}_2$  does not exceed the ion trap temperature by more than 10 K. For simplicity of presentation, we define the collisional temperature as  $T = T_{22\text{PT}} + 5$  K with an uncertainty of  $\pm 5$  K.

To illustrate the character of the results from the ion trap experiment, the typical data measured in an experiment with injection of  $\text{N}^+$  ions into the ion trap filled with He buffer gas and  $\text{H}_2$  reactant gas are shown in Fig. 1. Plotted are the time evolutions of normalized numbers ( $n_x(t)/n_{x0}$ ) of  $\text{N}^+$ ,  $\text{NH}^+$ ,  $\text{NH}_2^+$ ,  $\text{NH}_3^+$ ,  $\text{NH}_4^+$ , and  $\text{H}_3^+$  ions in the ion trap. The mass discriminations for particular ions are considered in the evaluation of the measured data. Owing to the large differences between the values of the rate coefficients of the reactions in chain (1), the data were collected in small time steps over a broad time interval. The dominant process after the injection of  $\text{N}^+$  ions is their reaction with  $\text{H}_2$  in which the  $\text{NH}^+$  ions are formed. The  $\text{NH}^+$  ions further react with  $\text{H}_2$ , producing  $\text{NH}_2^+$  and  $\text{H}_3^+$ . The  $\text{NH}_2^+$  ions consequently react with  $\text{H}_2$  to produce  $\text{NH}_3^+$ . Finally, the  $\text{NH}_4^+$  ions are produced in slow reactions of  $\text{NH}_3^+$  with  $\text{H}_2$ . From the time evolution of the normalized number of  $\text{H}_3^+$  ions it can be seen that at very long trapping times  $\text{H}_3^+$  ions are slowly removed by reactions with  $\text{N}_2$  (see e.g. Marquette et al. 1989) that penetrates into the trap volume from the ion source:



At low temperatures ( $\lesssim 40$  K), this process becomes negligible because the number density of  $\text{N}_2$  in the trap is reduced by condensation on the walls of the trap.

The time evolution of the numbers of ions in the ion trap after injection of  $\text{N}^+$  ions can be described by the following set of differential balance equations, which can be derived from the chemical Eqs. (2)–(6):

$$\frac{dn_{\text{N}^+}}{dt} = -r_{\text{N}^+} n_{\text{N}^+} \quad (7)$$

$$\frac{dn_{\text{NH}^+}}{dt} = r_{\text{N}^+} n_{\text{N}^+} - n_{\text{NH}^+} (r_{\text{NH}^+}^a + r_{\text{NH}^+}^b) \quad (8)$$

$$\frac{dn_{\text{H}_3^+}}{dt} = r_{\text{NH}^+}^b n_{\text{NH}^+} - r_{\text{H}_3^+} n_{\text{H}_3^+} \quad (9)$$

$$\frac{dn_{\text{NH}_2^+}}{dt} = r_{\text{NH}^+}^a n_{\text{NH}^+} - r_{\text{NH}_2^+} n_{\text{NH}_2^+} \quad (10)$$

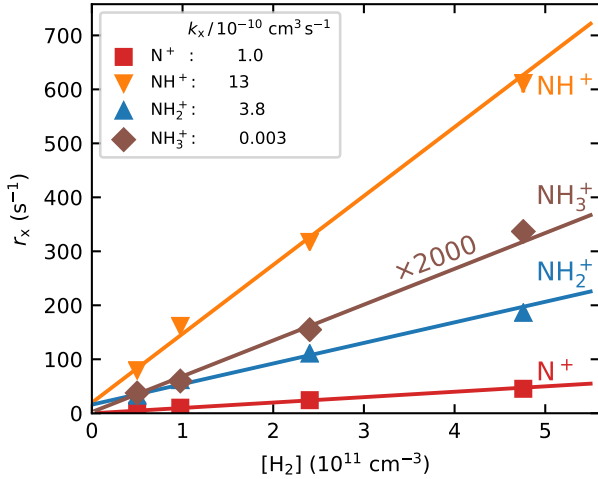
$$\frac{dn_{\text{NH}_3^+}}{dt} = r_{\text{NH}_2^+} n_{\text{NH}_2^+} - r_{\text{NH}_3^+} n_{\text{NH}_3^+} \quad (11)$$

$$\frac{dn_{\text{NH}_4^+}}{dt} = r_{\text{NH}_3^+} n_{\text{NH}_3^+}. \quad (12)$$

The reactions are parametrized by the reaction rates  $r_x$ , which are proportional to the reaction rate coefficients and the number densities of the neutral reactants, i.e.  $r_x = k_x[\text{H}_2]$  in the case of reactions with  $\text{H}_2$  and  $r_{\text{H}_3^+} = k_{\text{H}_3^+}[\text{N}_2]$  in Eq. (9). The rate coefficients of the reactions (2)–(6) can therefore be determined by fitting the solution of the corresponding set of differential rate equations to the experimental data with the reaction rates and initial numbers of ions as free parameters. From the fitted rates  $r_x$  at given hydrogen density  $[\text{H}_2]$ , we calculate the corresponding reaction rate coefficients as  $k_x = r_x/[\text{H}_2]$ . An example of a fitted solution is also shown in Fig. 1.

In measurements with  $\text{NH}^+$  and  $\text{NH}_2^+$  primary ions, the system can be simplified by setting  $n_{\text{N}^+}$  and  $(n_{\text{N}^+}, n_{\text{NH}^+}, n_{\text{H}_3^+})$  to zero, respectively. Special attention should be paid to the possible excitation of the intermediate ions produced in the sequence of reactions in trap. Details concerning the studies of specific reactions are discussed in Sect. 3.

To confirm that the observed reaction rates are indeed caused by binary reactions with  $\text{H}_2$  and to evaluate the possible loss of ions due to other (background) processes, we measured the time evolutions of the relative numbers of the ions at several number densities of hydrogen in the ion trap. In these particular experiments,  $\text{N}^+$  ions were injected into the ion trap (see example in Fig. 1). The examples of the dependences of the reaction loss rates ( $r_x$ ) on  $[\text{H}_2]$  for the reactions (2)–(5) measured at  $T = 18$  K are shown in Fig. 2. The values of  $r_{\text{NH}_3^+}$  were obtained from the fits of the measured time evolutions of relative number of ions at long trapping times, i.e. from evolutions of  $n_{\text{NH}_3^+}$  and  $n_{\text{NH}_4^+}$  (not shown in Fig. 1; see Fig. 9 below). The linearity of the dependences plotted in Fig. 2 confirms that the time evolutions of the relative numbers of particular ions in the ion trap are controlled by binary ion-molecule reactions with  $\text{H}_2$ . The measured loss rates can be expressed by the formula,  $r_x = k_x[\text{H}_2] + r_{\text{xbg}}$ , where  $r_{\text{xbg}}$  is the background loss rate for particular ions. The corresponding binary reaction rate coefficients  $k_{\text{N}^+}$ ,  $k_{\text{NH}^+}$ ,  $k_{\text{NH}_2^+}$ , and  $k_{\text{NH}_3^+}$  are given by the slope of the plotted dependences. The value of  $k_{\text{N}^+}(18 \text{ K}) = (1.0 \pm 0.4) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  obtained from the data plotted in Fig. 2 is in very good agreement with the values obtained in previous studies (Zymak et al. 2013). The values of the other reaction rate coefficients obtained from the data in Fig. 2 are discussed below.



**Fig. 2.** Measured reaction loss rates  $r_{N^+}$ ,  $r_{NH^+}$ ,  $r_{NH_2^+}$ , and  $r_{NH_3^+}$  of  $N^+$ ,  $NH^+$ ,  $NH_2^+$ , and  $NH_3^+$  ions, respectively, as a function of  $H_2$  number density. The primary  $N^+$  ions are injected into the ion trap from the SIS. The collisional temperature is  $T = 18$  K and helium number density is  $[He] = 8 \times 10^{13} \text{ cm}^{-3}$ . For the relatively slow reaction of  $NH_3^+$  ions, the plotted values of  $r_{NH_3^+}$  are multiplied by a factor of 2000. The corresponding binary reaction rate coefficients  $k_{N^+}$ ,  $k_{NH^+}$ ,  $k_{NH_2^+}$ , and  $k_{NH_3^+}$  given by the slope of the plotted dependences are indicated in the legend.

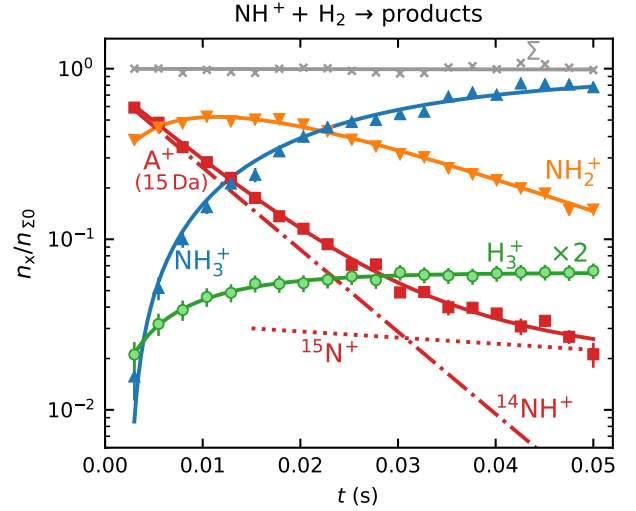
In order to increase the accuracy of the measured reaction rate coefficients for reactions of particular ions, these ions are produced in the SIS and injected into the ion trap. The corresponding reaction rate coefficient can then be calculated from the decay (time evolution) of the relative number of this particular ion. The obvious advantage of the determination of the rate coefficient from the decrease in the relative number of the studied ion is that the decrease in the relative number of the injected ions is influenced only by the reaction of particular ions with the reactant gas.

### 3. Results and discussion

#### 3.1. Reaction $NH^+ + H_2$

We measured the rate coefficients for the atom abstraction reaction (3a) and for the proton transfer reaction (3b) at the ion trap temperatures from 10 up to 130 K. In these experiments,  $NH^+$  ions were produced by electron bombardment of the mixture of  $H_2$  and  $N_2$  gases in the SIS and injected into the trap. The typical time evolutions of the normalized numbers of primary and product ions are shown in Fig. 3.

Since the mass filters in our instrument cannot resolve between  $^{14}NH^+$  and the  $^{15}N^+$  isotope, we also injected a small fraction of  $^{15}N^+$ . This mixture of  $^{14}NH^+$  and  $^{15}N^+$  ions (mass 15 Da) is denoted  $A^+(15 \text{ Da})$ . The relative populations of  $^{14}NH^+$  and  $^{15}N^+$  ions injected to the ion trap are influenced by the natural abundance of  $^{15}N$  in  $N_2$  and by kinetics in the ion source at particular conditions (partial pressures of gases, electron energy, and storage time in the SIS). The decrease in the normalized number of  $A^+(15 \text{ Da})$  is exponential (dash-dotted straight line in semi-log plot) with a leveling at  $\sim 0.025$  (dotted line indicated as  $^{15}N^+$ ). This leveling of the  $A^+(15 \text{ Da})$  numbers was not observed when  $^{14}NH^+$  ions were produced directly in the trap in the reaction of  $H_2$  with  $^{14}N^+$  ions (with well-known isotopic composition; see the example in Fig. 1). This confirms that the slowly reacting ions are  $^{15}N^+$  ions. The presence of  $^{15}N^+$  is considered



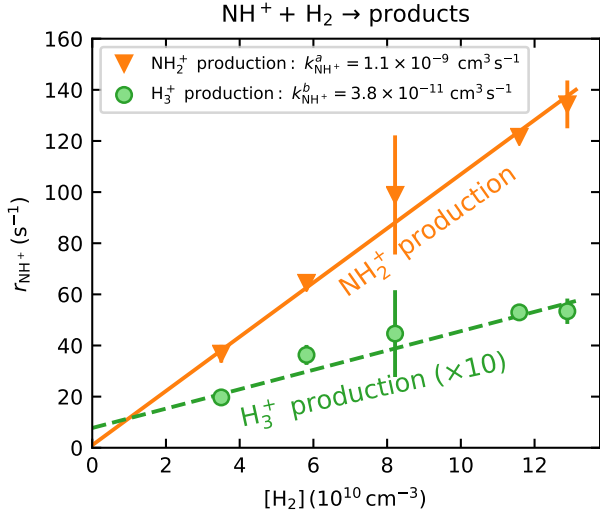
**Fig. 3.** Measured time evolutions of the normalized numbers ( $n_x(t)/n_{\Sigma 0}$ ) of indicated ions in the trap.  $A^+(15 \text{ Da})$  indicates a normalized number of ions with mass 15 Da (mixture of  $^{14}NH^+$  and  $^{15}N^+$  ions). The measurements were performed at  $T = 34$  K,  $[H_2] = 6.2 \times 10^{10} \text{ cm}^{-3}$ , and  $[He] \approx 10^{14} \text{ cm}^{-3}$ . The solid lines are least-squares fits, which were used to determine the reaction rate coefficients. The dash-dotted straight line indicates the decrease in the normalized numbers of  $^{14}NH^+$  ions and the dotted line indicates the normalized numbers of  $^{15}N^+$  ions. The crosses ( $\times$ ) indicate the sum of the normalized numbers of all ions in the ion trap,  $\Sigma(t)$ . The normalized numbers of  $H_3^+$  ions are increased by a factor of 2.

in the data analysis. Owing to its small rate coefficient, the reaction of  $^{15}N^+$  with  $H_2$  does not influence the evaluation of the rate coefficient  $k_{NH^+}$ . Nevertheless, we take it into account, assuming that it has the same rate coefficient as the  $^{14}NH^+ + H_2$  reaction (Zymak et al. 2013).

From the data plotted in Fig. 3, we can see the decrease in the relative number of  $NH^+$  ions and the production of  $NH_2^+$  and  $H_3^+$  ions. To obtain the branching ratio for the reactions (3a) and (3b), the ion detection system had to be calibrated. We obtained the discrimination of the detection system between ions of mass 3 Da and 16 Da by measuring the discrimination between mass 3 Da and 17 Da using a calibration reaction of  $H_3^+ + CH_4 \rightarrow CH_5^+ + H_2$  (Bohme et al. 1980) and by taking into account the discrimination between mass 16 Da and 17 Da known from the present experiment with reaction  $NH_2^+ + H_2 \rightarrow NH_3^+ + H$ .

We verified the binary character of the reactions (3a) and (3b) at the present experimental conditions by measuring the dependences of the reaction rates  $r_{NH^+}^a$  and  $r_{NH^+}^b$  on hydrogen number density. The examples of the dependences measured at temperature  $T = 20$  K are shown in Fig. 4. The number density of  $H_2$  leaking from the ion source has also been measured and is accounted for in our figures.

By fitting the measured time evolutions of the relative numbers of ions in the ion trap, the reaction rate coefficients  $k_{NH^+}^a$  and  $k_{NH^+}^b$  for the reaction channel (3a) and (3b) were obtained. The temperature dependences of the reaction rate coefficients  $k_{NH^+}^a$  and  $k_{NH^+}^b$  are shown in Fig. 5. Also plotted is the value of  $k_{NH^+}$  measured at 300 K by Kim et al. (1975). There is just one value of the reaction rate coefficient  $k_{NH^+}$  measured in the ion trap experiment at 15 K by Gerlich (1993). In the ion trap experiment, Gerlich measured the overall reaction rate coefficient  $k_{NH^+}$  (see Fig. 5). Production of  $H_3^+$  ions was observed only in the selected ion flow tube (SIFT) studies of Adams et al. (1980). At 300 K, they observed 85% of  $NH_2^+$  and 15% of  $H_3^+$



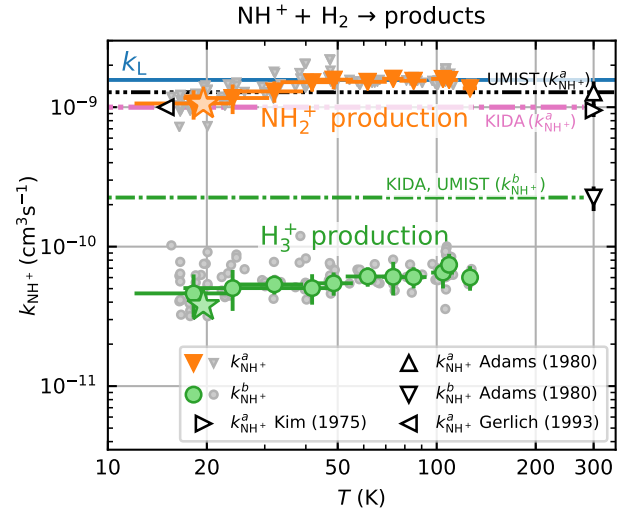
**Fig. 4.** Measured dependences of the reaction rates  $r_{\text{NH}^+}^a$  and  $r_{\text{NH}^+}^b$ , corresponding to the production of  $\text{NH}_2^+$  and  $\text{H}_3^+$  ions, on  $\text{H}_2$  number density. The plotted values of  $r_{\text{NH}^+}^b$  (corresponding to  $\text{H}_3^+$  production) are increased by a factor of 10. The mean helium number density was of the order of  $10^{14} \text{ cm}^{-3}$ . The collisional temperature is  $T = 20 \text{ K}$ . In these experiments, the primary  $\text{NH}^+$  ions were produced in the SIS and injected into the ion trap. The values of the binary reaction rate coefficients  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  obtained from the slopes of the corresponding dependences are shown in the legend.

products. The corresponding values of  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  are plotted in Fig. 5. If we extrapolate our data from 130 K towards 300 K as a constant, then there is good agreement with the previously obtained values of  $k_{\text{NH}^+}$  by Kim et al. (1975) and  $k_{\text{NH}^+}^a$  by Adams et al. (1980). However, the value of  $k_{\text{NH}^+}^b$  for the production of  $\text{H}_3^+$ , measured by Adams et al. (1980) is almost a factor of 4 higher than the present value at 130 K. Values recommended by the Kinetic Database for Astrochemistry (KIDA, Wakelam et al. 2012) and by the University of Manchester Institute of Science and Technology (UMIST) Database for Astrochemistry (McElroy et al. 2013) are also included in Fig. 5.

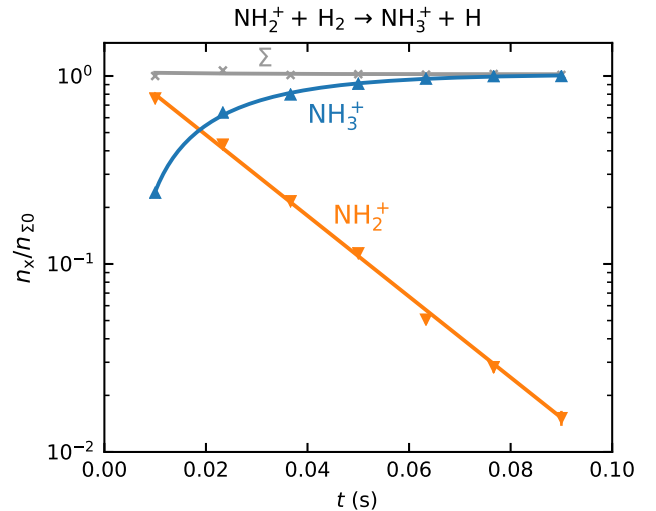
There have been several previous studies of reaction of  $\text{NH}^+$  with  $\text{H}_2$ ; however, this is the first study giving the temperature dependences of the reaction rate coefficients  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  from 15 up to 130 K, i.e. for astrophysically relevant temperatures.

### 3.2. Reaction $\text{NH}_2^+ + \text{H}_2$

As mentioned above, in the experiments where  $\text{N}^+$  ions were injected into the ion trap and  $\text{NH}_2^+$  ions were produced in the chain of reactions with  $\text{H}_2$ , we observed the dependence of the measured reaction rate coefficient  $k_{\text{NH}_2^+}$  on partial pressures of gases in the ion trap and on the trapping time. This is presumably connected with internal excitation of  $\text{NH}_2^+$  ions formed in exothermic reactions, which did not encounter enough collisions with He and  $\text{H}_2$  for collisional de-excitation prior to the hydrogen abstraction reaction. To avoid uncertainties, the  $\text{NH}_2^+$  ions were produced in the SIS and injected into the ion trap. The primary  $\text{NH}_2^+$  ions were produced by electron bombardment of the mixture of  $\text{N}_2$  and  $\text{H}_2$  with number density ratio  $[\text{N}_2]:[\text{H}_2] \approx 10:3$ . In these experiments, the temperature of the ion trap was varied from 10 K to 300 K. An example of measured time evolutions of the normalized numbers of primary  $\text{NH}_2^+$  and produced  $\text{NH}_3^+$  ions is shown in Fig. 6. The exponential decrease in the normalized numbers of  $\text{NH}_2^+$  ions over two



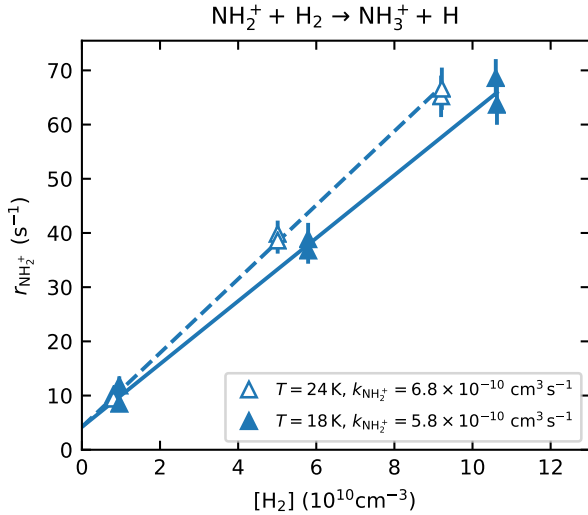
**Fig. 5.** Measured temperature dependences of the reaction rate coefficients  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  of the reactions (3a) and (3b), respectively. The data were binned and averaged for presentation. The raw data without binning are indicated with the smaller points. The stars (at 20 K) indicate the reaction rate coefficients obtained from the dependences of  $r_{\text{NH}^+}^a$  and  $r_{\text{NH}^+}^b$  on  $[\text{H}_2]$  (Fig. 4). The horizontal solid line ( $k_L$ ) indicates the value of the Langevin collisional rate coefficient. Results of Kim et al. (1975), Adams et al. (1980), and Gerlich (1993) are indicated with the open symbols. The values taken from the KIDA (Wakelam et al. 2012) and UMIST (McElroy et al. 2013) databases are also plotted.



**Fig. 6.** Measured time evolutions of the normalized numbers ( $n_x(t)/n_{\Sigma 0}$ ) of the primary  $\text{NH}_2^+$  ions and of the produced  $\text{NH}_3^+$  ions. The measurements were performed at the collisional temperature  $T = 39 \text{ K}$ , hydrogen number density  $[\text{H}_2] = 6.2 \times 10^{10} \text{ cm}^{-3}$ , and helium number density  $[\text{He}] \sim 10^{14} \text{ cm}^{-3}$ . The solid lines are fits of the measured data. The crosses ( $x$ ) indicate the sum of the normalized numbers of all ions in the ion trap,  $\Sigma(t)$ .

orders of magnitude indicates a reaction with a constant reaction rate coefficient.

Two examples of dependences of the reaction rates  $r_{\text{NH}_2^+}$  on  $[\text{H}_2]$  for reaction (4) measured at  $T = 18 \text{ K}$  and  $T = 24 \text{ K}$  are shown in Fig. 7. The linearity of the obtained dependences confirms that the time evolution of the relative numbers of  $\text{NH}_2^+$  ions in the ion trap is controlled by a binary reaction with  $\text{H}_2$ . The slope of the obtained linear dependence is given by the rate coefficient  $k_{\text{NH}_2^+}$  for the corresponding binary reaction.



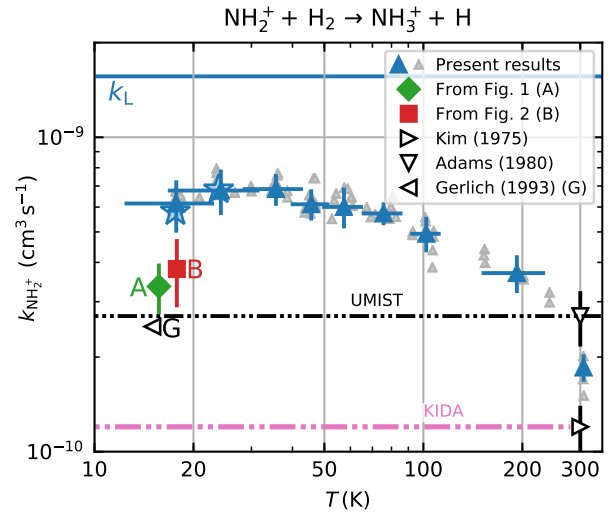
**Fig. 7.** Measured dependence of the reaction rate  $r_{\text{NH}_2^+}$  of  $\text{NH}_2^+$  ions on  $\text{H}_2$  number density. The helium number density is  $[\text{He}] \sim 1.1 \times 10^{14} \text{ cm}^{-3}$ . The collisional temperatures are  $T = 18 \text{ K}$  and  $24 \text{ K}$ . The corresponding values of the binary reaction rate coefficient  $k_{\text{NH}_2^+}$  are given by the slope of the plotted dependences. The primary  $\text{NH}_2^+$  ions were produced in the SIS and injected into the ion trap.

The measured temperature dependence of the rate coefficient  $k_{\text{NH}_2^+}$  between 15 and 300 K is shown in Fig. 8. Also plotted are the results of previous studies at 300 K (Kim et al. 1975; Adams et al. 1980), which are in agreement (within experimental accuracy) with the present value of  $k_{\text{NH}_2^+}(300 \text{ K})$ . Gerlich (1993) measured the rate coefficient  $k_{\text{NH}_2^+}$  of reaction (4) at 15 K using a 22-pole ion trap with injection of  $\text{N}^+$  ions. From time evolutions of the number of ions in the trap filled with mixture of He and  $\text{H}_2$ , he obtained the reaction rate coefficient  $k_{\text{NH}_2^+}(15 \text{ K}) = 2.5 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  (point G in Fig. 8). At similar conditions in present experiments using the injection of  $\text{N}^+$  ions into the ion trap, we obtained  $k_{\text{NH}_2^+}(16 \text{ K}) = (3.4 \pm 1.3) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  (point A, the value obtained from the data shown in Fig. 1) and  $k_{\text{NH}_2^+}(18 \text{ K}) = (3.8 \pm 1.6) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  (point B, the value obtained from the data shown in Fig. 2). The present results (point A and B) and the value from Gerlich (1993) (point G) are in rather good agreement, but we note again, that they were measured with short relaxation time, i.e. at conditions without sufficient relaxation of reacting ions (see also discussion in Gerlich 1993). The values of  $k_{\text{NH}_2^+}$  recommended by KIDA (Wakelam et al. 2012) and UMIST (McElroy et al. 2013) are also indicated in Fig. 8. KIDA and UMIST only use values measured at 300 K, and the differences in the measured temperature dependences are obvious.

In the experiments with injection of  $\text{N}^+$  or  $\text{NH}^+$  ions, we observed that at temperatures below  $\sim 100 \text{ K}$  the measured values of  $k_{\text{NH}_2^+}$  are dependent on processes of formation of  $\text{NH}_2^+$  ions. The detailed investigation of this phenomenon, including dependence on para/ortho population of  $\text{H}_2$ , will be a subject of further studies in our laboratory and are not discussed here.

### 3.3. Reaction $\text{NH}_3^+ + \text{H}_2$

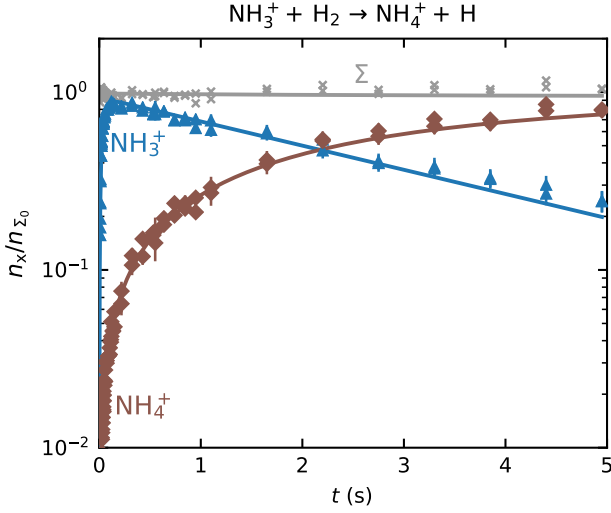
The exothermic reaction (5) of  $\text{NH}_3^+$  ion with  $\text{H}_2$  has been studied many times owing to fundamental interest and to its predicted role in the formation of  $\text{NH}_3$  in interstellar clouds (Herbst & Klemperer 1973; Le Gal et al. 2014; Gerin et al. 2016). Recent theoretical calculations of the temperature



**Fig. 8.** Temperature dependence of the rate coefficient  $k_{\text{NH}_2^+}$  of the reaction (4). The upward pointing triangles indicate the binned values measured in the present experiments with the injection of  $\text{NH}_2^+$  ions into the ion trap. The raw data without binning are indicated with the smaller points. The stars indicate the reaction rate coefficients obtained from the  $[\text{H}_2]$  dependences (see Fig. 7). The horizontal straight line ( $k_L$ ) indicates the value of the Langevin collisional rate coefficient. The values of  $k_{\text{NH}_2^+}$  measured in previous experiments at 300 K by Kim et al. (1975; ICR) and by Adams et al. (1980; SIFT) are indicated with the open triangles. The labels A and B indicate values of  $k_{\text{NH}_2^+}$  obtained in the present experiments with the injection of  $\text{N}^+$  ions into the trap at 16 and 18 K, respectively. The label G indicates the value of  $k_{\text{NH}_2^+}$  measured in similar conditions in ion trap at 15 K by Gerlich (1993).

dependence of the reaction rate coefficient  $k_{\text{NH}_3^+}$  for temperatures down to 20 K (Álvarez-Barcia et al. 2016) are in qualitative agreement with the experiments; nevertheless, at temperatures between 30 and 100 K the calculated reaction rate coefficient is smaller than the available experimental values. To provide further experimental evidence, we measured the reaction rate coefficients at temperatures ranging from 15 to 100 K. In these experiments,  $\text{N}^+$  ions were injected into the ion trap and in a sequence of hydrogen abstraction reactions,  $\text{NH}_3^+$  ions were formed there (see example in Fig. 1). Since reaction (5) is slow in the temperature range covered, the  $\text{NH}_3^+$  ions formed in the ion trap have thousands of collisions with He and  $\text{H}_2$  prior to the reaction. We can expect that in these collisions  $\text{NH}_3^+$  ions are thermalized. To see the eventual influence of excitation/de-excitation of  $\text{NH}_3^+$  ions in collisions with  $\text{H}_2$  the data were collected using a broad range of hydrogen densities (see the example of the data plotted in Fig. 2). The example of the measured time evolutions of the normalized numbers ( $n_x(t)/n_{x0}$ ) of the primary  $\text{NH}_3^+$  ions and of the produced  $\text{NH}_4^+$  ions measured at long storage time is shown in Fig. 9. The data plotted in Fig. 9 were measured at identical experimental conditions as data plotted in Fig. 1, the only difference is in the time scale. We also monitored the time evolution of the normalized numbers of  $\text{H}_3^+$ ,  $\text{N}^+$ ,  $\text{NH}^+$ , and  $\text{NH}_2^+$  ions in the trap (see the example plotted in Fig. 1), but in Fig. 9 these data are not included. As we can see from the data plotted in Fig. 1, the ions  $\text{N}^+$ ,  $\text{NH}^+$ , and  $\text{NH}_2^+$  are within 0.2 s removed from the trap in the sequence of fast reactions with  $\text{H}_2$ .

The values of  $k_{\text{NH}_3^+}$  obtained from the fits of measured time evolutions of  $n_x(t)$  at long trapping times (i.e. with long relaxation times) at temperatures from 15 up to 100 K are plotted in Fig. 10. We can see the agreement of the present results with the previous experimental results over the



**Fig. 9.** Measured time evolutions of normalized numbers ( $n_x(t)/n_{\Sigma 0}$ ) of  $\text{NH}_3^+$  and  $\text{NH}_4^+$  ions in the trap filed with  $\text{He}/\text{H}_2$  gas mixture after the injection of  $\text{N}^+$  ions from the SIS and fast formation of  $\text{NH}_3^+$  ions. The collisional temperature is  $T = 16$  K, hydrogen number density is  $[\text{H}_2] = 5.7 \times 10^{10} \text{ cm}^3 \text{ s}^{-1}$  and helium number density is  $[\text{He}] = 5.6 \times 10^{13} \text{ cm}^3 \text{ s}^{-1}$ . The fitted solution of the corresponding system of rate equations is indicated with the solid lines. The crosses ( $\times$ ) indicate the sum of the normalized numbers of all ions in the ion trap ( $\Sigma(t)$ ).

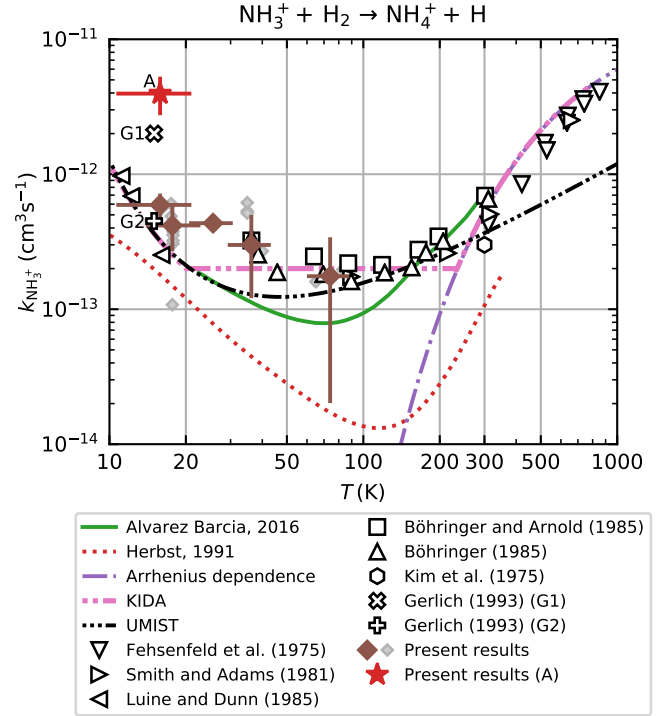
covered range of temperatures; in other words, our experiment confirms that the recently calculated reaction rate coefficient by [Álvarez-Barcia et al. \(2016\)](#) is systematically lower than the experimental data at low temperatures.

When the trapping time was short ( $\lesssim 100$  ms) and the time for the relaxation of produced  $\text{NH}_3^+$  ions was not long enough, we observed a substantial increase in the reaction rate coefficient  $k_{\text{NH}_3^+}$  in comparison with values obtained in experiments using long trapping time. An example of a value of  $k_{\text{NH}_3^+}$  obtained from the fit of time evolutions of  $n_x(t)$  at a short trapping time (i.e. with a short relaxation time) is plotted in Fig. 10, labelled A (non-thermalized  $k_{\text{NH}_3^+}^A(17 \text{ K}) = 4.0 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$ ). A similar increase in the reaction rate coefficient at short trapping time was observed by Gerlich with his ion trap experiment at 15 K ([Gerlich 1993](#)). His results obtained at short and long trapping times are labelled G1 and G2 in Fig. 10. At He pressures used in the ion trap in the present experiments (with  $\text{NH}_3^+$ ) we did not observe any effect of a ternary reaction. The temperature dependences of  $k_{\text{NH}_3^+}$  recommended by KIDA ([Wakelam et al. 2012](#)) and by UMIST ([McElroy et al. 2013](#)) are also plotted in Fig. 10.

#### 4. Summary

Following our previous studies of the reaction of  $\text{N}^+$  ions with molecular hydrogen, we studied the chain of hydrogen abstraction reactions of  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  ions with molecular hydrogen leading to the formation of  $\text{NH}_4^+$  ions at temperatures relevant for interstellar clouds. Using the 22-pole ion trap apparatus, we measured the temperature dependences of the rate coefficients of the reactions of  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  ions with  $\text{H}_2$  at temperatures down to 15 K.

In the reactions of  $\text{NH}^+$  ions, we observed two products:  $\text{NH}_2^+$  ( $\sim 97\%$ ) and  $\text{H}_3^+$  ( $\sim 3\%$ ) (see Figs. 3–5). There have been several previous studies of the reaction of  $\text{NH}^+$  with  $\text{H}_2$ , but this is the first study giving the temperature dependences of the reaction rate coefficients  $k_{\text{NH}^+}^a$  and  $k_{\text{NH}^+}^b$  at astrophysically



**Fig. 10.** The temperature dependence of the reaction rate coefficient  $k_{\text{NH}_3^+}$ . The binned present data measured with long trapping times are indicated with the full diamonds. The raw data without binning are indicated with the smaller points. Present data measured with short trapping time are indicated with a red star (label A). The points labelled as G1 and G2 are the values measured by [Gerlich \(1993\)](#) using short and long trapping time, respectively. Previous experimentally determined binary rate coefficient are taken from [Fehsenfeld et al. \(1975\)](#), [Kim et al. \(1975\)](#), [Smith & Adams \(1981\)](#), [Adams & Smith \(1984\)](#), [Luine & Dunn \(1985\)](#), [Böhlinger & Arnold \(1985\)](#), [Böhlinger \(1985\)](#), [Barlow & Dunn \(1987\)](#). The plotted theoretical temperature dependences were calculated by [Herbst et al. \(1991\)](#) (dashed line) and by [Álvarez-Barcia et al. \(2016\)](#) (solid line). The dash dotted line indicates the Arrhenius dependence obtained by the fit of the data measured at temperatures above 300 K ([Fehsenfeld et al. 1975](#)). Included are also the temperature dependences recommended by KIDA ([Wakelam et al. 2012](#)) and by UMIST ([McElroy et al. 2013](#)).

relevant temperatures. In the covered temperature range, from 15 to 130 K, both reaction rate coefficients are approximately constant (see Fig. 5). The value of  $k_{\text{NH}^+}^b$  obtained in our experiment at 130 K,  $k_{\text{NH}^+}^b = (6.0 \pm 2.4) \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$  is approximately 4 times lower than the only previously available value, which was obtained at 300 K ([Adams et al. 1980](#),  $k_{\text{NH}^+}^b = (23 \pm 5) \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$ ). Although these values are not directly comparable due to the difference in temperatures, it is unlikely that the reaction rate coefficient, which is constant below 130 K, would change by a factor of 4 between 130 K and 300 K. There are several systematic effects that can specifically affect the measurement of such a minor reaction channel. We are not able to discuss the possible sources of error in the experiment of [Adams et al. \(1980\)](#), such as different detection efficiencies and diffusion rates for  $\text{H}_3^+$  and  $\text{NH}_2^+$ . In our study, the detection efficiency was calibrated in experiments with  $\text{H}_3^+$  and  $\text{CH}_4$  (see Sect. 3.1) at otherwise identical conditions to the  $\text{NH}^+$  experiments. The standard deviation of the detection efficiency was smaller than 10% at given experimental conditions, although it can vary by tens of percentage points with change of potentials in the extraction and detection system. The other systematic effect is related to



the kinetic energy release of the reaction. Although we do not know what fraction of the 0.8 eV reaction exothermicity is converted to kinetic energy of the products, it is possible that the  $\text{H}_3^+$  ion gains sufficient energy to escape from the trap. However, we did not observe any variation in the branching ratio with trap RF amplitude or He pressure, which would indicate this effect.

The temperature dependence of the rate coefficient of the reaction of  $\text{NH}_2^+$  ion with  $\text{H}_2$  was measured for temperatures from 15 to 300 K (see Fig. 8). The only observed product of the reaction was the  $\text{NH}_3^+$  ion. The rate coefficient of this reaction drops monotonically from  $k_{\text{NH}_2^+}(17\text{ K}) = 6.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  to  $k_{\text{NH}_2^+}(300\text{ K}) = 2.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ . At 300 K, our value of  $k_{\text{NH}_2^+}$  is in agreement with the results of both previous studies (Kim et al. 1975; Adams et al. 1980). At 15 K, our value is significantly higher than that of Gerlich (1993). Our tests indicate that this might be due to insufficient relaxation of the  $\text{NH}_2^+$  ions in the experimental procedure of Gerlich (1993). Our experiments were carried out with an injection of  $\text{NH}_2^+$  ions from the SIS to the ion trap, and the decay of the number of primary ions was observed over long trapping times to exclude the influence of the process of formation and relaxation of  $\text{NH}_2^+$ . The time dependence of the  $\text{NH}_2^+$  ion formation process and its relaxation in He or  $\text{H}_2$  collisions is discussed. The detailed investigation of this phenomenon, including dependence on para/ortho population of  $\text{H}_2$ , will be a subject of further studies in our laboratory and is not discussed here. This is the first study giving the temperature dependence of the reaction rate coefficient  $k_{\text{NH}_2^+}$  from 15 K to 300 K.

Due to the astrophysical significance, the temperature dependence of the rate coefficient of the reaction of  $\text{NH}_3^+$  with  $\text{H}_2$ , producing  $\text{NH}_4^+$ , was also studied. The measured temperature dependence of the reaction rate coefficient has a minimum at temperatures around 70 K ( $k_{\text{NH}_3^+}(70\text{ K}) = 1.0 \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$ ). At lower temperatures the value of  $k_{\text{NH}_3^+}$  slowly increases with decreasing temperature. The presented data for the reaction of  $\text{NH}_3^+$  ions with  $\text{H}_2$  are in agreement with previous experimental values. The recently calculated reaction rate coefficient of Álvarez-Barcia et al. (2016) is systematically lower than all the experimental data at temperatures below 150 K, although the discrepancy is close to the experimental error.

Studies of the reactions of  $\text{NH}^+$ ,  $\text{NH}_2^+$ , and  $\text{NH}_3^+$  ions with  $\text{D}_2$ , HD, and para- $\text{H}_2$  are in preparation.

*Acknowledgements.* We thank the Technische Universität Chemnitz and the Deutsche Forschungsgemeinschaft for lending the 22-pole trap instrument to the Charles University team. We thank Prof. Dieter Gerlich for the discussion and helpful suggestions. This work was partly supported by the Czech Science Foundation (GACR 17-19459S, GACR 17-18067S), and by Charles University (project Nr. GAUK 1584217, 1144616, 1168216).

## References

Acharyya, K., & Herbst, E. 2015, *ApJ*, 812, 142  
Adams, N., & Smith, D. 1984, *IJMSI*, 61, 133

- Adams, N., & Smith, D. 1985, *CPL*, 117, 67  
Adams, N. G., Smith, D., & Paulson, J. F. 1980, *JChPh*, 72, 288  
Álvarez-Barcia, S., Russ, M.-S., Meisner, J., & Kästner, J. 2016, *FaDi*, 195, 69  
Barlow, S. E., & Dunn, G. H. 1987, *IJMSI*, 80, 227  
Bohme, D. K., Mackay, G. I., & Schiff, H. I. 1980, *JChPh*, 73, 4976  
Böhlinger, H. 1985, *CPL*, 122, 185  
Böhlinger, H., & Arnold, F. 1985, *Molecular Astrophysics, NATO ASI Series* (Dordrecht: Springer), 639  
Caselli, P., Bizzocchi, L., Keto, E., et al. 2017, *A&A*, 603, L1  
Cernicharo, J., Tercero, B., Fuente, A., et al. 2013, *ApJ*, 771, L10  
Cheung, A. C., Rank, D. M., Townes, C. H., Thornton, D. D., & Welch, W. J. 1968, *PhRvL*, 21, 1701  
Endres, E., Egger, G., Lee, S., et al. 2017, *JMoSp*, 332, 134  
Fehsenfeld, F. C., Schmeltekopf, A. L., & Ferguson, E. E. 1967, *JChPh*, 46, 2802  
Fehsenfeld, F. C., Lindinger, W., Schmeltekopf, A. L., Albritton, D. L., & Ferguson, E. E. 1975, *JChPh*, 62, 2001  
Florescu-Mitchell, A. I., & Mitchell, J. B. A. 2006, *PhR*, 430, 277  
Gerin, M., Neufeld, D. A., & Goicoechea, J. R. 2016, *ARA&A*, 54, 181  
Gerlich, D. 1992, *AdChP*, 82, 1  
Gerlich, D. 1993, *FaTr*, 89, 2199  
Gerlich, D. 1995, *PhyS*, 1995, 256  
Gerlich, D. 2008, *Low Temperatures and Cold Molecules* (Imperial College Press), 121  
Gerlich, D., Borodi, G., Luca, A., Mogo, C., & Smith, M. A. 2011, *ZPC*, 225, 5  
Harju, J., Daniel, F., Sipilä, O., et al. 2017, *A&A*, 600, A61  
Hauser, D., Lee, S., Carelli, F., et al. 2015, *NatPh*, 11, 467  
Hejduk, M., Dohnal, P., Varju, J., et al. 2012, *PSST*, 21, 024002  
Hejduk, M., Dohnal, P., Rubovič, P., et al. 2015, *JChPh*, 143, 044303  
Herbst, E., & Klemperer, W. 1973, *ApJ*, 185, 505  
Herbst, E., Defrees, D. J., & McLean, A. D. 1987, *ApJ*, 321, 898  
Herbst, E., DeFrees, D. J., Talbi, D., et al. 1991, *JChPh*, 94, 7842  
Hily-Blant, P., Pineau des Forêts, G., Faure, A., Le Gal, R., & Padovani, M. 2013, *A&A*, 557, A65  
Hunter, E. P. L., & Lias, S. G. 1998, *JPCRD*, 27, 413  
Kim, J. K., Theard, L. P., & Huntress, W. T. 1975, *JChPh*, 62, 45  
Knauth, D. C., Andersson, B. G., McCandliss, S. R., & Warren Moos, H. 2004, *Nature*, 429, 636  
Kovalenko, A., Tran, T., Rednyk, S., et al. 2018, *ApJ*, 856, 100  
Le Gal, R., Hily-Blant, P., Faure, A., et al. 2014, *A&A*, 562, A83  
Luine, J. A., & Dunn, G. H. 1985, *ApJ*, 299, L67  
Marquette, J. B., Rebrion, C., & Rowe, B. R. 1988, *JChPh*, 89, 2041  
Marquette, J. B., Rebrion, C., & Rowe, B. R. 1989, *A&A*, 213, L29  
McElroy, D., Walsh, C., Markwick, A. J., et al. 2013, *A&A*, 550, A36  
Meyer, D. M., & Roth, K. C. 1991, *ApJ*, 376, L49  
Ng, C., Baer, T., & Powis, I. 1994, *Unimolecular and Bimolecular Ion-Molecule Reaction Dynamics, Wiley Series In Ion Chemistry and Physics* (Wiley)  
Nieva, M.-F., & Przybilla, N. 2012, *A&A*, 539, A143  
Novotný, O., Berg, M., Bing, D., et al. 2014, *ApJ*, 792, 132  
Persson, C. M., Black, J. H., Cernicharo, J., et al. 2010, *A&A*, 521, L45  
Persson, C., Luca, M., Mookerjee, B., et al. 2012, *A&A*, 543, A145  
Plašil, R., Mehner, T., Dohnal, P., et al. 2011, *ApJ*, 737, 60  
Plašil, R., Zymak, I., Jusko, P., et al. 2012, *RSPTA*, 370, 5066  
Plašil, R., Zymak, I., Hejduk, M., et al. 2014, *J. Phys. Conf. Ser.*, 488, 122003  
Rist, C., Faure, A., Hily-Blant, P., & Le Gal, R. 2013, *JPCA*, 117, 9800  
Roučka, S., Rednyk, S., Kovalenko, A., et al. 2018, *A&A*, 615, L6  
Scott, G. B. I., Freeman, C. G., & McEwan, M. J. 1997, *MNRAS*, 290, 636  
Smith, D., & Adams, N. G. 1981, *MNRAS*, 197, 377  
van Dishoeck, E. F., Jansen, D. J., Schilke, P., & Phillips, T. G. 1993, *ApJ*, 416, L83  
Wakelam, V., Herbst, E., Loison, J.-C., et al. 2012, *ApJS*, 199, 21  
Wyrowski, F., Güsten, R., Menten, K. M., et al. 2016, *A&A*, 585, A149  
Zymak, I., Hejduk, M., Mulin, D., et al. 2013, *ApJ*, 768, 86

Article II: OH<sup>+</sup> Formation in the Low-temperature O<sup>+</sup>(<sup>4</sup>S) + H<sub>2</sub>  
Reaction

Kovalenko, A., Tran, T. D., Rednyk, S., Roučka, Š., Dohnal, P., Plašil, R.,  
Gerlich, J., Glosík, J.

2018

*The Astrophysical Journal*, 856(2), 100(1-6)



# OH<sup>+</sup> Formation in the Low-temperature O<sup>+</sup>(<sup>4</sup>S) + H<sub>2</sub> Reaction

Artem Kovalenko , Thuy Dung Tran , Serhiy Rednyk , Štěpán Roučka , Petr Dohnal , Radek Plašil ,  
Dieter Gerlich , and Juraj Glosík

Department of Surface and Plasma Science, Faculty of Mathematics and Physics, Charles University, Prague, Czech Republic; [radek.plasil@mff.cuni.cz](mailto:radek.plasil@mff.cuni.cz)

Received 2017 December 18; revised 2018 February 7; accepted 2018 February 18; published 2018 March 28

## Abstract

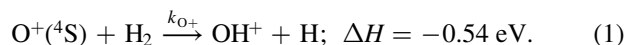
Formation of OH<sup>+</sup> in collisions of ground-state O<sup>+</sup>(<sup>4</sup>S) ions with normal H<sub>2</sub> has been studied using a variable temperature 22-pole RF ion trap. From 300 to 30 K the measured reaction rate coefficient is temperature-independent, with a small decrease toward 15 K. The recent wave packet calculation predicts a slightly steeper temperature dependence. The rate coefficients at 300 and 15 K are almost the same,  $(1.4 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  and  $(1.3 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ , respectively. The influence of traces of the two metastable ions, O<sup>+</sup>(<sup>2</sup>D) and O<sup>+</sup>(<sup>2</sup>P), has been examined by monitoring the H<sup>+</sup> products of their reactions with H<sub>2</sub>, as well as by chemically probing them with N<sub>2</sub> reactant gas.

*Key words:* astrochemistry – ISM: molecules – molecular data – molecular processes

## 1. Introduction

Oxygen, the third most abundant element in the universe, plays an important role in the chemistry of many environments. In the interstellar medium, the formation of OH<sup>+</sup> ions in reactions of O<sup>+</sup> with H<sub>2</sub> is a key process in the synthesis of water molecules (Gerin et al. 2016). Subsequent reactions of OH<sup>+</sup> with H<sub>2</sub> lead to H<sub>2</sub>O<sup>+</sup> and H<sub>3</sub>O<sup>+</sup> formation, which, following dissociative recombination with electrons, produce H<sub>2</sub>O (Jensen et al. 2000; Neau et al. 2000). The vital importance of H<sub>2</sub>O in interstellar space has recently been discussed by Hollenbach et al. (2012). New ground- and space-based observations (e.g., Pilbratt et al. 2010; Wyrowski et al. 2010; Hollenbach et al. 2012; Gómez-Carrasco et al. 2014; Indriolo et al. 2015; Muller et al. 2016; Neufeld & Wolfire 2016, 2017) that focused on the role of OH<sub>n</sub><sup>+</sup> in the interstellar medium indicate that in H<sub>2</sub>-dominated regions, there are two main routes to form OH<sup>+</sup> (Hollenbach et al. 2009, 2012; Gerin et al. 2016). In the first one, cosmic rays (CR) ionize H<sub>2</sub>. The H<sub>2</sub><sup>+</sup> reacts with H<sub>2</sub> to H<sub>3</sub><sup>+</sup>, which transfers a proton to O (de Ruette et al. 2016). In the second one, CR first forms protons. Electron transfer from O to H<sup>+</sup> leads to O<sup>+</sup> and OH<sup>+</sup> is formed by the title reaction. The importance of these reactions and many other competing processes depends on the conditions prevailing in the different regions of space. To understand the astronomical observations, more theoretical and experimental studies of the formation and destruction of OH<sup>+</sup> ions are needed (Indriolo et al. 2015), especially at the low temperatures of dense interstellar clouds.

In the present study, we focus predominantly on the temperature dependence of the exoergic reaction of ground-state O<sup>+</sup>(<sup>4</sup>S) ion with H<sub>2</sub> molecule:



The rate coefficient is denoted  $k_{\text{O}^+}$ . The reaction enthalpy  $\Delta H$  at 0 K was calculated from tabulated enthalpies of formation, ionization potentials, and dissociation energies (Wiedmann et al. 1992; Chase 1998; Sansonetti & Martin 2005; Liu et al. 2009). The estimated errors do not exceed 1 on the last digit of all enthalpies given in this paper.

As thoroughly reviewed in previous papers (Bulut et al. 2015, for instance) and summarized below in Section 4, the reaction (1) has been studied often, both experimentally and theoretically. Nonetheless, there are no reports of measurements below 300 K in the literature so far, although an understanding of the title reaction at low temperatures is required for the astronomical modeling. Therefore, we decided to extend the temperature range of measurements of the title reaction from room temperature down to 15 K.

In the following, we first briefly describe the present experiments. In Section 3 the new experimental results for the title reaction are presented. The new data are compared with those from previous experiments and with recent theoretical predictions.

## 2. Experimental

### 2.1. 22PT Instrument

The cryogenic 22-pole RF ion trap instrument used in the present study has been described previously (see, e.g., Gerlich et al. 2011, 2012, 2013; Zymak et al. 2011, 2013; Plašil et al. 2012; Mulin et al. 2015). Therefore, only a brief description is given here. Atomic oxygen ions, O<sup>+</sup>, are produced in the storage ion source by bombarding N<sub>2</sub>O with electrons of energies  $E_e$ . For most experiments,  $E_e = 50 \text{ eV}$  has been used. To vary the fraction of metastable O<sup>+</sup> ions (O<sup>+</sup>(<sup>2</sup>D) and O<sup>+</sup>(<sup>2</sup>P)), the electron energy has been varied between 21 and 145 eV. The formed ions are extracted from the ion source, selected according to their mass-to-charge ratio, and injected into the ion trap. Typically, one filling of the trap consists of a few hundred primary reactant ions. The kinetic energy of the ions is cooled in collisions with helium atoms of the buffer gas, while a small admixture of H<sub>2</sub> leads to reaction ( $[\text{He}]/[\text{H}_2] \approx 100$ ). After a preselected trapping time ( $t$ , here up to 500 ms), all ions are extracted, their mass is analyzed with a quadrupole mass spectrometer, and they are counted using a micro channel plate detector.

### 2.2. Gas Number Density

The number density of buffer and reactant gases inside the trap is determined using a spinning rotor gauge connected

directly to the trap. In addition, an ion gauge mounted onto the vacuum chamber containing the trap is used to monitor the gas flow. It is calibrated from time to time using the spinning rotor gauge. The uncertainty of the reactant number density is estimated conservatively to be 20%.

### 2.3. Ortho- and Para- $H_2$

In the present study we used normal hydrogen, composed of  $\frac{1}{4}$  para- $H_2$  and  $\frac{3}{4}$  ortho- $H_2$ . It has been checked (Zymak et al. 2013) that there is no ortho-para conversion in the gas inlet system or in the trap at low temperatures. Previous experimental studies of the title reaction  $O^+ + H_2$  (by FA, SIFT, SIFDT, GIB, and ICR experiments, the details and the references are given below) were carried out with hydrogen at 300 K, i.e., with normal hydrogen.

### 2.4. Collision Temperature

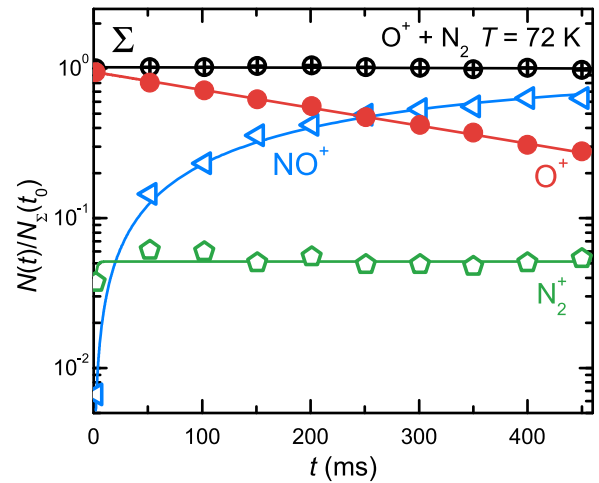
In most radiofrequency (RF) ion traps, the translational temperature of the thermalized ions is slightly higher than that of the trap ( $T_{22PT}$ ), as discussed in the literature (Gerlich & Horning 1992; Gerlich 2008; Asvany & Schlemmer 2009; Chakrabarty et al. 2013; Endres et al. 2017). Due to the favorable ionic and molecular mass ratio 16:2 in reaction (1), the collisional temperature, denoted here simply by  $T$ , is mainly determined by the cold  $H_2$ . Note that  $T$  is the mass-weighted mean value of the temperature of the two reactants. We estimate that in our trap  $T < T_{22PT} + 10$  K and for simplicity we define it as  $T = T_{22PT} + (5 \pm 5)$  K. For more details and discussions see Plašil et al. (2012), Zymak et al. (2013), Mulin et al. (2015), and Tran et al. (2018).

### 2.5. Internal Excitation of $O^+$ Ions

Depending on the electron energy and the number density, the  $O^+$  ions from the storage ion source are either in the ground-state  $^4S$  or in one of two excited metastable states ( $^2D$  or  $^2P$ ). Quenching the excited  $O^+$  in the trap via collisions with He atoms is inefficient. See, e.g., confirmation in a helium-buffered SIFT experiment (Glosík et al. 1978). The radiative life-times of these metastables are larger than  $5.6 \times 10^3$  s and 4.9 s, respectively (Zeippen 1987; Godefroid & Fischer 1984), which are long in comparison with the length of typical storage time used in the present experiments ( $< 0.5$  s).

To determine the population of trapped  $O^+$  ions in the  $^2D$  and  $^2P$  states, we studied their interaction with  $N_2$  in the trap. Note that for the  $O^+(^4S)$  ground-state ions the charge transfer reaction is endoergic by 1.96 eV. In analyses of the experimental data we have used the results of previous experimental studies in which a very slow reaction of  $O^+(^4S)$  with  $N_2$  producing  $NO^+$  and a very efficient charge transfer reaction of excited  $O^+(^2D, ^2P)$  ions with  $N_2$  producing  $N_2^+$  were observed (Glosík et al. 1978; Smith et al. 1978; Johnsen & Biondi 1980; Hierl et al. 1997; Le Garrec et al. 2003). Thus, by monitoring the production of  $NO^+$  and  $N_2^+$ , it is possible to determine the relative populations of the ground-state  $O^+(^4S)$  ions and of the excited  $O^+(^2D, ^2P)$  ions in the ion trap.

The measured data are in good agreement with results of previous studies (Glosík et al. 1978; Hierl et al. 1997; Le Garrec et al. 2003). An example of data obtained at collision temperature  $T = 72$  K,  $[N_2] = 3.9 \times 10^{11}$  cm $^{-3}$ ,  $[He] = 5.7 \times 10^{13}$  cm $^{-3}$ , and electron energy  $E_e = 75$  eV is shown in Figure 1. We observed that only 5% of the injected



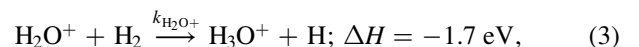
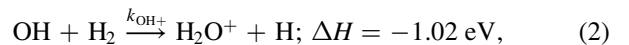
**Figure 1.** Time evolutions of the numbers of trapped  $O^+$ ,  $N_2^+$ , and  $NO^+$  ions. The time dependence of the total number of ions ( $\Sigma$ ) is also indicated. The data were measured at a temperature  $T = 72$  K, the number densities in the trap  $[N_2] = 3.9 \times 10^{11}$  cm $^{-3}$ ,  $[He] = 5.7 \times 10^{13}$  cm $^{-3}$ , and the energy of ionizing electrons  $E_e = 75$  eV.

$O^+$  were converted into  $N_2^+$  ions. The rate coefficient obtained for the reaction  $O^+(^4S) + N_2$  is equal to  $(6.6 \pm 0.1) \times 10^{-12}$  cm $^3$  s $^{-1}$ . This value agrees within a few percent with the CRESU results measured at the same temperature (Le Garrec et al. 2003). On the base of these studies we can safely conclude that over 90% of the trapped  $O^+$  ions are in the  $^4S$  ground state when using the energy of ionizing electrons  $E_e < 75$  eV. In the experiments where the rate coefficients  $k_{O^+}$  of reaction (1) were measured as a function of collision temperature, the energy of ionizing electrons  $E_e = 50$  eV was used.

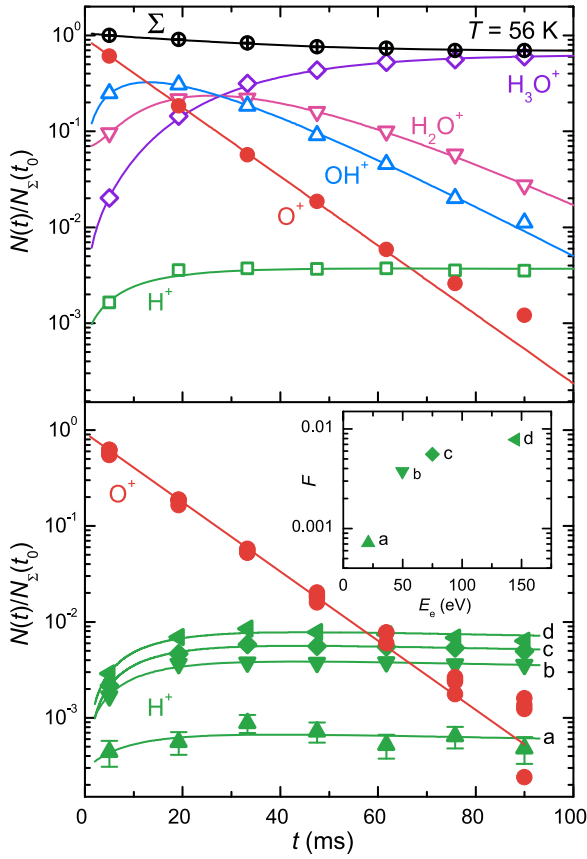
## 3. Results

A typical result for trapping of  $O^+$  ions in a He/ $H_2$  mixture is shown in the upper panel of Figure 2. The number of different ions counted after storage times  $t$  is denoted as  $N_X(t)$ , where the index X refers to the various ions in the trap, e.g.,  $N_{O^+}(t)$  is the number of  $O^+$  ions. The total number of detected ions at time  $t$  is  $N_\Sigma(t) = \sum N_X(t)$ . For a better comparison of the results from different experiments, we divide the ion counts by the total number of ions measured at the shortest trapping time  $t_0$ , i.e., we plot normalized values  $N_X(t)/N_\Sigma(t_0)$ . The plotted numbers of  $H^+$  are corrected for the difference in detection efficiency relative to  $O^+$ , which was obtained using a separate calibration reaction ( $H^+ + CH_4$  producing  $CH_3^+$  and  $CH_4^+$ ). The decrease of  $\Sigma = N_\Sigma(t)/N_\Sigma(t_0)$  (crossed open circles in Figure 2) with time indicates smaller detection efficiencies for  $OH^+$ ,  $H_2O^+$ , and  $H_3O^+$  relative to  $O^+$ . The detection efficiencies of these heavier ions are not accounted for in the plots; however, they are treated as free parameters in the fits.

Inspection of the upper panel of Figure 2 reveals that  $O^+$  ions react to produce  $OH^+$ , followed by a sequence of additional exothermic reactions with  $H_2$ :



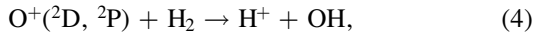
with the rate coefficients denoted as  $k_{OH+}$  and  $k_{H_2O+}$ . The final product ion is protonated water. The given 0 K enthalpy



**Figure 2.** Time dependence of the normalized numbers of the indicated ions after injecting  $O^+$  ions into the trap. The measurements were performed at  $T = 56$  K,  $[H_2] = 5.4 \times 10^{10} \text{ cm}^{-3}$ , and  $[He] = 5.5 \times 10^{12} \text{ cm}^{-3}$ . All lines indicate the results from the fits (see the explanation in the text). Upper panel: electron energy  $E_e = 50$  eV. Note that almost all injected  $O^+$  ions are finally converted into  $H_3O^+$  ions via sequential reactions with  $H_2$ . The small amount of  $H^+$  ions produced (growing to  $\approx 0.5\%$ ) is indicative of the relative population of metastable  $O^+(^2D, ^2P)$  ions in the trap. Lower panel: electron energies  $E_e = 21, 50, 75,$  and  $145$  eV. The inset in the lower panel reveals that the fraction of  $H^+$  products,  $F$ , increases with  $E_e$ .

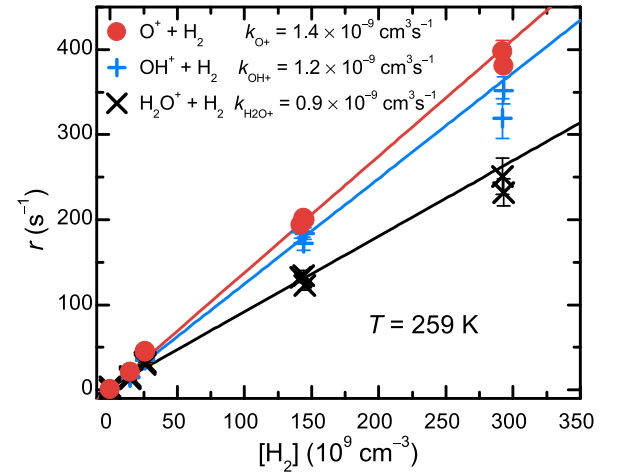
changes given in eV were calculated from the relevant heats of formation (Haney & Franklin 1969; Chase 1998), ionization potentials (Wiedmann et al. 1992; Lauzin et al. 2015), and dissociation energies (Liu et al. 2009).

The observed  $H^+$  ions are not due to  $O^+(^4S) + H_2$ , since this reaction is not only endoergic by 0.06 eV but also spin-forbidden (Flesch & Ng 1991; Li et al. 1997). These protons are produced in reactions of electronically excited metastable  $O^+(^2D)$  and  $O^+(^2P)$  ions with  $H_2$



where  $O^+(^2D, ^2P)$  indicates an unspecified mixture of  $O^+(^2D)$  and  $O^+(^2P)$  ions, with respective reaction enthalpies  $\Delta H(^2D) = -3.27$  eV,  $\Delta H(^2P) = -4.96$  eV.

The low fraction of  $H^+$  products in comparison with the population of  $O^+(^2D, ^2P)$  obtained in experiments with  $N_2$  indicates that  $OH^+$  is also produced in reactions of  $O^+(^2D, ^2P)$ . Based on reactions (1)–(4), the time variations of the primary ions and all product ions were fitted using a kinetic model. Free parameters are the rate coefficients, the initial numbers of all ions, as well as the detection efficiencies, as mentioned above. The overall agreement of the fits (solid lines) and the data



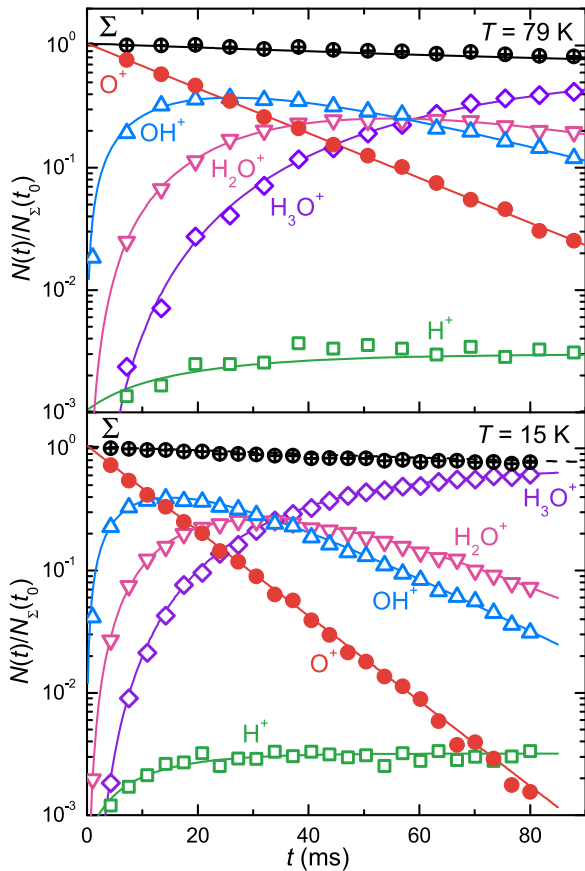
**Figure 3.** Loss rates  $r_{O^+}$ ,  $r_{OH^+}$ , and  $r_{H_2O^+}$  of  $O^+$ ,  $OH^+$ , and  $H_2O^+$  ions as a function of  $H_2$  number density. The collisional temperature  $T = 259$  K. The corresponding binary reaction rate coefficients  $k_{O^+}$ ,  $k_{OH^+}$ , and  $k_{H_2O^+}$  are given by the slope of the plotted dependencies (see the text).

points is very good. The number of  $O^+$  is well-described by a mono-exponential decay over almost 3 orders of magnitude.

The lower panel of Figure 2 illustrates the increase of the fraction of metastable  $O^+$  ions with energy of the ionizing electrons  $E_e$  increasing from 21 to 145 eV. The fraction of  $H^+$  ions produced via reaction (4), which is proportional to the population of  $O^+(^2D, ^2P)$  ions in the trap, reaches up to 1% (see inset). This is consistent with the results of probing with  $N_2$ , which revealed that over 90% of trapped  $O^+$  ions are in ground-state  $O^+(^4S)$  for  $E_e < 75$  eV (see Section 2.5).

The reaction rate coefficient obtained for  $OH^+$  production,  $k_{OH^+}(56 \pm 5 \text{ K}) = (1.5 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ , is within the uncertainty of our data constant for all electron energies, which confirms that the measured values of  $k_{O^+}$  are not influenced by the presence of metastable  $O^+(^2D, ^2P)$  ions. This is because the fraction of  $O^+(^2D, ^2P)$  is low (below 0.1) and the rate coefficient for the reaction of  $O^+(^4S)$  is nearly collisional, i.e., the contribution from  $O^+(^2D, ^2P)$  ions to the decay rate of  $N_{O^+}(t)$  can be on the order of few percent. To minimize the influence of  $O^+(^2D, ^2P)$  ions, the electron energy was kept at  $E_e = 50$  eV in all measurements of  $k_{O^+}$  and the production of  $H^+$  ions was monitored.

Another experimental test carried out was the dependence of the reaction rate  $r_{O^+}$  on the reactant gas number density  $[H_2]$  measured at  $T = 259$  K, which is shown in Figure 3. The linearity of the dependencies obtained confirms that the decrease of the number of primaries in the trap is due to a binary ion-molecule reaction with  $H_2$ , i.e., the rate can be expressed by the formula,  $r_{O^+} = k_{O^+}[H_2] + r_{bg}$ . The level of the background loss rate ( $r_{bg}$ ) measured in the experiment with pure He without  $H_2$  is  $r_{bg} = 0.9 \text{ s}^{-1}$ . At lower temperatures, cryopumping is more efficient in reducing the residual background gas (mostly  $N_2O$ ,  $H_2O$  from the ion source). Figure 3 also includes reaction rates  $r_{OH^+}$  and  $r_{H_2O^+}$  obtained from the fits of the same set of the data. The reaction rate coefficients obtained from the linear fits of the data in Figure 3 are:  $k_{O^+}(259 \text{ K}) = (1.4 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ ,  $k_{OH^+}(259 \text{ K}) = (1.2 \pm 0.2) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ , and  $k_{H_2O^+}(259 \text{ K}) = (0.90 \pm 0.18) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ . These values are in good agreement with results from previous experiments at 300 K (see, e.g., ref. SIFT (Jones et al. 1981), FA (Fehsenfeld



**Figure 4.** Time dependence of the normalized number of indicated ions in the trap after injection of  $O^+$  ions. The open circles indicate the normalized total number of ions ( $\Sigma$ ). Upper panel:  $T = 79$  K,  $[H_2] = 2.7 \times 10^{10} \text{ cm}^{-3}$ ,  $[He] = 6.7 \times 10^{12} \text{ cm}^{-3}$ . Lower panel:  $T = 15$  K,  $[H_2] = 6.0 \times 10^{10} \text{ cm}^{-3}$ ,  $[He] = 1.5 \times 10^{14} \text{ cm}^{-3}$ . The experimental data are fitted by a kinetic model (solid lines) resulting in the rate coefficients for reactions (1)–(3).

et al. 1967), ICR (Kim et al. 1975), and VT-SIFT (Martinez et al. 2015)).

No dependence of the measured reaction rate coefficients on the helium number density was observed in the range of  $5 \times 10^{12} - 5 \times 10^{13} \text{ cm}^{-3}$ .

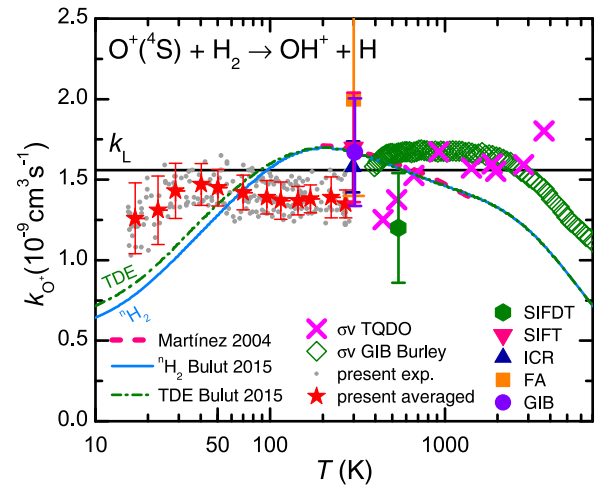
The change of the ion composition with increasing storage time at collisional temperatures  $T = 15$  and  $79$  K is shown in Figure 4 together with the results from fits. Similar to Figure 2, sequential addition of hydrogen atoms to  $O^+$  leads finally to  $H_3O^+$  and less than one percent of  $H^+$  ions are formed. The decay rate of the number of  $O^+$  ions has been studied at several collisional temperatures, ranging from  $15$  K to  $300$  K.

The temperature dependence obtained for the rate coefficient  $k_{O^+}$  of the title reaction is plotted in Figure 5. The spread of the data is indicative of the reproducibility of our measurements. Plotted is also a collection of results from previous experiments and theories. All these results will be discussed below.

## 4. Discussion

### 4.1. The Role of Metastable $O^+$

As is illustrated in the lower part of Figure 2, the variation of electron energy, used in the ion source for production of  $O^+$  from  $N_2O$ , results in different populations of the  $^4S$  ground state and of the metastable excited states,  $^2D$  and  $^2P$ . By monitoring the production of  $N_2^+$  ions in the reaction of trapped  $O^+$  ions



**Figure 5.** Dependence of the rate coefficient  $k_{O^+}$  for the reaction of the ground-state  $O^+(^4S)$  ions with normal  $H_2$  on the collision temperature. The dots are the present data, as measured and explained in the text, while the filled stars are weighted averages over logarithmically spaced bins in temperature. The error bars indicate the sum of statistical uncertainties and the influence of temperature uncertainty. The horizontal straight line ( $k_L$ ) indicates the value of the Langevin collisional rate coefficient. The filled symbols at  $300$  and  $500$  K are the results measured with the methods indicated (see the text). The points extending above  $1000$  K are effective rate coefficients measured by the guided ion beam technique (GIB; Burley et al. 1987) and TQDO (Li et al. 1997). The solid and dashed-dotted lines are reaction rate coefficients for normal  $H_2$  and for  $H_2$  with a thermal population of nuclear spin states, respectively, based on theoretical cross sections of Bulut et al. (2015). The theoretical rate coefficients of Martínez et al. (2004) are indicated by the dashed line.

with  $N_2$  we have found operating conditions where the population of the excited metastable ions is  $\approx 5\%$ . Under these conditions ( $E_e < 50$  eV), the influence of metastable  $O^+$  ions on the determination of  $k_{O^+}$  is negligible (see the lower panel of Figure 2 and related text). This is because  $5\%$  of  $O^+(^2D, ^2P)$  can add to a measured value of the rate coefficient by at most  $5\%$  of the value of collisional rate coefficient ( $\approx 0.08 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ ).

### 4.2. Temperature Dependence

At first sight, Figure 5 shows a typical temperature dependence expected for an exothermic ion-molecule reaction: the measured rate coefficients,  $k_{O^+}$ , are close to the Langevin rate coefficient  $k_L$  (horizontal straight line). The two sets of data above room temperature have been measured by guiding an ion beam through a  $300$  K target gas cell in experiments GIB (Burley et al. 1987) and TQDO (Li et al. 1997). Their agreement with  $k_L$  below and deviations from  $k_L$  above  $0.3$  eV are discussed in the relevant papers and in Bulut et al. (2015). Note that the plotted points are effective rate coefficients, as derived from the measured effective cross sections just by multiplying them with the mean relative velocity. Additionally, we have corrected the center of mass energy of Li et al. (1997) for the thermal motion of the neutral target with temperature  $T_{\text{GAS}} \approx 300$  K by adding a term  $(3/2)k_B T_{\text{GAS}} m_{O^+}/(m_{O^+} + m_{H_2})$ , where  $m_{O^+}$  and  $m_{H_2}$  are the  $O^+$  and  $H_2$  masses, respectively. The large filled points around  $300$  K summarize 50 years of ion-molecule reaction studies using well-established techniques such as flowing afterglow (FA) (Fehsenfeld et al. 1967), ion cyclotron resonance (ICR) ion trap (Kim et al. 1975), and selected ion flow tube (SIFT) (Smith et al. 1978). The selected ion flow-drift tube (SIFDT) value was measured at a mean relative kinetic energy of  $0.07$  eV (Federer et al. 1984).

We do not know of any previous experimental data for the studied reaction measured below room temperature. Taking into account only the indicated statistical errors, our ion trap data are slightly below  $k_L$ . On one side this may indicate that a fraction of the collision complexes decays back to the reactants. On the other side, this may be due to a systematic error, which is estimated to be up to 20% for the trapping experiments. From 300 to 30 K the measured reaction rate coefficient is temperature-independent. Below there may be a weak temperature dependence, as also indicated by theoretical results, shown as a solid line (Bulut et al. 2015).

Bulut et al. (2015) used time-dependent wave packet calculations (TDWP) to determine the integral cross sections for reactive collisions of  $O^+$  with  $H_2(j)$  at collision energies from 1 eV down to 1 meV and for  $j = 0, 1$ , and 2 (Bulut et al. 2015). They used a global potential energy surface (PES) calculated for the ground electronic state  $1^4A'$  of  $H_2O^+$  by Martínez et al. (2004). On this surface, reactants can progress to the  $OH^+ + H$  product channel without a barrier. In accordance with this, the cross sections they obtained are in very good agreement with the predictions of the Langevin model for a 300 K population of the rotational states (see Bulut et al. 2015, Figure 8). Important for our low-temperature results is that the theoretical cross sections level off below 10 meV. Thermal rate coefficients have been calculated using these results and accounting for the  $T$ -dependent  $j$ -population of the normal  $H_2$  and of  $H_2$  with thermal population of nuclear spin states. The results are indicated by the solid and dashed-dotted lines in Figure 5, respectively. The difference between the rate coefficients at low temperatures is caused by a small ( $\approx 10\%$ ) but significant decrease of the cross sections, with  $j$  increasing from 0 to 2.

It is too early to speculate whether the rate coefficients for reaction (1) really decrease at low temperatures and why. The authors of Bulut et al. (2015) also do not yet know why the TDWP cross sections depart from the Langevin model below 10 meV. They provide thermal rate coefficients only above 200 K and suspect that the leveling off may be due to errors in the theoretical treatment or due to incorrect long-range behavior of the PES.

## 5. Conclusion and Outlook

The temperature range of experimentally determined rate coefficients for the title reaction has been extended down to a nominal trap temperature of  $T_{22PT} = 10$  K, corresponding to a collisional temperature of 15 K. Overall, the hydrogen abstraction reaction follows the behavior of many exothermic ion-molecule reactions, as described by the Langevin capture model. The values obtained at 300 K are in relatively good agreement with the results of previous experimental studies. Nevertheless, we have to keep in mind that in those previous experiments, the internal excitation of reacting  $O^+$  ions was not specified, with the exception of the TQDO study of Li et al. (1997) and of the GIB study of Burley et al. (1987).









There are some indications that the rate coefficient for the title reaction decreases slightly below 30 K in accordance with the TDWP cross sections, which level off below 10 meV (Bulut et al. 2015). However, to test this, more experiments are needed and the theory also has to be re-examined for collision energies relevant to 10 K. Measurements with para- $H_2$  may reveal additional changes at low temperatures. Also, an extension of

the collision energy below 10 K is within the reach of modern cryogenic traps.

We also obtained rate coefficients for the reactions (2) and (3), leading finally to  $H_3O^+$ . The values obtained at room temperature are in good agreement with previous results from SIFT (Jones et al. 1981), FA (Fehsenfeld et al. 1967), ICR (Kim et al. 1975), and VT-SIFT (Martinez et al. 2015) experiments. The results derived from the fits at low temperatures indicate that there are no surprising changes in this reaction rate coefficient. The results of our detailed studies of reactions (2) and (3) with directly injected and cooled  $OH^+$  and  $H_2O^+$  ions have been published (Tran et al. 2018).

We thank the Technische Universität Chemnitz and the Deutsche Forschungsgemeinschaft for lending the 22-pole trap instrument to the Charles University. We thank prof. David Smith for discussion and helpful suggestions. This work was partly supported by the Czech Science Foundation (GACR 15-15077S, GACR 17-19459S, GACR 17-18067S), and by the Charles University (project Nr. GAUK 1584217, 1144616, 1168216).

## ORCID iDs

Artem Kovalenko  <https://orcid.org/0000-0001-9521-6821>  
 Thuy Dung Tran  <https://orcid.org/0000-0002-9894-1647>  
 Serhiy Rednyk  <https://orcid.org/0000-0002-0408-0170>  
 Štěpán Roučka  <https://orcid.org/0000-0002-2419-946X>  
 Petr Dohnal  <https://orcid.org/0000-0003-0341-0382>  
 Radek Plašil  <https://orcid.org/0000-0001-8520-8983>  
 Dieter Gerlich  <https://orcid.org/0000-0002-3550-305X>  
 Juraj Glosík  <https://orcid.org/0000-0002-2638-9435>

## References

- Asvany, O., & Schlemmer, S. 2009, *IJMSp*, 279, 147  
 Bulut, N., Castillo, J. F., Jambrina, P. G., et al. 2015, *JPCA*, 119, 11951  
 Burley, J. D., Ervin, K. M., & Armentrout, P. B. 1987, *IJMP*, 80, 153  
 Chakrabarty, S., Holz, M., Campbell, E. K., et al. 2013, *JPCL*, 4, 4051  
 Chase, M. W. 1998, NIST-JANAF Thermochemical Tables (4th ed.; New York: AIP)  
 de Ruelle, N., Miller, K. A., O'Connor, A. P., et al. 2016, *ApJ*, 816, 31  
 Endres, E. S., Egger, G., Lee, S., et al. 2017, *JMoSp*, 332, 134  
 Federer, W., Villinger, H., Howorka, F., et al. 1984, *PhRvL*, 52, 2084  
 Fehsenfeld, F. C., Schmeltekopf, A. L., & Ferguson, E. E. 1967, *JChPh*, 46, 2802  
 Flesch, G. D., & Ng, C. Y. 1991, *JChPh*, 94, 2372  
 Gerin, M., Neufeld, D. A., & Goicoechea, J. R. 2016, *ARA&A*, 54, 181  
 Gerlich, D. 2008, Low Temperatures and Cold Molecules (Singapore: Imperial College Press), 121  
 Gerlich, D., Borodi, G., Luca, A., Mogo, C., & Smith, M. A. 2011, *ZPC*, 225, 475  
 Gerlich, D., & Horning, S. 1992, *ChRv*, 92, 1509  
 Gerlich, D., Jusko, P., Roučka, Š., et al. 2012, *ApJ*, 749, 22  
 Gerlich, D., Plašil, R., Zymak, I., et al. 2013, *JPCA*, 117, 10068  
 Glosík, J., Rakshit, A. B., Twiddy, N. D., Adams, N. G., & Smith, D. 1978, *JPhB*, 11, 3365  
 Godefroid, M., & Fischer, C. F. 1984, *JPhB*, 17, 681  
 Gómez-Carrasco, S., Godard, B., Lique, F., et al. 2014, *ApJ*, 794, 33  
 Haney, M. A., & Franklin, J. L. 1969, *JChPh*, 50, 2028  
 Hierl, P. M., Dotan, I., Seeley, J. V., et al. 1997, *JChPh*, 106, 3540  
 Hollenbach, D., Kaufman, M. J., Bergin, E. A., & Melnick, G. J. 2009, *ApJ*, 690, 1497  
 Hollenbach, D., Kaufman, M. J., Neufeld, D., Wolfire, M., & Goicoechea, J. R. 2012, *ApJ*, 754, 105  
 Indriolo, N., Neufeld, D. A., Gerin, M., et al. 2015, *ApJ*, 800, 40  
 Jensen, M. J., Bilodeau, R. C., Safvan, C. P., et al. 2000, *ApJ*, 543, 764  
 Johnsen, R., & Biondi, M. A. 1980, *JChPh*, 73, 190

- Jones, J. D. C., Birkinshaw, K., & Twiddy, N. D. 1981, [CPL](#), **77**, 484
- Kim, J. K., Theard, L. P., & Huntress, W. T. 1975, [JChPh](#), **62**, 45
- Lauzin, C., Jacovella, U., & Merkt, F. 2015, [MolPh](#), **113**, 3918
- Le Garrec, J.-L., Carles, S., Speck, T., et al. 2003, [CPL](#), **372**, 485
- Li, X., Huang, Y.-L., Flesch, G. D., & Ng, C. Y. 1997, [JChPh](#), **106**, 564
- Liu, J., Salumbides, E. J., Hollenstein, U., et al. 2009, [JChPh](#), **130**, 174306
- Martínez, O., Ard, S. G., Li, A., et al. 2015, [JChPh](#), **143**, 114310
- Martínez, R., Millán, J., & González, M. 2004, [JChPh](#), **120**, 4705
- Mulin, D., Roučka, Š., Jusko, P., et al. 2015, [PCCP](#), **17**, 8732
- Muller, S., Müller, H. S. P., Black, J. H., et al. 2016, [A&A](#), **595**, A128
- Neau, A., Al Khalili, A., Rosén, S., et al. 2000, [JChPh](#), **113**, 1762
- Neufeld, D. A., & Wolfire, M. G. 2016, [ApJ](#), **826**, 183
- Neufeld, D. A., & Wolfire, M. G. 2017, [ApJ](#), **845**, 163
- Pilbratt, G. L., Riedinger, J. R., Passvogel, T., et al. 2010, [A&A](#), **518**, L1
- Plašil, R., Zymak, I., Jusko, P., et al. 2012, [RSPTA](#), **370**, 5066
- Sansonetti, J. E., & Martin, W. C. 2005, [JPCRD](#), **34**, 1559
- Smith, D., Adams, N. G., & Miller, T. M. 1978, [JChPh](#), **69**, 308
- Tran, T. D., Rednyk, S., Kovalenko, A., et al. 2018, [ApJ](#), **854**, 25
- Wiedmann, R. T., Tonkyn, R. G., White, M. G., Wang, K., & McKoy, V. 1992, [JChPh](#), **97**, 768
- Wyrowski, F., van der Tak, F., Herpin, F., et al. 2010, [A&A](#), **521**, L34
- Zeippen, C. J. 1987, [A&A](#), **173**, 410
- Zymak, I., Hejduk, M., Mulin, D., et al. 2013, [ApJ](#), **768**, 86
- Zymak, I., Jusko, P., Roučka, Š., et al. 2011, [EPJAP](#), **56**, 24010



Article III: Formation of  $\text{H}_2\text{O}^+$  and  $\text{H}_3\text{O}^+$  Cations in Reactions of  $\text{OH}^+$  and  $\text{H}_2\text{O}^+$  with  $\text{H}_2$ : Experimental Studies of the Reaction Rate Coefficients from  $T = 15$  to 300 K

Tran, T. D., Rednyk, S., Kovalenko, A., Roučka, Š., Dohnal, P., Plašil, R.,  
Gerlich, J., Glosík, J.

2018

*The Astrophysical Journal*, 854(1), 25(1-5)



# Formation of $\text{H}_2\text{O}^+$ and $\text{H}_3\text{O}^+$ Cations in Reactions of $\text{OH}^+$ and $\text{H}_2\text{O}^+$ with $\text{H}_2$ : Experimental Studies of the Reaction Rate Coefficients from $T = 15$ to 300 K

Thuy Dung Tran , Serhiy Rednyk , Artem Kovalenko , Štěpán Roučka , Petr Dohnal , Radek Plašil ,  
Dieter Gerlich , and Juraj Glosík

Department of Surface and Plasma Science, Faculty of Mathematics and Physics, Charles University, V Holešovičkách 2,  
Prague, 180 00, Czech Republic; [stepan.roucka@mff.cuni.cz](mailto:stepan.roucka@mff.cuni.cz)

Received 2017 October 24; revised 2017 December 8; accepted 2017 December 8; published 2018 February 7

## Abstract

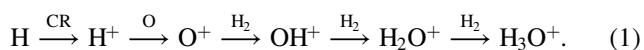
$\text{OH}^+$  and  $\text{H}_2\text{O}^+$  cations play a significant role in the chemistry of the cold interstellar medium and hence their hydrogen abstraction reactions with  $\text{H}_2$  have to be included in ion chemical models. The reactions lead directly or indirectly to  $\text{H}_3\text{O}^+$  ions that subsequently recombine with electrons and dissociate into H atoms and  $\text{H}_2\text{O}$ . The experiments described in this paper provide rate coefficients ( $k_{\text{OH}^+}$  and  $k_{\text{H}_2\text{O}^+}$ ) for the reactions of  $\text{OH}^+$  and  $\text{H}_2\text{O}^+$  with  $\text{H}_2$  over a wide temperature range (from 15 to 300 K). A cryogenic 22-pole RF ion trap instrument was employed for this purpose. It was found that  $k_{\text{OH}^+}$  increases from  $(0.76 \pm 0.30) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  at 17 K to  $(1.24 \pm 0.25) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  at 263 K while  $k_{\text{H}_2\text{O}^+}$  is nearly constant, varying from  $(0.93 \pm 0.35) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  at 17 K to  $(1.00 \pm 0.25) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  at 218 K.

*Key words:* astrochemistry – ISM: molecules – methods: laboratory: molecular – molecular data – molecular processes

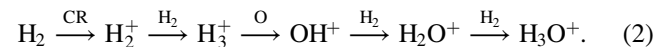
## 1. Introduction

Hydrides, molecules containing a heavy atom and one or more hydrogen atoms, were among the first molecules observed in interstellar space, oxygen hydrides being one of the key species (see recent reviews of Hollenbach et al. 2009; van Dishoeck et al. 2013, 2014; Gerin et al. 2016 and references therein). Major progress in the astronomical studies of oxygen hydrides recently has been made in connection with the Herschel space observatory, the Atacama Large Millimeter/submillimeter Array, the Atacama Pathfinder Experiment, and other observatories. Of particular interest is the detection of OH,  $\text{OH}^+$ ,  $\text{H}_2\text{O}$ ,  $\text{H}_2\text{O}^+$ , and  $\text{H}_3\text{O}^+$  in diffuse and dense Galactic interstellar media (Gerin et al. 2010; Ossenkopf et al. 2010; Pilbratt et al. 2010; Wyrowski et al. 2010; Hollenbach et al. 2012; Gómez-Carrasco et al. 2014; Muller et al. 2016; Neufeld & Wolfire 2016). Hence, the formation of neutral and ionized oxygen hydrides under conditions of the interstellar medium has become an important problem in contemporary laboratory astrophysics (Gerin et al. 2016) and has motivated theoretical studies of these reactions (Ard et al. 2014; Bulut et al. 2015; Martinez et al. 2015; Song et al. 2016a, 2016b, 2017). The results also impact determinations of cosmic ray (CR) ionization rates (Hollenbach et al. 2012; Indriolo et al. 2015; Neufeld & Wolfire 2016, 2017).

The dissociative recombination of  $\text{H}_3\text{O}^+$  ions with electrons is an important source of  $\text{H}_2\text{O}$  molecules (see Larsson & Orel 2008 and references therein). The formation of  $\text{OH}^+$ ,  $\text{H}_2\text{O}^+$ , and  $\text{H}_3\text{O}^+$  has been discussed in detail by Hollenbach et al. (2009, 2012). In the hydrogen-rich environment, there are two routes that lead to  $\text{H}_3\text{O}^+$  ions (see also Hollenbach et al. 2009, 2012; van Dishoeck et al. 2014; Gerin et al. 2016). If the H atom concentration is high,  $\text{H}^+$  ions are generated by CR ionization. In a next step,  $\text{O}^+$  ions are formed from O atoms in the slightly endothermic electron transfer to  $\text{H}^+$ . A series of further reactions with  $\text{H}_2$  leads to  $\text{OH}^+$ ,  $\text{H}_2\text{O}^+$ , and finally to  $\text{H}_3\text{O}^+$ . The overall simplified scheme is



The second reaction chain to  $\text{H}_3\text{O}^+$  starts with the formation of  $\text{H}_2^+$  by the CR ionization of  $\text{H}_2$  followed by the formation of  $\text{H}_3^+$  via a reaction with  $\text{H}_2$ . The  $\text{H}_3^+$  ion can eventually recombine with an electron (Glosík et al. 2010; Hejduk et al. 2015) or react by proton transfer with a less abundant O atom. Subsequently, the  $\text{OH}^+$  ions react with  $\text{H}_2$  and a sequence of hydrogen atom abstraction reactions leads to the formation of  $\text{H}_3\text{O}^+$  (Hollenbach et al. 2012; Indriolo & McCall 2013; van Dishoeck et al. 2013):



The formation of  $\text{OH}^+$  and other oxygen hydrides in various astrophysical environments has been investigated by many authors (see, e.g., van Dishoeck et al. 2013, 2014; Indriolo et al. 2015; Muller et al. 2016; Neufeld & Wolfire 2017, and a recent review of Gerin et al. 2016).

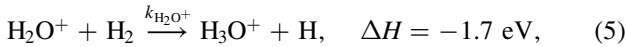
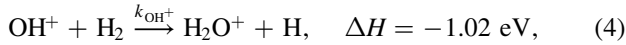
The measurements reported here build on our previous studies in which we used the cryogenic 22-pole RF ion trap apparatus to determine the rate coefficient ( $k_{\text{O}^+}$ ) of the reaction of ground state  $\text{O}^+(^4\text{S})$  ions with  $\text{H}_2$ ,



in the temperature range of 15–300 K. The results and experimental procedures required to obtain the ground state  $\text{O}^+(^4\text{S})$  ions are described in a separate publication (A. Kovalenko et al. 2017, in preparation). The reaction enthalpy  $\Delta H$  at 0 K was calculated using tabulated enthalpies of formation, dissociation energies, and ionization potentials (Wiedmann et al. 1992; Chase 1998; Sansonetti & Martin 2005; Liu et al. 2009). Further theoretical and experimental studies of this reaction would be highly desirable to understand the observed data and also for chemical modeling of the interstellar medium (Indriolo et al. 2015; Markus et al. 2016). The rate coefficient  $k_{\text{O}^+}$  measured by A. Kovalenko et al. (2017, in preparation) is nearly constant in the studied

temperature range with values of  $k_{O^+}$  (15 K) =  $(1.3 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  and  $k_{O^+}$  (300 K) =  $(1.4 \pm 0.3) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  at 15 K and 300 K, respectively. The results are in rather good agreement with flowing-afterglow (FA) data at 300 K (Fehsenfeld et al. 1967), the ion cyclotron resonance (ICR) ion trap data of Kim et al. (1975), as well as the selected ion flow tube (SIFT) data of Smith et al. (1978). The cross-sections measured at collision energies from 0.02 to 7 eV (Li et al. 1997) and the theoretical results (Bulut et al. 2015) are also in good agreement.

The present experimental study is focused on the temperature dependence of the two reactions



where the reaction enthalpies were calculated from the heats of formation (Haney & Franklin 1969; Chase 1998), dissociation energies (Liu et al. 2009), and ionization potentials (Wiedmann et al. 1992; Lauzin et al. 2015). Reaction (4) has been studied several times using various experiments including FA (Fehsenfeld et al. 1967), SIFT (Jones et al. 1981; Shul et al. 1988), and ICR (Kim et al. 1975). Reaction (5) has also been studied experimentally using FA (Fehsenfeld et al. 1967), ICR (Kim et al. 1975), drift tube (DT) (Rakshit & Warneck 1980, 1981), flow-drift tube (FDT) (Dotan et al. 1980), and SIFT (Jones et al. 1981). However, these studies were performed at 300 K with the exception of the FDT experiment that covered relative kinetic energies from 0.04 up to 0.3 eV. More recently the temperature dependencies of  $k_{\text{OH}^+}$  from 200 to 600 K and  $k_{\text{H}_2\text{O}^+}$  from 100 to 600 K have been measured using a variable temperature VT-SIFT experiment (Ard et al. 2014; Martinez et al. 2015). Reactions (4) and (5) have been studied theoretically, including rotational and isotopic effects, for higher collision energies (in the range 0.02–2 eV; Song et al. 2016a, 2016b, 2017).

To our knowledge, there are no experimental studies of the rate coefficients of Reactions (4) and (5) for temperatures below 200 K and 100 K, respectively. The wide range of astrophysical conditions, temperatures in particular, for which the formation and the destruction of the oxygen hydrides are relevant (Hollenbach et al. 2012; van Dishoeck et al. 2014; Indriolo et al. 2015; Gerin et al. 2016; Muller et al. 2016) and the scarcity of experimental data for the temperatures below 300 K are the main motivation for our present study.

In the following text, we shortly describe the experimental arrangement and parameters. The results are presented and discussed in Section 3. They are compared with values from previous experiments and with available theoretical predictions.

## 2. Experiment

The experiments were carried out using a cryogenic 22-pole RF ion trap (AB-22PT instrument). This apparatus allows us to study the interaction of a neutral reactant gas with a cold thermal ensemble of ions confined in a radiofrequency electric field. It has been described in detail many times (see e.g., Gerlich et al. 2011, 2012, 2013; Zymak et al. 2011, 2013; Plašil et al. 2012; Mulin et al. 2015). The body of the 22-pole ion trap is connected to a cold head, which can reach nominal

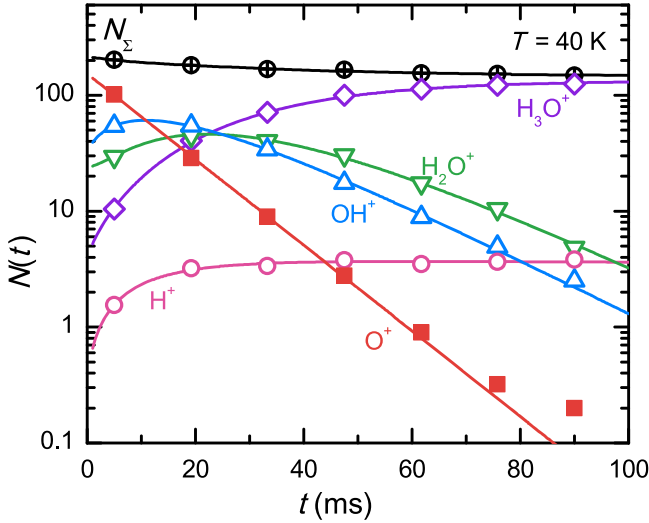
temperatures ( $T_{22\text{PT}}$ ) as low as 10 K. It can be safely assumed on the basis of our previous tests (Plašil et al. 2012; Zymak et al. 2013; Mulin et al. 2015; Š. Roučka et al. 2017, in preparation) that the temperature of the buffer and reactant gases (He and H<sub>2</sub>) in the volume of the trap does not exceed the measured trap temperature by more than 10 K, hence the gas temperature is taken as  $T_{\text{GAS}} = T_{22\text{PT}} + (5 \pm 5) \text{ K}$ . Likewise, the collisional temperature ( $T$ ) of collisions between studied ions and H<sub>2</sub> is taken as  $T = T_{22\text{PT}} + (5 \pm 5) \text{ K}$ . The uncertainties in the temperatures are taken into account when evaluating the gas number densities and consequently also the rate coefficients. For more details and discussion see Paul et al. (1995), Chakrabarty et al. (2013), Jusko et al. (2014), and Endres et al. (2017).

The primary reactants O<sup>+</sup>, OH<sup>+</sup>, and H<sub>2</sub>O<sup>+</sup> are generated by electron impact in a storage ion source using a mixture of N<sub>2</sub>O and H<sub>2</sub> gases. We are using N<sub>2</sub>O because it is safer to handle than O<sub>2</sub> and it has higher vapor pressure than, e.g., H<sub>2</sub>O or H<sub>2</sub>O<sub>2</sub>. The ions are stored for typically 100 ms in the ion source before extraction. Once the source is opened, the ions pass through a mass filter and they are injected into the ion trap filled with a mixture of He and H<sub>2</sub>. The injected ions are cooled in collisions with He atoms. Since the He density is typically 100 times higher than the hydrogen density, the kinetic energy of almost all of the ions is thermalized to  $T_{\text{GAS}}$  prior to a collision with a molecule of H<sub>2</sub>. The uncertainty of the reactant gas pressure in the trap is less than 20%.

After the storage time period, varied in the present studies between 5 and 100 ms, the remaining ions are extracted from the trap, their mass is selected by passing through a quadrupole mass spectrometer, and they are counted by a microchannel plate detector. The number of detected ions is proportional to the number of ions of a particular mass in the trap at the moment of opening. The detection efficiency depends on the mass of the ions. However, in the majority of the present experiments we detect O<sup>+</sup>, OH<sup>+</sup>, H<sub>2</sub>O<sup>+</sup>, and H<sub>3</sub>O<sup>+</sup> ions, whose masses are similar and therefore in the first approximation, the effect of the mass discrimination can be neglected. Furthermore, reaction rates are inferred mainly from the decay of the number of primary ions, so the resulting rate coefficients are not affected by the mass discrimination. Nevertheless, we always monitor the mass discrimination by comparing the observed rate of disappearance of reactant ions with the rate of formation of product ions.

We used normal hydrogen, which is a 300 K statistical mixture of nuclear spin states containing one-fourth of para-H<sub>2</sub> and three-fourths of ortho-H<sub>2</sub> with total nuclear spins of  $I = 0$  and  $I = 1$ , respectively. If hydrogen is cooled down from 300 K without a catalyst, the para to ortho ratio stays constant, as confirmed by chemical probing with N<sup>+</sup> ions in the present experimental setup (Zymak et al. 2013).

In principle, the reaction rate coefficients  $k_{O^+}$ ,  $k_{\text{OH}^+}$ , and  $k_{\text{H}_2\text{O}^+}$  can all be determined by injecting O<sup>+</sup> ions into the trap and observing the whole chain of hydrogen abstraction reactions at once. Such data are shown in Figure 1 for illustration. In this experiment, a small fraction of O<sup>+</sup> was in excited metastable states O<sup>+(2D)</sup> and O<sup>+(2P)</sup> and these ions produced H<sup>+</sup> in reactions with H<sub>2</sub> (see our related publication concerning this reaction for details; A. Kovalenko et al. 2017, in preparation). Figure 1 also shows the total number of ions in the trap ( $N_{\Sigma}(t)$ ). The slight decrease of the sum over time, which is partly due to loss of H<sup>+</sup> from the trap and partly due to



**Figure 1.** Measured decay of the number of the primary  $O^+$  ions and the time evolutions of the numbers of the  $H^+$ ,  $OH^+$ ,  $H_2O^+$ , and  $H_3O^+$  product ions.  $N_{\Sigma}(t)$  is the sum of the numbers of all of the ions in the trap. The experiment was carried out under the following conditions: collisional temperature  $T = 40 \pm 5$  K, gas number densities  $[H_2] = 6.4 \times 10^{10} \text{ cm}^{-3}$ , and  $[He] = 6.5 \times 10^{12} \text{ cm}^{-3}$ . The lines are computed solutions of the adequate rate equations, using the fitted rate coefficients. The rate coefficients and their overall uncertainties are  $k_{O^+}(40 \text{ K}) = (1.33 \pm 0.34) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ ,  $k_{OH^+}(40 \text{ K}) = (8.8 \pm 2.3) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ , and  $k_{H_2O^+}(40 \text{ K}) = (9.5 \pm 2.5) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ .

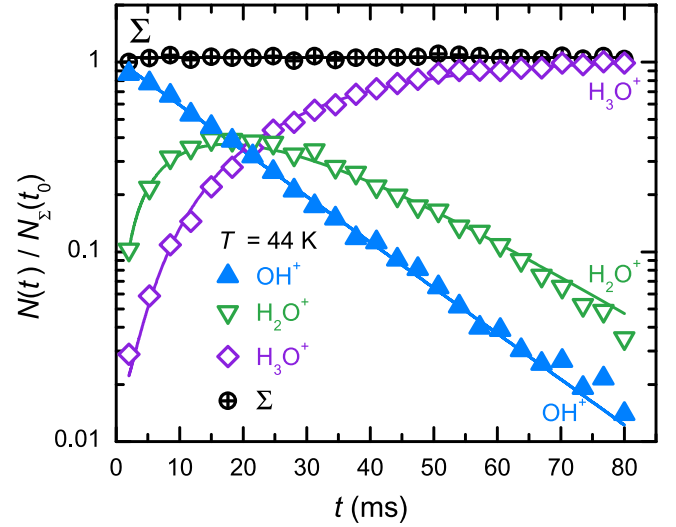
mass discrimination, is taken into account in the data analysis. Minor parasitic reactions due to impurities are observed when only pure He is admitted into the trap. This background loss rate is less than 2% of the reaction loss rate and it is subtracted in the data analysis.

The rate coefficients of Reactions (4) and (5) can be measured directly by producing these ions in the ion source and injecting them into the trap. This is more accurate because we do not have two or more simultaneous processes influencing the number of a particular ion in the trap and, furthermore, the intermediate products  $OH^+$  and  $H_2O^+$  in the reaction chain can be produced with internal excitation, which can influence their reaction rates. In the data graphs, the numbers of ions are divided by the total number of trapped ions measured at time  $t_0$  (time of the first measurements after injection).

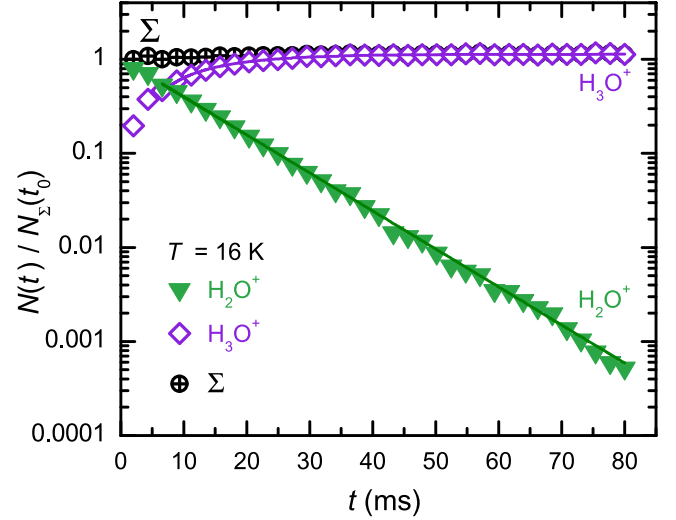
### 3. Results and Discussion

Typical examples of the time evolutions of the normalized numbers of the primary  $OH^+$  ions, the product  $H_2O^+$  ions, and the secondary product  $H_3O^+$  ions measured at the collisional temperature  $T = 44$  K are shown in Figure 2. Also plotted is the sum  $\Sigma(t)$  of the normalized numbers of ions in the trap. Its nearly constant value indicates low losses of ions in the trap due to their reaction with residual gases and gases penetrating from the ion source. Note the mono-exponential decay of the number of  $OH^+$  ions in the trap due to the reaction with  $H_2$ , which indicates good thermalization of the ions. The reaction rate coefficients  $k_{OH^+}$  and  $k_{H_2O^+}$  were obtained simultaneously from the fit of the data.

The rate coefficient  $k_{H_2O^+}$  for the reaction of  $H_2O^+$  with  $H_2$  was also obtained in the simpler and more accurate way, which consists of generating  $H_2O^+$  ions in the ion source. Figure 3 shows the mono-exponential decay of  $H_2O^+$  ions obtained in this type of experiment.



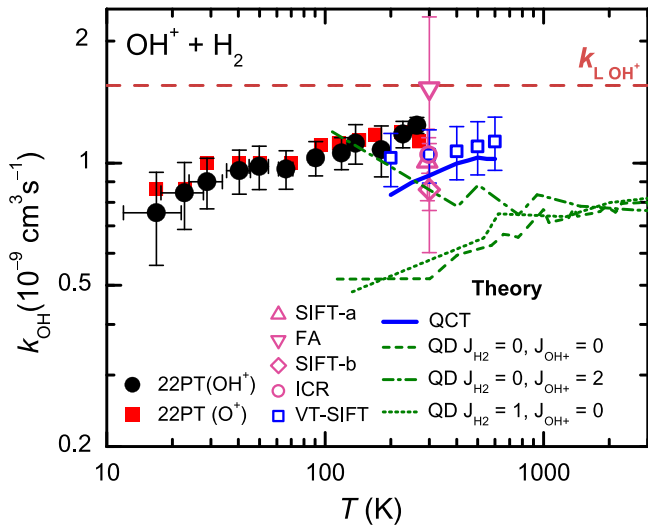
**Figure 2.** Measured time evolution of the normalized numbers of the primary  $OH^+$  and product  $H_2O^+$  and  $H_3O^+$  ions. The lines indicate solutions to the set of rate equations. The open circles indicate the sum ( $\Sigma$ ) of the normalized numbers of ions in the trap. The experimental parameters are  $T = 44 \pm 5$  K,  $[H_2] = 6.4 \times 10^{10} \text{ cm}^{-3}$ , and  $[He] = 1.8 \times 10^{13} \text{ cm}^{-3}$ . The fitted reaction rate coefficients and their overall uncertainties are  $k_{OH^+}(44 \text{ K}) = (8.7 \pm 2.2) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  and  $k_{H_2O^+}(44 \text{ K}) = (9.4 \pm 2.4) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ .



**Figure 3.** Time evolution of the normalized relative numbers of the primary  $H_2O^+$  and secondary  $H_3O^+$  ions. The lines indicate solutions to the set of rate equations. The open circles indicate the sum ( $\Sigma$ ) of the normalized numbers of ions in the trap. The experimental parameters are:  $T = 16 \pm 5$  K,  $[H_2] = 1.0 \times 10^{11} \text{ cm}^{-3}$ , and  $[He] = 2.9 \times 10^{13} \text{ cm}^{-3}$ . The obtained reaction rate coefficient is  $k_{H_2O^+}(16 \text{ K}) = (9.2 \pm 3.2) \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ .

The binary character of the studied reactions was confirmed by measuring the dependencies of the reaction rates on the reactant gas number density  $[H_2]$ . The linearity of the measured dependencies confirmed that the loss of  $OH^+$  (and  $H_2O^+$ ) ions in the trap was caused by a binary ion–molecule reaction with  $H_2$ .

The measured temperature dependence of the rate coefficient  $k_{OH^+}$  of the reaction of  $OH^+$  with normal  $H_2$  is shown in Figure 4. The reaction rate coefficients obtained from experiments with  $OH^+$  injection (see Figure 2) and  $O^+$  injection (see Figure 1) are compared in the figure. Both temperature dependencies are in very good agreement in the

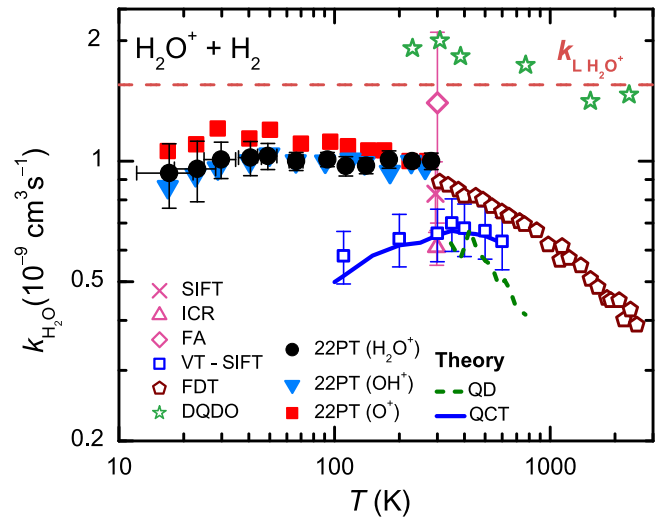


**Figure 4.** Temperature dependence of the rate coefficient  $k_{\text{OH}^+}$  of the reaction of  $\text{OH}^+$  with normal hydrogen. The averaged data obtained in experiments with  $\text{OH}^+$  and  $\text{O}^+$  ions injected into the trap are indicated by full circles and squares, respectively. The systematic error due to pressure measurement is 20%. The dashed horizontal line ( $k_{\text{L,OH}^+}$ ) indicates the Langevin collisional rate coefficient. The previous results at 300 K are FA (Fehsenfeld et al. 1967), ICR (Kim et al. 1975), SIFT-a (Jones et al. 1981), and SIFT-b (Shul et al. 1988). The temperature dependencies of  $k_{\text{OH}^+}$  calculated (QCT) and measured (VT-SIFT) by Martínez et al. (2015) are indicated by the full line and open squares, respectively. The dashed, dotted, and dash-dotted lines represent the phenomenological rate coefficients ( $\nu\sigma$ ) derived from the theoretical QD cross-sections (Song et al. 2016a) corresponding to different rotational states of reactants as indicated in the legend.

whole covered temperature range, indicating that a possible excitation of the intermediate products does not influence the present experimental results. The error bars of our data with  $\text{OH}^+$  injection indicate the statistical errors combined with the systematic uncertainty due to temperature measurement. The uncertainties of the data with  $\text{O}^+$  injection are of similar size.

Figure 4 also shows  $k_{\text{OH}^+}$  measured in several previous experiments at 300 K (FA, Fehsenfeld et al. 1967; ICR, Kim et al. 1975; SIFT, Jones et al. 1981; SIFT, Shul et al. 1988), the temperature dependent data from the VT-SIFT experiment (Martínez et al. 2015), and the theoretical temperature dependence obtained using the quasiclassical trajectory method (QCT; Martínez et al. 2015). For qualitative comparison with our data, the cross-sections ( $\sigma$ ) of Song et al. (2016a) obtained using quantum dynamics (QD) for  $\text{OH}^+$  and  $\text{H}_2$  in rotational states ( $J_{\text{H}_2} = 0, J_{\text{OH}^+} = 0$ ); ( $J_{\text{H}_2} = 0, J_{\text{OH}^+} = 2$ ); ( $J_{\text{H}_2} = 1, J_{\text{OH}^+} = 0$ ) were converted to phenomenological rate coefficients by multiplying them with the relative velocity ( $\nu$ ).

The agreement with the previously published values measured at 300 K is within the combined accuracy of the older and present values. The agreement of the present results with the experimental and theoretical dependencies obtained by Martínez et al. (2015) is good with a small systematic shift. The reaction rate coefficients calculated from the cross-sections of Song et al. (2016a) for ground state reactants are approximately by factor of 2 lower in comparison with present data, but also in comparison with the previous data measured at 300 K. The higher value of the thermal experimental rate coefficients is probably due to the thermal population of  $\text{OH}^+$  rotational states. The  $\text{OH}^+(J_{\text{OH}^+} = 2)$ , which is still significantly populated at 100 K, has a theoretical reaction rate coefficient almost twice as large as the ground state  $\text{OH}^+(J_{\text{OH}^+} = 0)$



**Figure 5.** Temperature dependence of the reaction rate coefficient  $k_{\text{H}_2\text{O}^+}$  of the reaction of  $\text{H}_2\text{O}^+$  with normal hydrogen. The averaged data obtained in experiments with  $\text{H}_2\text{O}^+$ ,  $\text{OH}^+$ , and  $\text{O}^+$  ions injected into the trap are indicated by full circles, triangles, and squares, respectively. The systematic error due to pressure determination is 20%. The dashed horizontal line ( $k_{\text{L,H}_2\text{O}^+}$ ) indicates the Langevin collisional rate coefficient. The previous results at 300 K are FA (Fehsenfeld et al. 1967), ICR (Kim et al. 1975), FDT (Dotan et al. 1980), and SIFT (Jones et al. 1981). The values measured (VT-SIFT) and calculated (QCT) by Ard et al. (2014) are indicated by the open squares and by the full line (QCT), respectively. The dashed line and stars represent the phenomenological rate coefficients ( $\nu\sigma$ ) derived from the theoretical QD, and experimental DQDO cross-sections (Song et al. 2016b). The uncertainty of the DQDO results is 50%.

(Song et al. 2016a). The decrease of our measured reaction rate coefficient toward low temperatures may also be explained by the enhancement of the reaction rate coefficients via rotational excitation of  $\text{OH}^+$ , because about 90% of  $\text{OH}^+$  is in the ground state at 15 K. The rotational excitation of  $\text{H}_2$  has a much smaller effect according to the theory (Song et al. 2016a), even though it contains more energy. The influence of the  $\text{H}_2$  rotational excitation will be further investigated by measuring with para-enriched hydrogen (Zymak et al. 2013).

The temperature dependencies of the reaction rate coefficient  $k_{\text{H}_2\text{O}^+}$  measured in the experiments with the injection of  $\text{O}^+$ ,  $\text{OH}^+$ , and  $\text{H}_2\text{O}^+$  ions into the trap are shown in Figure 5. The experiments were carried out at nominal trap temperatures from 10 to 300 K. The results of experiments with the injection of  $\text{O}^+$  and  $\text{OH}^+$  ions are in agreement over the whole temperature range with the most reliable temperature dependence measured with the injection of  $\text{H}_2\text{O}^+$  ions (see the example in Figure 3). Figure 5 also includes  $k_{\text{H}_2\text{O}^+}$  measured in previous experiments at 300 K (FA, Fehsenfeld et al. 1967; ICR, Kim et al. 1975; FDT, Dotan et al. 1980; SIFT, Jones et al. 1981) and temperature dependent theoretical (QCT) and experimental (VT-SIFT) data between 100 and 600 K by Ard et al. (2014). The cross-sections of Song et al. (2016b) obtained theoretically using the QD method and experimentally using the double-quadrupole double-octopole (DQDO) apparatus are averaged over the thermal population of the  $\text{H}_2\text{O}^+$  rotational states and represented by phenomenological rate coefficients.

The agreement with previous values obtained at 300 K is also within the combined accuracy of the old and present values. However, our results show that the rate coefficient is not decreasing with temperature as claimed by Ard et al. (2014). Their experimental and theoretical data are systematically lower than the present ion trap data and also the

previously published FDT results at higher temperatures (Dotan et al. 1980).








#### 4. Conclusion and Outlook

We have reported results of experimental studies of the temperature dependencies of the rate coefficients for the  $\text{OH}^+ + \text{H}_2$  and  $\text{H}_2\text{O}^+ + \text{H}_2$  reactions using the cryogenic 22-pole RF ion trap with the electron impact storage ion source. The data cover a much wider temperature range (15–300 K) than previous experimental results. The agreement of previous and present data in the overlapping temperature range confirms the accuracy of our experimental methods and we recommend incorporating these new low-temperature data into astrophysical models. We are preparing further low-temperature studies of the reactions of  $\text{O}^+$ ,  $\text{OH}^+$ , and  $\text{H}_2\text{O}^+$  with  $\text{D}_2$  and also with HD (deuterium hydride), and also studies of reactivity with  $\text{H}_2$  in ortho and para nuclear spin configurations. The latter studies may reveal effects of the rotational state of  $\text{H}_2$  that have been predicted.

We thank the Technische Universität Chemnitz and the Deutsche Forschungsgemeinschaft for lending the 22-pole ion trap instrument to the Charles University. We are grateful to professor Rainer Johnsen for consultation. This work was partly supported by the Czech Science Foundation (GACR 14-14715P, GACR 15-15077S, GACR 17-19459S) and by the Charles University (project Nr. GAUK 572214, 1584217, 1144616, 1168216).

*Software:* Imfit (Newville et al. 2014).

#### ORCID iDs

Thuy Dung Tran  <https://orcid.org/0000-0002-9894-1647>  
 Serhiy Rednyk  <https://orcid.org/0000-0002-0408-0170>  
 Artem Kovalenko  <https://orcid.org/0000-0001-9521-6821>  
 Štěpán Roučka  <https://orcid.org/0000-0002-2419-946X>  
 Petr Dohnal  <https://orcid.org/0000-0003-0341-0382>  
 Radek Plašil  <https://orcid.org/0000-0001-8520-8983>  
 Dieter Gerlich  <https://orcid.org/0000-0002-3550-305X>  
 Juraj Glosík  <https://orcid.org/0000-0002-2638-9435>

#### References

- Ard, S. G., Li, A., Martinez, O., et al. 2014, *JPCA*, **118**, 11485  
 Bulut, N., Castillo, J. F., Jambrina, P. G., et al. 2015, *JPCA*, **119**, 11951  
 Chakrabarty, S., Holz, M., Campbell, E. K., et al. 2013, *JPhCh Lett.*, **4**, 4051  
 Chase, M. W. (ed.) 1998, NIST-JANAF Thermochemical Tables (Woodbury, NY: AIP)  
 Dotan, I., Lindinger, W., Rowe, B., et al. 1980, *CPL*, **72**, 67  
 Endres, E. S., Egger, G., Lee, S., et al. 2017, *JMoSp*, **332**, 134  
 Fehsenfeld, F. C., Schmeltekopf, A. L., & Ferguson, E. E. 1967, *JChPh*, **46**, 2802  
 Gerin, M., De Luca, M., Black, J., et al. 2010, *A&A*, **518**, L110  
 Gerlich, D., Plašil, R., Zymak, I., et al. 2013, *JPCA*, **117**, 10068  
 Gerlich, D., Borodi, G., Luca, A., Mogo, C., & Smith, M. A. 2011, *ZPC*, **225**, 475  
 Gerlich, D., Jusko, P., Roučka, Š., et al. 2012, *ApJ*, **749**, 22  
 Gerlich, D., Plašil, R., Zymak, I., et al. 2013, *JPCA*, **117**, 10068  
 Glosík, J., Plašil, R., Kotrčík, T., et al. 2010, *MolPh*, **108**, 2253  
 Gómez-Carrasco, S., Godard, B., Lique, F., et al. 2014, *ApJ*, **794**, 33  
 Haney, M. A., & Franklin, J. L. 1969, *JChPh*, **50**, 2028  
 Hejduk, M., Dohnal, P., Rubovič, P., et al. 2015, *JChPh*, **143**, 044303  
 Hollenbach, D., Kaufman, M. J., Bergin, E. A., & Melnick, G. J. 2009, *ApJ*, **690**, 1497  
 Hollenbach, D., Kaufman, M. J., Neufeld, D., Wolfire, M., & Goicoechea, J. R. 2012, *ApJ*, **754**, 105  
 Indriolo, N., & McCall, B. J. 2013, *Chem. Soc. Rev.*, **42**, 7763  
 Indriolo, N., Neufeld, D. A., Gerin, M., et al. 2015, *ApJ*, **800**, 40  
 Jones, J. D. C., Birkinshaw, K., & Twiddy, N. D. 1981, *CPL*, **77**, 484  
 Jusko, P., Asvany, O., Wallerstein, A.-C., Brünken, S., & Schlemmer, S. 2014, *PhRvL*, **112**, 253005  
 Kim, J. K., Theard, L. P., & Huntress, W. T. 1975, *JChPh*, **62**, 45  
 Larsson, M., & Orel, A. E. 2008, Dissociative Recombination of Molecular Ions (Cambridge: Cambridge Univ. Press)  
 Lauzin, C., Jacovella, U., & Merkt, F. 2015, *MolPh*, **113**, 3918  
 Li, X., Huang, Y.-L., Flesch, G. D., & Ng, C. Y. 1997, *JChPh*, **106**, 564  
 Liu, J., Salumbides, E. J., Hollenstein, U., et al. 2009, *JChPh*, **130**, 174306  
 Markus, C. R., Hodges, J. N., Perry, A. J., et al. 2016, *ApJ*, **817**, 138  
 Martinez, O., Ard, S. G., Li, A., et al. 2015, *JChPh*, **143**, 114310  
 Mulin, D., Roučka, Š., Jusko, P., et al. 2015, *PCCP*, **17**, 8732  
 Muller, S., Müller, H. S. P., Black, J. H., et al. 2016, *A&A*, **595**, A128  
 Neufeld, D. A., & Wolfire, M. G. 2016, *ApJ*, **826**, 183  
 Neufeld, D. A., & Wolfire, M. G. 2017, *ApJ*, **845**, 163  
 Newville, M., Stensitzki, T., Allen, D. B., & Ingargiola, A. 2014, LMFIT: Non-linear Least-Square Minimization and Curve-Fitting for Python, Tech. Rep., Zenodo, doi:10.5281/zenodo.11813  
 Ossenkopf, V., Müller, H. S. P., Lis, D. C., et al. 2010, *A&A*, **518**, L111  
 Paul, W., Lücke, B., Schlemmer, S., & Gerlich, D. 1995, *IJMSI*, **149**, 373  
 Pilbratt, G. L., Riedinger, J. R., Passvogel, T., et al. 2010, *A&A*, **518**, L1  
 Plašil, R., Zymak, I., Jusko, P., et al. 2012, *RSPTA*, **370**, 5066  
 Rakshit, A. B., & Warneck, P. 1980, *J. Chem. Soc., Faraday Trans. 2*, **76**, 1084  
 Rakshit, A. B., & Warneck, P. 1981, *JChPh*, **74**, 2853  
 Sansonetti, J. E., & Martin, W. C. 2005, *JPCRD*, **34**, 1559  
 Shul, R. J., Passarella, R., DiFazio, L. T., Keese, R. G., & Castleman, A. W. 1988, *JPhCh*, **92**, 4947  
 Smith, D., Adams, N. G., & Miller, T. M. 1978, *JChPh*, **69**, 308  
 Song, H., Li, A., & Guo, H. 2016a, *JPCA*, **120**, 4742  
 Song, H., Li, A., Guo, H., et al. 2016b, *PCCP*, **18**, 22509  
 Song, H., Li, A., Yang, M., & Guo, H. 2017, *PCCP*, **19**, 17396  
 van Dishoeck, E. F., Bergin, E. A., Lis, D. C., & Lunine, J. I. 2014, Protostars and Planets VI (Tucson, AZ: Univ. Arizona Press), **835**  
 van Dishoeck, E. F., Herbst, E., & Neufeld, D. A. 2013, *ChRv*, **113**, 9043  
 Wiedmann, R. T., Tonkyn, R. G., White, M. G., Wang, K., & McKoy, V. 1992, *JChPh*, **97**, 768  
 Wyrowski, F., van der Tak, F., Herpin, F., et al. 2010, *A&A*, **521**, L34  
 Zymak, I., Hejduk, M., Mulin, D., et al. 2013, *ApJ*, **768**, 86  
 Zymak, I., Jusko, P., Roučka, Š., et al. 2011, *EPJAP*, **56**, 24010

Article IV: Effect of rotational excitation of H<sub>2</sub> on isotopic  
exchange reaction with OD<sup>-</sup> at low temperatures

Roučka, Š., Rednyk, S., Kovalenko, A., Tran, T. D., Plašil, R., Kálosi, Á.,  
Dohnal, P., Gerlich, D., Glosík, J.

2018

*Astronomy & Astrophysics*, 615(L6), 1-5

LETTER TO THE EDITOR

# Effect of rotational excitation of H<sub>2</sub> on isotopic exchange reaction with OD<sup>-</sup> at low temperatures

Š. Roučka, S. Rednyk, A. Kovalenko, T. D. Tran, R. Plašil, Á. Kálosi, P. Dohnal, D. Gerlich, and J. Glošík

Department of Surface and Plasma Science, Faculty of Mathematics and Physics, Charles University, V Holešovičkách 2, 180 00 Prague, Czech Republic  
e-mail: stepan.roucka@mff.cuni.cz

Received 19 April 2018 / Accepted 21 May 2018

## ABSTRACT

**Aims.** This paper presents experimentally obtained rate coefficients for the weakly endothermic reaction OD<sup>-</sup> + H<sub>2</sub> → OH<sup>-</sup> + HD with ortho- and para-hydrogen at astrophysically relevant temperatures between 10 and 300 K.

**Methods.** The reaction was studied with normal and para-enriched (99.5% para-H<sub>2</sub>) hydrogen in a 22-pole ion trap. The measured temperature dependencies of reaction rate coefficients are analyzed using a model which assumes that the rotational energies of the two reactants are equivalent to the translational energy in driving the reaction.

**Results.** At room temperature, the rate coefficients of reactions with both nuclear spin variants reach 7 × 10<sup>-11</sup> cm<sup>3</sup> s<sup>-1</sup>, which is in good agreement with the previous results from ion trap and swarm experiments with normal hydrogen. Cooling down the trap slows down the reaction and leads, at a nominal trap temperature of 11 K, to a rate coefficient below 10<sup>-14</sup> cm<sup>3</sup> s<sup>-1</sup> for para-enriched hydrogen. The fitted reaction endothermicity of 25.3 ± 2.2 meV agrees well with the literature value calculated in the Born-Oppenheimer approximation, ΔH<sup>0</sup> = 24.0 meV. A simpler evaluation procedure, fitting the data with Arrhenius functions, results in <sup>p</sup>k = 16.8 × 10<sup>-11</sup> exp(-234 K/T) cm<sup>3</sup> s<sup>-1</sup> for pure para-hydrogen and <sup>o</sup>k = 9.4 × 10<sup>-11</sup> exp(-101 K/T) cm<sup>3</sup> s<sup>-1</sup> for pure ortho-hydrogen.

**Key words.** astrochemistry – molecular data – molecular processes – methods: laboratory: molecular – ISM: molecules

## 1. Introduction

In the present study we investigate the weakly endothermic H/D exchange reaction



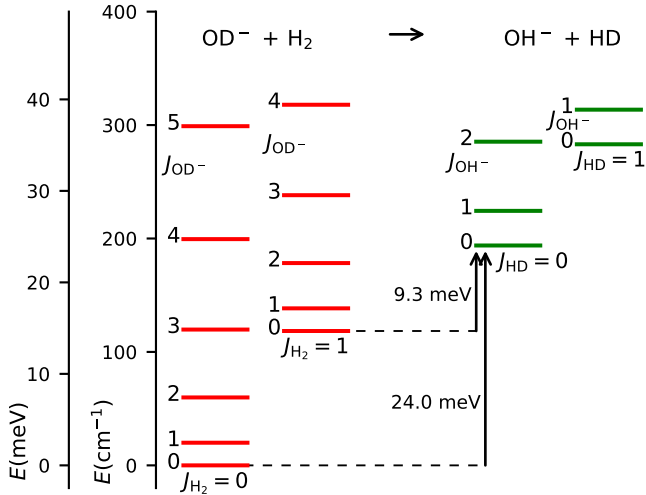
Isotope exchange between different molecules is always nearly thermoneutral; however, due to differences in zero-point energies of the reactants and the products, endothermicities or exothermicities of some tens of meV are obtained. Hence, rate coefficients for H/D exchange tend to be very sensitive to temperatures between 10 and 1000 K, which are typical for the interstellar medium. One consequence is that deuterated molecular species are important tracers for the physical conditions prevailing in the interstellar medium (Roberts & Millar 2000). For example, observations of deuteration have been used, in combination with chemical models, to explore the origin of solar-system molecules (Cleeves et al. 2014b), to constrain the ionization rates in protoplanetary disks (Miettinen et al. 2011; Cleeves et al. 2014a), or to measure the age of prestellar cores (Pagani et al. 2013). The observations of the H/D ratios in molecules are generally concentrating on cations or neutrals, while not much is known about deuteration of anions. Negatively charged ions have recently been observed in space (McCarthy et al. 2006; Cernicharo et al. 2007; Brünken et al. 2007), and their role is increasingly more frequently investigated using astrochemical networks (Walsh et al. 2009; McElroy et al. 2013). In

particular, UMIST-based models (McElroy et al. 2013) predict that OH<sup>-</sup> ions are present in cold dark clouds in quantities comparable to its positive counterpart, OH<sup>+</sup>. While OH and OH<sup>+</sup> are routinely observed (see, e.g., Gerin et al. 2010; Wampfler et al. 2010), OH<sup>-</sup> is still just a candidate for detection in space (Cazzoli & Puzzarini 2006).

The importance of the OH<sup>-</sup> ion is underlined by a number of recent laboratory and theoretical studies of its destruction (Hlavenka et al. 2009; Hauser et al. 2015a), formation (Jusko et al. 2015; Plašil et al. 2017), inelastic collisions (Hauser et al. 2015b), or its rotational spectrum (Cazzoli & Puzzarini 2006; Matsushima et al. 2006; Jusko et al. 2014; Lee et al. 2016). Additionally, the OH<sup>-</sup> anion has become a tool for thermometry of ions in ion traps (Otto et al. 2013; Endres et al. 2017) and storage rings (DESIREE, Schmidt et al. 2017; CSR, Meyer et al. 2017), based on threshold photodetachment spectroscopy.

This work presents rate coefficients measured for reaction (1) with normal and para-enriched H<sub>2</sub> at trap temperatures ranging from 10 to 300 K. Previous studies, performed at room temperature (Grabowski et al. 1983) or in the ranges of 10–300 K (Mulin et al. 2015) and 300–508 K (Viggiano & Morris 1994), used only normal H<sub>2</sub>. Experimental determination of state-specific rate coefficients for J<sub>H<sub>2</sub></sub> = 0 and 1 is essential because the reactivity is highly sensitive to the internal excitation of the reactants. Moreover, the interstellar ortho-/para-H<sub>2</sub> population is often far from thermal equilibrium.





**Fig. 1.** Rotational energy levels of the reactants (left-hand side) and products (right-hand side) for reaction (1). The rotational constants for  $\text{OD}^-$ ,  $\text{OH}^-$ ,  $\text{HD}$ , and  $\text{H}_2$  are taken from the literature (Huber & Herzberg 1979; Rehfuss et al. 1986; Matsushima et al. 2006). The two arrows indicate that 24.0 meV of translational energy are needed to reach the threshold at 0, while 9.3 meV are sufficient for reaction with ortho- $\text{H}_2$  ( $J_{\text{H}_2} = 1$ ).

To compare internal excitation with the endothermicity, Fig. 1 shows the lowest rotational states of the involved diatomic molecules. The nuclear spin states  $I=0$  (para) and 1 (ortho) are linked to the even and odd rotational states, respectively. The 0 K reaction enthalpy of 24 meV was calculated in the Born-Oppenheimer approximation, as discussed in detail by Mulin et al. (2015). At meV accuracy, the energetics are influenced by isotopic electronic shifts of the potential energy surfaces of  $\text{OD}^-$  and  $\text{OH}^-$ , as shown for several other isotope exchange reactions with  $\text{H}_2$  and  $\text{D}_2$  (Kleinman & Wolfsberg 1973, 1974; Adohip-Krou et al. 2004). In particular, spectroscopic studies of OH and OD neutrals suggest that the isotopic shift of electronic ground-state potential energy surface in this system can be up to 2.5 meV (see note 78 in Ruscic et al. 2002). We here also derive an experimental value of the reaction endothermicity.

## 2. Experiment

The experiments have been carried out using the RF 22-pole ion trap machine described by Gerlich (1995, 2008) and Zymak et al. (2011). This instrument has been used to study isotope effects in anion-neutral reactions (Mulin et al. 2015; Roučka et al. 2015; Plašil et al. 2017), therefore we only recall a few specific details. The 22-pole trap was operated at an RF frequency of 17.4 MHz and at amplitudes up to 60 V. The nominal trap temperature ( $T_{22\text{PT}}$ ) was varied from 300 K down to 10 K.

$\text{OD}^-$  ions were produced in an electron impact storage ion source (SIS) using a mixture of  $\text{N}_2\text{O}$  and  $\text{D}_2$  (Jusko et al. 2013, 2014). They were extracted from the source, mass selected by a quadrupole mass filter, and guided into the ion trap, where reactions with  $\text{H}_2$  were followed as a function of time and  $\text{H}_2$  number density. A  $\text{He}/\text{H}_2$  gas mixture was introduced directly into the trap volume. After various well-defined trapping times, the trap was opened and ions were extracted, mass selected by a second quadrupole mass spectrometer, and counted by an MCP detector. The numbers of detected ions were proportional to the numbers of trapped ions, and minor differences in the overall

detection efficiency for  $\text{OD}^-$  and  $\text{OH}^-$  were accounted for in the data analysis.

The addition of helium buffer gas ensures that most of the ions are collisionally thermalized before reacting with  $\text{H}_2$ . Typically used  $\text{He}/\text{H}_2$  mixtures result in ten collisions with  $\text{He}$  for every one collision with  $\text{H}_2$  on average. Moreover, at low temperatures,  $\text{H}_2$  molecules also act as a buffer gas because of the low probability for a reactive collision. Recent experiments in our apparatus (Plašil et al. 2012; Zymak et al. 2013) as well as in other 22-pole trap experiments (Hauser et al. 2015b; Endres et al. 2017) have confirmed that the collisional temperature is slightly higher than the temperature of the trap. In the present case we can safely assume that the collisional temperature in interaction of  $\text{OD}^-$  with  $\text{H}_2$  does not exceed the trap temperature by more than 10 K. For simplicity of presentation, we define the collisional temperature as  $T = T_{22\text{PT}} + 5$  K with an uncertainty of  $\pm 5$  K.

We used normal hydrogen ( $^n\text{H}_2$ ) and para-enriched hydrogen ( $^e\text{H}_2$ ) as reactants, with number densities varying from  $10^{12}$  up to  $10^{13}$   $\text{cm}^{-3}$ . In normal  $\text{H}_2$ , the para/ortho ratio is 1/3, corresponding to the 300 thermal equilibrium. The para-enriched hydrogen, containing 99.5% of para- $\text{H}_2$ , was produced using a para-hydrogen generator with paramagnetic catalyst (Hejduk et al. 2012; Dohnal et al. 2012; Zymak et al. 2013). The amount of ortho impurities was determined in situ using chemical probing with  $\text{N}^+$  ions (Zymak et al. 2013). The nuclear spin-state-specific rate coefficients  $^p k_1$  and  $^o k_1$  for the reaction of  $\text{OD}^-$  with para- and ortho- $\text{H}_2$  were then calculated from the measured rate coefficients  $^n k_1$  and  $^e k_1$  for the reaction with  $^n\text{H}_2$  and with  $^e\text{H}_2$ .

To evaluate the measured time dependencies, especially at low temperatures, we have to take into account reconversion of  $\text{OH}^-$  products back to  $\text{OD}^-$  via the fast exothermic reaction (Mulin et al. 2015)



The  $\text{D}_2$  traces in the trap originate from the ion source. Equilibrium between this backward reconversion reaction and reaction (1) is established at long enough trapping times.

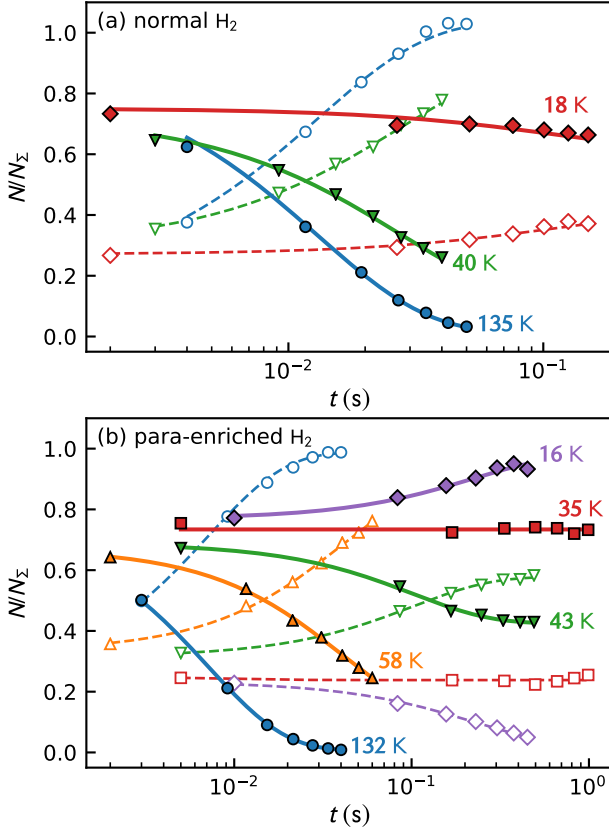
## 3. Results and discussion

The change in numbers of detected  $\text{OD}^-$  and  $\text{OH}^-$  ions in the trap can be described with the balance equations

$$\frac{dN_{\text{OD}^-}}{dt} = -k_1 N_{\text{OD}^-} [\text{H}_2] + k_2 \xi N_{\text{OH}^-} [\text{D}_2], \quad (3)$$

$$\frac{dN_{\text{OH}^-}}{dt} = -k_2 N_{\text{OH}^-} [\text{D}_2] + k_1 \frac{N_{\text{OD}^-}}{\xi} [\text{H}_2], \quad (4)$$

where  $k_1$  and  $k_2$  are the rate coefficients of reactions (1) and (2), and  $[\text{H}_2]$  and  $[\text{D}_2]$  are the number densities of  $\text{H}_2$  and  $\text{D}_2$  in the trap, respectively. The factor  $\xi$ , which is close to 1, accounts for the detection efficiency of  $\text{OD}^-$  relative to  $\text{OH}^-$ . In the following, the additional index n and e is used, that is,  $^n k_i$  and  $^e k_i$ , to indicate reactions with  $^n\text{H}_2$  and  $^e\text{H}_2$ , respectively. The reaction rate coefficients are determined by fitting the measured time dependencies of the numbers of detected ions,  $N_{\text{OD}^-}(t)$  and  $N_{\text{OH}^-}(t)$ , with solution of Eqs. (3) and (4). Free parameters of the fit are the reaction rates  $r_1 = k_1 [\text{H}_2]$  and  $r_2 = k_2 [\text{D}_2]$ , the initial numbers of detected ions,  $N_{\text{OD}^-}(t=0)$  and  $N_{\text{OH}^-}(t=0)$ , and the relative detection efficiency  $\xi$ .



**Fig. 2.** Time dependencies of the normalized numbers of OD<sup>-</sup> (filled symbols) and OH<sup>-</sup> (open symbols), measured with <sup>1</sup>H<sub>2</sub> (panel a) and <sup>2</sup>H<sub>2</sub> (panel b). The collisional temperatures are indicated in the figure, and the densities of <sup>1</sup>H<sub>2</sub>, <sup>2</sup>H<sub>2</sub>, He, and D<sub>2</sub> are listed in Table 1.

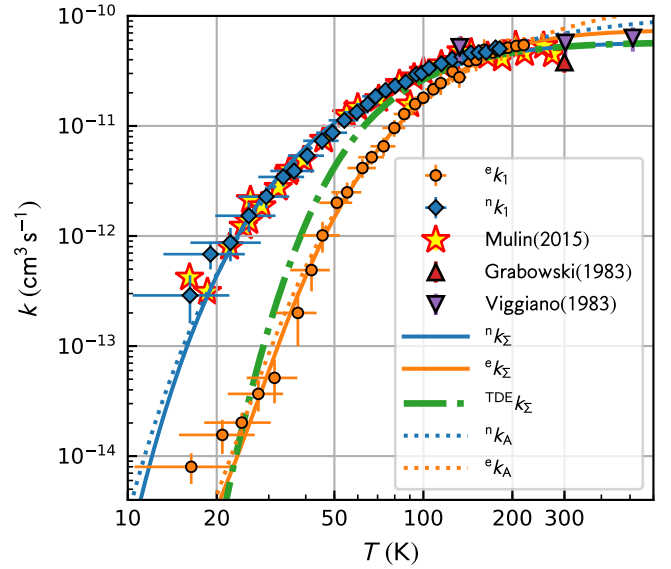
Examples of measured and fitted time dependencies are plotted in Fig. 2. For clarity, the numbers are normalized by the total number of ions,  $N_{\Sigma}$ , measured directly after filling the trap. The upper panel of Fig. 2 shows data obtained with <sup>1</sup>H<sub>2</sub> as target gas, and the lower panel those for <sup>2</sup>H<sub>2</sub>. The corresponding densities of <sup>1</sup>H<sub>2</sub>, <sup>2</sup>H<sub>2</sub>, He, and D<sub>2</sub> in the trap are listed in Table 1. The density of D<sub>2</sub> penetrating from the ion source is typically 1000 times lower than that of H<sub>2</sub>. Nevertheless, the equilibrium numbers of OD<sup>-</sup> are comparable to or even higher than those of OH<sup>-</sup> at the lowest temperatures because of the very slow forward reaction and the fast backward reversion reaction. The effect of backward reversion is even more pronounced for a reaction with para-enriched H<sub>2</sub>. The dependence of the reaction rate  $r_1$  on the hydrogen number density was measured to be linear, confirming the binary character of reaction (1).

Figure 3 compares the rate coefficients,  ${}^n k_1$  and  ${}^e k_1$ , measured between 15 and 300 with previous results for  ${}^n k_1$  from the same experimental arrangement (Mulin et al. 2015; shown as a function of collisional temperature,  $T = T_{22PT} + 5$  K). The agreement is very good (within a few percent). We also show results from the FDT experiment of Viggiano & Morris (1994) and the SIFT experiment of Grabowski et al. (1983). The data we present were simplified by binning the results of nearly 300 measurements into logarithmically spaced bins in temperature. The indicated uncertainties of the reaction rate coefficients take into account the estimated fit errors and the 5 uncertainty of the collisional temperature, which also influences the reactant number density. The uncertainty due to pressure measurement is 20%.

**Table 1.** Experimental conditions (collisional temperatures  $T$ , number densities of H<sub>2</sub>, He, and D<sub>2</sub>, and fractions of ortho-H<sub>2</sub>) used to measure the data shown in Fig. 2.

$T$ K	[H <sub>2</sub> ] 10 <sup>12</sup> cm <sup>-3</sup>	[ <sup>o</sup> H <sub>2</sub> ] [H <sub>2</sub> ]	[He] 10 <sup>12</sup> cm <sup>-3</sup>	[D <sub>2</sub> ] 10 <sup>9</sup> cm <sup>-3</sup>
135	1.8	0.75	16	3.0
40	6.8	0.75	30	5.4
18	10.3	0.75	45	8.2
132	3.7	0.005	22	2.2
58	7.2	0.005	33	3.3
43	8.3	0.005	38	3.8
35	9.2	0.005	42	4.2
16	10.4	0.005	33	6.3

**Notes.** The H<sub>2</sub> and He gases are introduced to the trap volume via the inlet system, and D<sub>2</sub> is diffusing into the trap from the ion source. The uncertainties of the number densities are close to 20%.

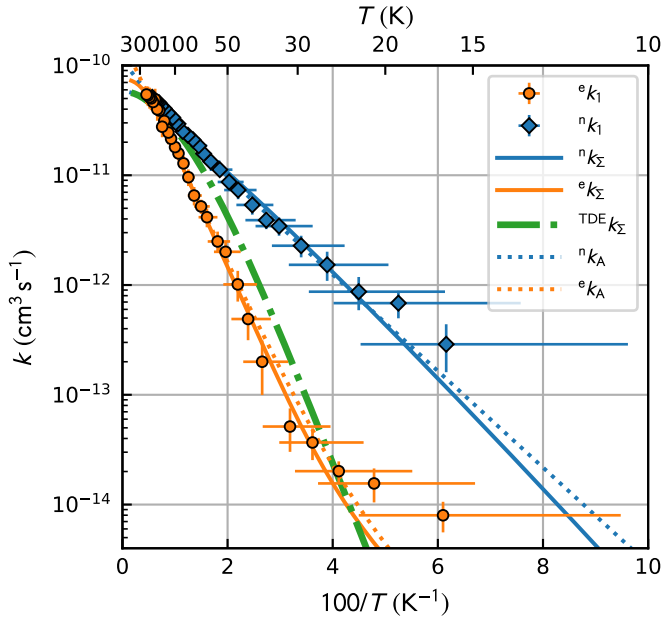


**Fig. 3.** Temperature dependencies of the rate coefficients,  ${}^n k_1$  and  ${}^e k_1$ , measured for reaction (1) with <sup>1</sup>H<sub>2</sub> and with <sup>2</sup>H<sub>2</sub>. We also plot data obtained in our previous experiments with <sup>1</sup>H<sub>2</sub> using the same experimental arrangement (Mulin et al. 2015), FDT data of Viggiano & Morris (1994), and SIFT data of Grabowski et al. (1983). The solid and dotted lines show the result of a state-specific model and an Arrhenius model, respectively. For details, see the text. The dash-dotted line indicates the thermal rate coefficient, i.e., the ortho/para ratio is also equilibrated to  $T$ .

The Arrhenius plot of the measured reaction rate coefficients shown in Fig. 4 reveals a nearly linear dependence of  ${}^n k_1$  and  ${}^e k_1$  over two and three orders of magnitude, respectively. The deviations from linearity for temperatures below 25 are smaller than the estimated temperature uncertainty of  $\pm 5$  K.

We analyzed our data using the simple statistical model described by Mulin et al. (2015), which takes into account any non-equilibrium population of H<sub>2</sub> rotational states. Generally, the rate coefficient  $k_1$  of reaction (1) can be expressed as a weighted sum  $k_{\Sigma}$  of state-specific rate coefficients  $k_{J_{H_2} J_{OD^-}}$ , averaged over the rotational states  $J_{H_2}$  and  $J_{OD^-}$  of H<sub>2</sub> and OD<sup>-</sup>. In particular, we define rate coefficients  ${}^n k_{\Sigma}$  and  ${}^e k_{\Sigma}$  for reaction with normal and para-enriched H<sub>2</sub> as

$${}^{n/e} k_{\Sigma} = \sum_{J_{H_2}, J_{OD^-}} {}^{n/e} P_{J_{H_2}} P_{J_{OD^-}} k_{J_{H_2} J_{OD^-}}, \quad (5)$$



**Fig. 4.** Arrhenius plot of  ${}^e k_1$  and  ${}^n k_1$  measured in experiments with  ${}^e \text{H}_2$  and  ${}^n \text{H}_2$ . The solid and dotted lines show the results from a state-specific model and an Arrhenius model, respectively. The dash-dotted line shows the fully thermalized rate coefficients. For details, see the text.

where  ${}^{n/e} P_{J_{\text{H}_2}}$  and  $P_{J_{\text{OD}^-}}$  are the corresponding populations of rotational states. Our model uses the assumption that all state-specific rate coefficients follow the Arrhenius temperature dependence

$$k_{J_{\text{H}_2} J_{\text{OD}^-}} = {}^{p/o} k_0 \exp\left(-\frac{\Delta E_{J_{\text{H}_2} J_{\text{OD}^-}}}{k_B T}\right), \quad (6)$$

with the activation energy given by

$$\Delta E_{J_{\text{H}_2} J_{\text{OD}^-}} = \max\{0, \Delta H_{\text{fit}}^0 - E_{J_{\text{H}_2}} - E_{J_{\text{OD}^-}}\}. \quad (7)$$

The model assumes that both the rotational energy of  $E_{J_{\text{H}_2}}$  and  $E_{J_{\text{OD}^-}}$  and the translational energy are equivalent in driving the reaction. In Eq. (6) we also assume that there is a global pre-exponential factor  ${}^p k_0$  for reactions with para- $\text{H}_2$  ( $J_{\text{H}_2}$  even) and  ${}^o k_0$  for all reactions of ortho- $\text{H}_2$  ( $J_{\text{H}_2}$  odd). The only free parameters of the model are the reaction endoergicity  $\Delta H_{\text{fit}}^0$  and the pre-exponential factors  ${}^p k_0$  and  ${}^o k_0$ . The model was fitted globally to both temperature dependencies of rate coefficients of reactions with ortho- and para- $\text{H}_2$ , resulting in

$$\Delta H_{\text{fit}}^0 = 25.3_{-2.1}^{+2.4} \text{ meV},$$

$${}^p k_0 = 7.4_{-0.4}^{+0.7} \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}, \quad {}^o k_0 = 5.2_{-1.3}^{+1.3} \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}.$$

These mean values were obtained from fits corresponding to the collisional temperature  $T = T_{22\text{PT}} + 5 \text{ K}$ . The quoted uncertainties include the statistical errors and the systematic errors due to temperature uncertainty, which were estimated by fitting the data with collisional temperatures defined as  $T^- = T_{22\text{PT}}$  and  $T^+ = T_{22\text{PT}} + 10 \text{ K}$ . The pre-exponential factors  ${}^p k_0$  and  ${}^o k_0$  have additional uncertainty of 20% due to the uncertainty of the pressure. The fitted curves are included in Figs. 3 and 4.

The true thermal reaction rate coefficient  ${}^{\text{TDE}} k_{\Sigma}$  plotted in Figs. 3 and 4 was derived from our model by inserting the equilibrium population of all  $\text{H}_2$  rotational states,  ${}^{\text{TDE}} P_{J_{\text{H}_2}}$ , into Eq. (5), that is, also assuming equilibrium ortho/para population.

To facilitate the incorporation of our data into chemical models, we also fit our data with the simple Arrhenius model, described by the equation

$${}^{n/e} k_A = (1 - {}^{n/e} f) {}^p k_A + {}^{n/e} f {}^o k_A, \quad (8)$$

where  ${}^{n/e} f = [{}^o \text{H}_2]/[{}^{n/e} \text{H}_2]$  is the fraction of ortho- $\text{H}_2$  in the normal/para-enriched  $\text{H}_2$  and

$${}^{p/o} k_A = {}^{p/o} k_{A0} \exp(-{}^{p/o} E_A / (k_B T)). \quad (9)$$

The agreement of the fit with our data (Fig. 4) indicates that the linear combination of Arrhenius dependencies (8) is a good approximation. The resulting parameters are

$${}^o E_A = 8.7_{-1.7}^{+2.1} \text{ meV}, \quad {}^o k_{A0} = 9.4_{-2.2}^{+3.1} \times 10^{-11} \text{ cm}^3 \text{ s}^{-1},$$

$${}^p E_A = 20.1_{-2.3}^{+2.7} \text{ meV}, \quad {}^p k_{A0} = 16.8_{-2.7}^{+3.6} \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}.$$

The errors were estimated by the same procedure as discussed above for the state-specific model parameters.

The fitted endothermicity  $\Delta H_{\text{fit}}^0$  is in good agreement with the value  $\Delta H^0 = 24.0 \text{ meV}$  discussed in the Introduction. This confirms that the isotopic energy shifts are not significantly larger than our experimental accuracy ( $\approx 2.5 \text{ meV}$ ). The discrepancy between the Arrhenius activation energy  ${}^p E_A$  and the endothermicities  $\Delta H_{\text{fit}}^0$  and  $\Delta H^0$  is not surprising because  ${}^p E_A$  is an empirical parameter that does not account for the internal excitation of the reactants. Nevertheless, the activation energy  ${}^o E_A$  of the reaction with ortho- $\text{H}_2$  is in good agreement with the Born-Oppenheimer estimate of  $9.3 \text{ meV}$ . It is also in agreement with the activation energy of reaction with normal  $\text{H}_2$ ,  $E_{A\text{-exp}} = (7.9 \pm 0.3) \text{ meV}$  determined by Mulin et al. (2015), which confirms that reaction with ortho- $\text{H}_2$  was the dominant process in that experiment.

## 4. Conclusions

We have extended our previous measurements (Mulin et al. 2015) of the isotope exchange reaction between  $\text{OD}^-$  and  $\text{H}_2$  using nearly pure para-hydrogen. The data allowed us to extract nuclear-spin specific rate coefficients at temperatures from 15 to 300. As expected, the two results are equal at room temperature and agree well with previous data (Viggiano & Morris 1994; Grabowski et al. 1983). Over a wide range of temperatures, the rate coefficients fall with decreasing temperature in accordance with the Arrhenius formula. For astrochemical models, reactions with pure para- or ortho- $\text{H}_2$  can be simply characterized with Arrhenius functions  ${}^p k = 16.8 \times 10^{-11} \exp(-234 \text{ K}/T) \text{ cm}^3 \text{ s}^{-1}$  and  ${}^o k = 9.4 \times 10^{-11} \exp(-101 \text{ K}/T) \text{ cm}^3 \text{ s}^{-1}$ , respectively.

*Acknowledgements.* This work was partly supported by the Czech Grant Agency GACR 17-19459S and by the Charles University (GAUK 1584217, 1144616, 1168216). We thank the Chemnitz University of Technology and the DFG for lending the 22-pole trap instrument to the Charles University.

## References

- Adohi-Krou, A., Martin, F., Ross, A. J., Linton, C., & Le Roy R. J. 2004, *J. Chem. Phys.*, **121**, 6309
- Brünken, S., Gupta, H., Gottlieb, C. A., McCarthy, M. C., & Thaddeus, P. 2007, *ApJ*, **664**, L43
- Cazzoli, G., & Puzzarini, C. 2006, *ApJ*, **648**, L79
- Cernicharo, J., Guélin, M., Agúndez, M., et al. 2007, *A&A*, **467**, L37
- Cleeves, L. I., Bergin, E. A., & Adams, F. C. 2014a, *ApJ*, **794**, 123
- Cleeves, L. I., Bergin, E. A., Alexander, C. M. O., et al. 2014b, *Science*, **345**, 1590

- Dohnal, P., Hejduk, M., Varju, J., et al. 2012, *Phil. Trans. R. Soc. A*, **370**, 5101
- Endres, E. S., Egger, G., Lee, S., et al. 2017, *J. Mol. Spectr.*, **332**, 134
- Gerin, M., De Luca, M., Black, J., et al. 2010, *A&A*, **518**, L110
- Gerlich, D. 1995, *Phys. Scr.*, **1995**, 256
- Gerlich, D. 2008, *Low Temperatures and Cold Molecules* (London: Imperial College Press)
- Grabowski, J. J., DePuy, C. H., & Bierbaum, V. M. 1983, *J. Am. Chem. Soc.*, **105**, 2565
- Hauser, D., Lakhmanskaya, O., Lee, S., Roučka, Š., & Wester, R. 2015a, *New J. Phys.*
- Hauser, D., Lee, S., Carelli, F., et al. 2015b, *Nat. Phys.*, **11**, 467
- Hejduk, M., Dohnal, P., Varju, J., et al. 2012, *Plasma Sources Sci. Technol.*, **21**, 024002
- Hlavenka, P., Otto, R., Trippel, S., et al. 2009, *J. Chem. Phys.*, **130**, 061105
- Huber, K. P., & Herzberg, G. 1979, *Molecular Spectra and Molecular Structure: Constants of Diatomic Molecules* (New York: Van Nostrand Reinhold)
- Jusko, P., Roučka, Š., Plašil, R., & Glosík, J. 2013, *Int. J. Mass Spectrom.*, **352**, 19
- Jusko, P., Asvany, O., Wallerstein, A.-C., Brünken, S., & Schlemmer, S. 2014, *Phys. Rev. Lett.*, **112**, 253005
- Jusko, P., Roučka, Š., Mulin, D., et al. 2015, *J. Chem. Phys.*, **142**, 014304
- Kleinman, L. I., & Wolfsberg, M. 1973, *J. Chem. Phys.*, **59**, 2043
- Kleinman, L. I., & Wolfsberg, M. 1974, *J. Chem. Phys.*, **60**, 4740
- Lee, S., Hauser, D., Lakhmanskaya, O., et al. 2016, *Phys. Rev. A*, **93**, 032513
- Matsushima, F., Yonezu, T., Okabe, T., Tomaru, K., & Moriwaki, Y. 2006, *J. Mol. Spectrosc.*, **235**, 261
- McCarthy, M. C., Gottlieb, C. A., Gupta, H., & Thaddeus, P. 2006, *ApJ*, **652**, L141
- McElroy, D., Walsh, C., Markwick, A. J., et al. 2013, *A&A*, **550**, A36
- Meyer, C., Becker, A., Blaum, K., et al. 2017, *Phys. Rev. Lett.*, **119**, 023202
- Miettinen, O., Hennemann, M., & Linz, H. 2011, *A&A*, **534**, A134
- Mulin, D., Roučka, Š., Jusko, P., et al. 2015, *Phys. Chem. Chem. Phys.*, **17**, 8732
- Otto, R., von Zastrow, A., Best, T., & Wester, R. 2013, *Phys. Chem. Chem. Phys.*, **15**, 612
- Pagani, L., Lesaffre, P., Jorfi, M., et al. 2013, *A&A*, **551**, A38
- Plašil, R., Zymak, I., Jusko, P., et al. 2012, *Phil. Trans. R. Soc. A*, **370**, 5066
- Plašil, R., Tran, T. D., Roučka, Š., et al. 2017, *Phys. Rev. A*, **96**, 062703
- Reh fuss, B. D., Crofton, M. W., & Oka, T. 1986, *J. Phys. Phys.*, **85**, 1785
- Roberts, H., & Millar, T. J. 2000, *A&A*, **361**, 388
- Roučka, Š., Mulin, D., Jusko, P., et al. 2015, *J. Phys. Chem. Lett.*, **6**, 4762
- Ruscic, B., Wagner, A. F., Harding, L. B., et al. 2002, *J. Phys. Chem. A*, **106**, 2727
- Schmidt, H. T., Eklund, G., Chartkunchand, K. C., et al. 2017, *Phys. Rev. Lett.*, **119**, 073001
- Viggiano, A. A., & Morris, R. A. 1994, *J. Chem. Phys.*, **100**, 2748
- Walsh, C., Harada, N., Herbst, E., & Millar, T. J. 2009, *ApJ*, **700**, 752
- Wampfler, S. F., Herczeg, G. J., Bruderer, S., et al. 2010, *A&A*, **521**, L36
- Zymak, I., Jusko, P., Roučka, Š., et al. 2011, *Eur. Phys. J. Appl. Phys.*, **56**, 24010
- Zymak, I., Hejduk, M., Mulin, D., et al. 2013, *ApJ*, **768**, 86

Article V: The reaction of  $O^+(^4S)$  ions with  $H_2$ , HD, and  $D_2$  at low temperatures: Experimental study of the isotope effect.

Kovalenko, A., Roučka, Š., Tran, T.D., Rednyk, S., Plašil, R., Dohnal, P.,  
Gerlich, D., Glosík, J.

*The Astrophysical Journal*

(The paper is under review)

# The reaction of $O^+(^4S)$ ions with $H_2$ , HD, and $D_2$ at low temperatures: Experimental study of isotope effect.

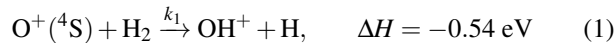
A. Kovalenko,<sup>1</sup> Š. Roučka,<sup>1</sup> T. D. Tran,<sup>1</sup> S. Rednyk,<sup>1</sup> R. Plašil,<sup>1</sup> P. Dohnal,<sup>1</sup> and J. Glosík<sup>1</sup>  
*Charles University, Faculty of Mathematics and Physics, Department of Surface and Plasma Science, V Holešovičkách 2, 180 00 Prague, Czech Republic*

(Dated: December 22, 2020)

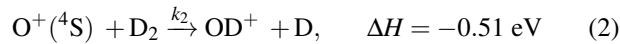
The reactions of the  $^4S$  ground electronic state of  $O^+$  ions with  $D_2$  and HD were studied in a cryogenic 22-pole radio-frequency ion trap in the temperature range of 15 to 300 K. The obtained reaction rate coefficients for both reactions are, considering the experimental errors, nearly independent of temperature and close to the values of the corresponding Langevin collisional reaction rate coefficients. The obtained branching ratios for the production of  $OH^+$  and  $OD^+$  in the reaction of  $O^+(^4S)$  with HD do not change significantly with temperature and are consistent with results obtained at higher collisional energies by other groups. Particular attention was given to ensure that the  $O^+$  ions in the trap are in the ground electronic state.

## I. INTRODUCTION

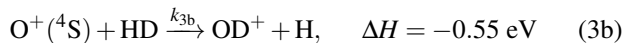
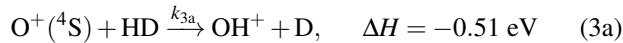
The present study of the reactions of  $O^+$  ions with  $D_2$  and HD is a continuation of our previous study<sup>1</sup> of the reaction of the ground state  $O^+(^4S)$  ions with  $H_2$



where the temperature dependence of the reaction rate coefficient  $k_1$  was measured from 300 K down to 15 K. The present study is focused on finding the temperature dependencies of the rate coefficients of the title reactions



with reaction rate coefficient  $k_2$  and



with reaction rate coefficients  $k_{3a}$  and  $k_{3b}$ , respectively. The overall reaction rate coefficient  $k_3$  is given by the sum of the rate coefficients for both channels;  $k_3 = k_{3a} + k_{3b}$ . The reaction enthalpies  $\Delta H$  at 0 K were calculated from tabulated enthalpies of formation, ionization potentials, and dissociation energies<sup>2-7</sup>.

The main aim of the present study is to determine the temperature dependencies of the reaction rate coefficients for three different isotopologues of  $H_2$  and the branching ratio for products  $OH^+$  and  $OD^+$  of reaction with HD. Despite the fundamental character of all three reactions and their importance for the chemistry of many plasmatic environments, the temperature dependencies of the rate coefficients of the reactions (2) and (3) were not measured up to now for subthermal temperatures and for the reaction (1) such measurements were made only very recently in our laboratory<sup>1</sup>. It is even more surprising if we realize that oxygen is the third most abundant element in the Universe and, as such, it plays a significant role in the chemistry of many astrophysically important environments. The reaction of  $O^+$  with molecular hydrogen, which

is the dominant molecule in many interstellar environments, is a key process in the formation of  $OH^+$  which initiates a sequence of reactions with  $H_2$  that can lead to the formation of  $H_3O^{+8,9}$ . Consequently, an  $H_2O$  molecule can be formed in dissociative recombination of  $H_3O^+$  ion with an electron<sup>10</sup>. If HD or  $D_2$  are involved in some reaction in the sequence, then deuterated or partly deuterated  $H_3O^+$  and  $H_2O$  can be produced. To understand the astronomical observations of these deuterated molecules, more experimental and theoretical studies of reactions with HD and  $D_2$  and more astronomical observations are needed<sup>11-14</sup>.

The paper is structured as follows. In Section II A we briefly describe the experimental apparatus and experimental procedure. In Section II B we discuss the relation of nominal and collisional temperature in the reaction of  $O^+$  ions with reactant molecules. The characterization of the internal excitation of the trapped  $O^+$  ions is highlighted in Section II C. The experimental results for reaction of  $O^+(^4S)$  ground-state ions with  $D_2$  and HD are presented in Section III A and III B, respectively. The influence of the metastable  $O^+(^2D)$  and  $O^+(^2P)$  ions present in small quantities in the ion trap on the measured results is discussed in Section IV A. In Section IV B, the isotopic effects are discussed and the temperature dependencies of the reaction rate coefficients  $k_1$ ,  $k_2$ , and  $k_3$  are compared at low temperatures. The summary is given in Section V.

## II. EXPERIMENTAL

We have used a temperature-variable cryogenic linear 22-pole RF (radiofrequency) ion trap (22PT) to measure the rate coefficients of the title reactions and their temperature dependencies in the temperature range of 15 — 300 K. The apparatus, the principle of ion trapping, and operating have been described previously, so only a very short description will be given here with an emphasis on some specific aspects of the present study (for more details, see Refs. 15–24).

### A. 22PT instrument

The primary  $O^+$  ions are produced by bombardment of  $N_2O$  with electrons in a storage ion source (SIS). The energy of ionizing electrons ( $E_e$ ) is determined by the accelerating voltage applied on the electron-emitting filament in the ion source and its influence on the internal excitation of the formed  $O^+$  ions is discussed in subsection II C. Ions are periodically extracted from the ion source, mass-selected by a quadrupole mass filter and injected into the ion trap via a quadrupole bender (see schematic drawings in Figures 1 and 2 in Ref. 20, and in Fig. 1 in Ref. 24). The volume of the trap is filled with He buffer gas with a small admixture of reactant gas ( $H_2$ , HD, or  $D_2$ ). In the text and figures, the number densities of  $H_2$ , HD,  $D_2$  and He gases in the trap are denoted as  $[H_2]$ ,  $[HD]$ ,  $[D_2]$ , and  $[He]$ , respectively. Typical He number density in the trap is in the range of  $10^{13} - 10^{14} \text{ cm}^{-3}$ , the number density of the reactant gas was varied from  $10^{10}$  up to  $10^{12} \text{ cm}^{-3}$ . Because of the high relative number density of helium buffer gas, the thermalization of translational energy in collisions with He proceeds at a much higher rate than collisions with reactant molecules.

After a preselected trapping time, the ions are extracted from the ion trap and after passing through the second quadrupole mass spectrometer, the mass-selected ions are counted by a multichannel plate detector (MCP). The efficiency of the detection system (extraction from the ion trap, mass selection, and detection), which is dependent on the mass of the ion (mass discrimination), is considered in data analysis. The relative detection efficiencies of oxygen hydride and deuteride ions with respect to  $O^+$  were treated as free parameters in the studies of reactions of  $O^+$  ions with  $H_2$  or  $D_2$ , respectively. These fitted detection efficiencies were used in the analysis of the  $O^+ + HD$  data. The difference between detection efficiencies for  $O^+$  and  $D^+$  ions was measured by using a calibration reaction of  $D^+$  with  $CH_4$  (for details and further references, see Ref.<sup>1</sup>). The measured relative numbers of ions presented in this paper are corrected for the mass discrimination.

By measuring the time evolution of the numbers of primary and product ions in the trap, the reaction rate coefficient of primary reaction can be determined for the given temperature and reactant gas number density.

At this point, we should comment on the internal excitation of reactant molecules. Molecules  $H_2$  and  $D_2$  can be in a para or ortho nuclear spin configurations, which are coupled with rotational states of molecules. As a result, some transitions are forbidden and the thermalization of rotational states is slow. In the present experiments, we are using normal hydrogen and normal deuterium gases, where the populations of ortho and para states correspond to thermal equilibrium at 300 K. The populations of ortho and para states in thermal equilibrium at 300 K is close to the statistical ratio of 1/3 for para- $H_2$ /ortho- $H_2$  and 1/2 for para- $D_2$ /ortho- $D_2$ . In our setup, we are using stainless steel gas handling system and we observed that para- or ortho- $H_2$  populations do not change while passing the reactant gas from a reservoir into the ion trap volume<sup>21,25</sup>. Because of spin-forbidden transitions, the rota-

tional states of  $H_2$  (or  $D_2$ ) are thermalized only within nuclear spin manifolds<sup>21,25</sup>. In the case of HD molecule, there are no nuclear spin forbidden transitions and the thermal relaxation of rotational states is fast. Having in mind that in experiments we are using normal  $H_2$  and normal  $D_2$ , we will use the designation normal only to stress special features, otherwise we will use simple names hydrogen and deuterium ( $H_2$  and  $D_2$ ).

### B. Collision temperature

The 22-pole trap is mounted within a scattering cell that is attached to the cold head of a closed-cycle helium refrigerator. The temperature of the scattering cell ( $T_{22PT}$ ) is directly measured by silicon diodes and it can reach values down to 10 K. It is safe to assume on the basis of our previous studies<sup>21-23,26</sup> that the translational temperature of He and of the reactant gas,  $T_g$ , in the volume of the trap does not exceed  $T_{22PT}$  by more than 10 K.

The translational energy of the injected ions is thermalized in collisions with atoms of helium buffer gas filling the trap volume. For a particular construction of the ion trap and experimental conditions, the translational energy distribution of the trapped ions can be characterized by measuring the Doppler broadening of their absorption lines<sup>27,28</sup>. We will assume that the translational energy distribution of the trapped ions is close to Maxwell-Boltzmann distribution and can be characterized by translational temperature  $T_t$ .

The collisional temperature  $T$  is a mass-weighted average of ion and gas translational temperatures. Several recent experiments with linear 22-pole RF ion trap instruments<sup>21,22,26,29,30</sup> have shown that the collisional temperature can be slightly higher than the trap temperature,  $T_{22PT}$ . Based on these experimental studies, we can safely assume that in the present study, the collisional temperature in the interaction of  $O^+$  ions with HD and  $D_2$  does not exceed  $T_{22PT}$  by more than 10 K. For the presentation of our data, we define the collisional temperature as  $T = T_{22PT} + 5 \text{ K}$  with an uncertainty of  $\pm 5 \text{ K}$ .

### C. Internal excitation of $O^+$ ions

In several previous ion trap experiments with  $O^+$  ions, it was observed that trapped  $O^+$  ions can be in the ground-state  $^4S$  or in one of the two excited metastable states ( $^2D$  or  $^2P$ ). The radiative lifetimes of ions in  $^2D$  or  $^2P$  states are larger than  $5.6 \times 10^3 \text{ s}$  and  $4.9 \text{ s}$ , respectively<sup>31,32</sup>. The quenching of excited metastable ions  $O^+(^2D)$  and  $O^+(^2P)$  (here simply denoted as  $O^+(^2D,^2P)$ ) in collisions with He atoms is inefficient. E.g., in SIFT experiments at 300 K with He buffer gas (with pressure above 60 Pa), the metastable  $O^+(^2D,^2P)$  ions live for many milliseconds and survive thousands of collisions with atoms of He buffer gas without de-excitation<sup>33</sup>.

Due to the presence of  $O^+(^4S)$  and  $O^+(^2D,^2P)$  ions in the ion trap, the loss rate of  $O^+$  ions due to a reaction with reactant gas can depend on the actual fraction of excited ions, which is primarily given by the used source gas and by the

energy of electrons used in the ion source. The relative numbers of  $O^+(^4S)$  and  $O^+(^2D,^2P)$  ions in the ion trap can be measured by monitoring the production of  $NO^+$  and  $N_2^+$  ions in reaction of trapped  $O^+$  ions with  $N_2$  (see similar studies in Refs. 1, 33–37). In our previous study<sup>1</sup> of the reaction of  $O^+$  with  $H_2$ , we measured the dependence of the population of the excited metastable  $O^+(^2D,^2P)$  ions in the ion trap on the energy of ionizing electrons,  $E_e$ . From these measurements, we concluded that for  $E_e < 50$  eV, the fraction ( $F$ ) of excited metastable  $O^+(^2D,^2P)$  ions in the ion trap is below 5% (see Figure 2 in Ref. 1). In the present experiments, we used the same experimental configuration and the same conditions in the ion source. Therefore, we can assume the same fraction of excited metastable  $O^+(^2D,^2P)$  ions in the ion trap.

To verify that  $F < 5\%$  for  $E_e < 50$  eV, we studied the production rate of  $D^+$  ions in the reaction of  $O^+$  ions with  $D_2$  as a function of  $E_e$ . The production of  $D^+$  (and  $D_2^+$ ) ions is a clear indication of the presence of excited  $O^+(^2D,^2P)$  ions in the trap, as it is endothermic for the ground state  $O^+(^4S)$  ions but exothermic for the excited metastable  $O^+(^2D,^2P)$  ions<sup>4</sup>. The  $D_2^+$  ions are converted to  $D_3^+$  in further reactions with  $D_2$ . Therefore, we prefer to monitor the production of  $D^+$  ions, which do not react with  $D_2$  (very slow association can be neglected), despite the favorable branching ratio<sup>4</sup> for the production of  $D_2^+$ .

A typical result obtained for trapping of  $O^+$  ions in a He/ $D_2$  mixture at  $T = 54$  K is shown in panel (a) of Figure 1. The numbers of detected ions ( $n_X(t)$ ) as a function of trapping time  $t$  are normalized by  $n_\Sigma(t_0)$ , which is the number of all ions detected in the first measurement after the injection of  $O^+$  ions into the trap (at trapping time  $t_0$ ). The figure shows the time evolutions of normalized numbers of  $O^+$ ,  $OD^+$ ,  $D_2O^+$ , and  $D_3O^+$  ions. We can see that a sequence of reactions with  $D_2$  leads to the consequent formation of  $OD^+$ ,  $D_2O^+$ , and finally  $D_3O^+$ . Note that the decrease of the normalized number of  $O^+$  ions is mono-exponential over nearly four orders in magnitude.

The lower panel of Figure 1 shows the time evolutions of normalized numbers of primary  $O^+$  and product  $D^+$  ions measured at several electron energies  $E_e$ ; in these particular experiments,  $E_e$  was varied from 20 up to 150 eV. The measured loss rate of  $O^+$  ions is almost independent on  $E_e$ , whereas the measured fraction,  $F_{D^+}$ , of produced  $D^+$  ions with respect to all ions is increasing with increasing  $E_e$ .

The fraction  $F_{D^+}$  is quantitatively evaluated from the fit of the time evolution of the numbers of  $O^+$  and  $D^+$  ions. The obtained dependence of  $F_{D^+}$  on  $E_e$  shown in Figure 2 is consistent with results obtained from probing the population of  $O^+(^2D,^2P)$  ions in the trap with  $N_2$  and from monitoring the production of  $H^+$  in the reaction of  $O^+$  with  $H_2$  (for further details, see Refs. 1 and 33). Based on the present and previous studies<sup>1</sup>, we can conclude that over 95% of trapped  $O^+$  ions are in the ground  $O^+(^4S)$  state if  $E_e < 50$  eV.

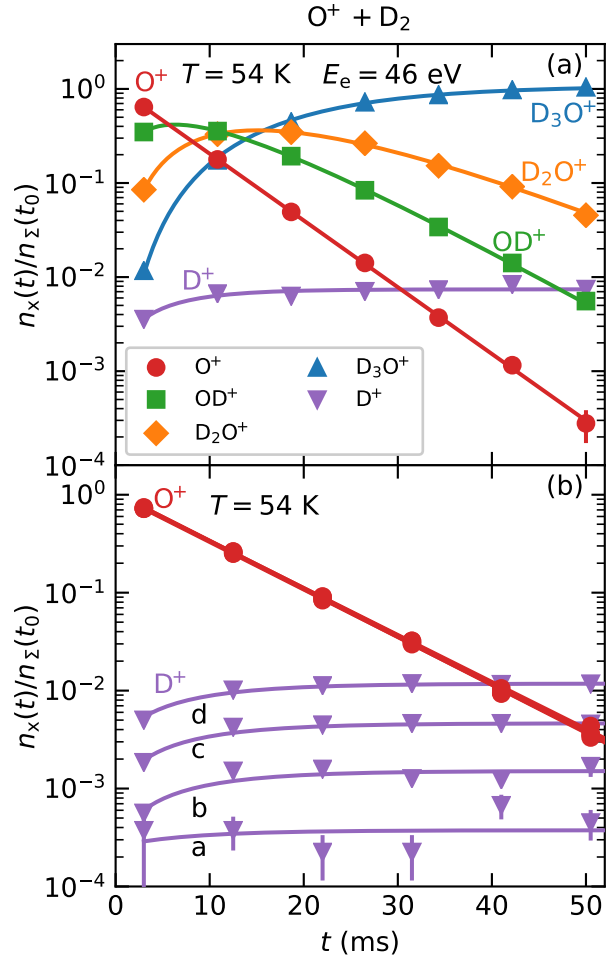


Figure 1. Measured time evolutions of normalized numbers ( $n_X(t)/n_\Sigma(t_0)$ ) of the indicated ions after injection of  $O^+$  ions into the trap filled with He buffer gas and  $D_2$  reactant gas (normal  $D_2$ ). Panel (a) – the time evolutions measured at  $T = 54$  K with  $[D_2] = 1.6 \times 10^{11} \text{ cm}^{-3}$ ,  $[He] = 2.6 \times 10^{13} \text{ cm}^{-3}$  and electron energy in the ion source  $E_e = 46$  eV. The lines indicate the results of the fit to the data. The corresponding reaction rate coefficient is  $k_2(54 \text{ K}) = (1.03 \pm 0.22) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ . Panel (b) – the time evolutions of the normalized numbers of primary  $O^+$  ions and produced  $D^+$  ions measured at several electron energies  $E_e$  in the range from 20 up to 150 eV. The letters a, b, c, and d indicate the values measured at an electron energy of 20, 24, 35, and 120 eV, respectively. The data were measured at  $T = 54$  K with  $[D_2] = 9.2 \times 10^{10} \text{ cm}^{-3}$  and  $[He] = 2.4 \times 10^{13} \text{ cm}^{-3}$ . Obtained reaction rate coefficient  $k_2(54 \text{ K}) = (1.21 \pm 0.25) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  is independent of the electron energy.

#### D. Systematic uncertainties

The overall uncertainty of the measured reaction rate coefficients consists of the following components: 1) the statistical uncertainty,  $\sigma_{\text{stat}}$ , due to the finite number of observed ions; 2) the error arising from the uncertainty in collisional and reactant gas temperatures,  $\sigma_T$ , which is dominant for tem-



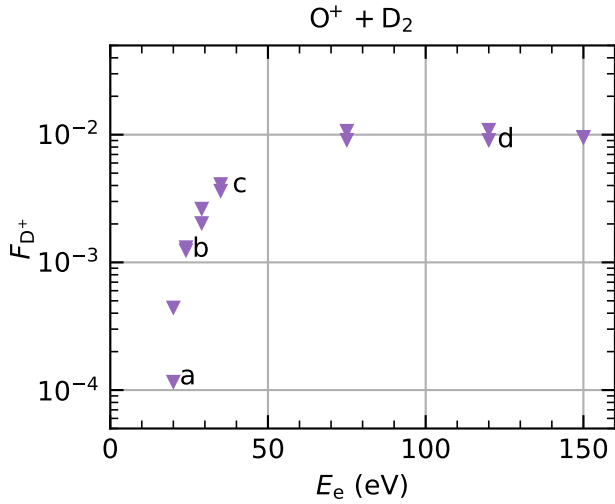


Figure 2. The dependence of the fraction  $F_{D^+}$  of the detected  $D^+$  ions on electron energy  $E_e$  used in the ion source. The letters a, b, c, and d indicate the values obtained at electron energy of 20, 24, 35, and 120 eV, respectively and correspond to the appropriate data plotted in the lower panel of Figure 1

peratures below 30 K; 3) the 20% systematic error  $\sigma_{\text{sys}}$  given by the uncertainty in reactant number density determination. The error bars of the obtained reaction rate coefficients displayed in Figures 4, 7, and 9 below include statistical error  $\sigma_{\text{stat}}$  and the temperature error  $\sigma_T$ . The constant systematic error  $\sigma_{\text{sys}}$  is not shown for clarity. Hence, the error bars indicate the relative uncertainty of our data. The errors of the reaction rate coefficients quoted throughout the text include all the mentioned uncertainties (i.e., the absolute uncertainty). As an example, for the data obtained at 54 K plotted in the upper panel of Figure 1 we have  $\sigma_{\text{stat}} = 3 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$ ,  $\sigma_T = 7.4 \times 10^{-11} \text{ cm}^3 \text{ s}^{-1}$  and  $\sigma_{\text{sys}} = 2.0 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$  giving total error  $\sigma_{\text{total}} = 2.2 \times 10^{-10} \text{ cm}^3 \text{ s}^{-1}$ .

### III. RESULTS

#### A. The reaction of $O^+(^4S)$ ions with $D_2$

The balance equation for  $O^+$  ions in the ion trap can be written in the form:

$$\frac{dn_{O^+}}{dt} = -k_2[D_2]n_{O^+} = -r_2n_{O^+} \quad (4)$$

where  $k_2$  and  $r_2$  are the rate coefficient and the rate of the binary reaction (2), respectively. The values of  $k_2$  and  $r_2$  are related by equation  $r_2 = k_2[D_2]$ . By monitoring time evolutions of the number of  $O^+(^4S)$  ions in the ion trap at a pre-selected temperature  $T$  and given  $D_2$  and He number densities (see, e.g., Figure 1), the rate coefficient  $k_2(T)$  of the reaction (2) can be obtained. To confirm the binary character of the reaction of  $O^+(^4S)$  ions with  $D_2$  taking place in the ion trap, we measured the dependence of the reaction rate  $r_2$  on the

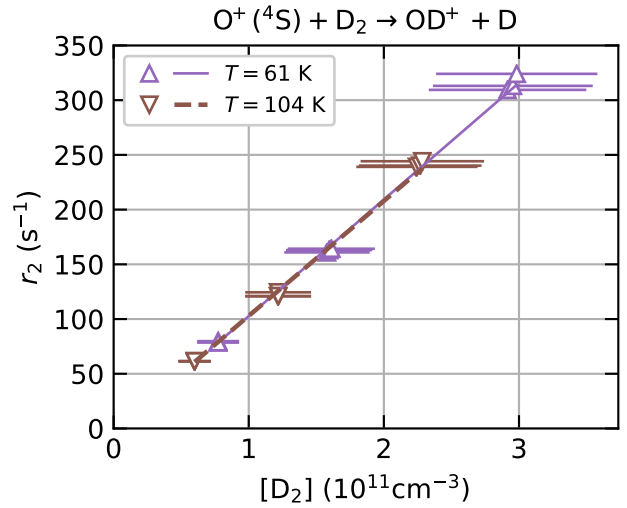


Figure 3. Dependencies of the loss rates  $r_2$  of  $O^+(^4S)$  ions on number density of  $D_2$  measured at collisional temperatures  $T = 61$  K (triangle up,  $[\text{He}] = 1.65 \times 10^{13} \text{ cm}^{-3}$ ) and  $T = 104$  K (triangle down,  $[\text{He}] = 2.15 \times 10^{13} \text{ cm}^{-3}$ ). The values of the binary reaction rate coefficient  $k_2$  obtained as the slopes of the plotted dependencies are:  $k_2(61 \text{ K}) = (1.07 \pm 0.16) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  and  $k_2(104 \text{ K}) = (1.05 \pm 0.16) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ . The displayed error bars include the systematic uncertainty of reactant number density determination that is estimated as 20 % of the measured value.

reactant density  $[D_2]$  at several temperatures and buffer gas densities. Examples of dependencies of  $r_2$  on  $[D_2]$  obtained from measurements at  $T = 61$  K and  $T = 104$  K are shown in Figure 3. In the experiments, the primary  $O^+(^4S)$  ions were produced using  $E_e = 45.5$  eV. The linearity of the obtained dependencies confirms that the time evolutions of the number of  $O^+(^4S)$  ions in the ion trap are controlled by a binary reaction with  $D_2$ . The values of the binary reaction rate coefficients  $k_2$  are given by the slope of the plotted dependencies. More precisely, we should include losses due to reactions with molecules of a background gas and write:  $r_2 = (k_2[D_2] + r_{2\text{bg}})$ . From the extrapolation of the data plotted in Figure 3 towards  $[D_2] = 0$ , we can see that the level of the background loss rate ( $r_{2\text{bg}}$ ) is very low. This conclusion is also confirmed by experiments with pure He without adding  $D_2$ . The dependence of the reaction rate  $r_2$  on  $[\text{He}]$  was also monitored in the experiments, but we did not find any significant variation when changing  $[\text{He}]$  from  $4 \times 10^{12} \text{ cm}^{-3}$  to  $4.4 \times 10^{13} \text{ cm}^{-3}$ .

By measuring the time evolutions of the numbers of  $O^+$  ions in the ion trap at temperatures from 15 up to 300 K, the temperature dependence of the rate coefficient  $k_2$  of the reaction (2) was obtained. Low energy electrons ( $E_e = 45.5$  eV) were used in the SIS to create a dominant population of the ground state  $O^+(^4S)$  ions in the ion trap (see the discussion in section II C). The obtained data and final temperature dependence of the rate coefficient  $k_2$  of the reaction of  $O^+(^4S)$  ions with  $D_2$  are shown in the panel (a) of Figure 4. See section II D for a discussion of uncertainties.

In the panel (b) of Figure 4, the temperature dependence of the reaction rate coefficient  $k_2$  obtained in the present study

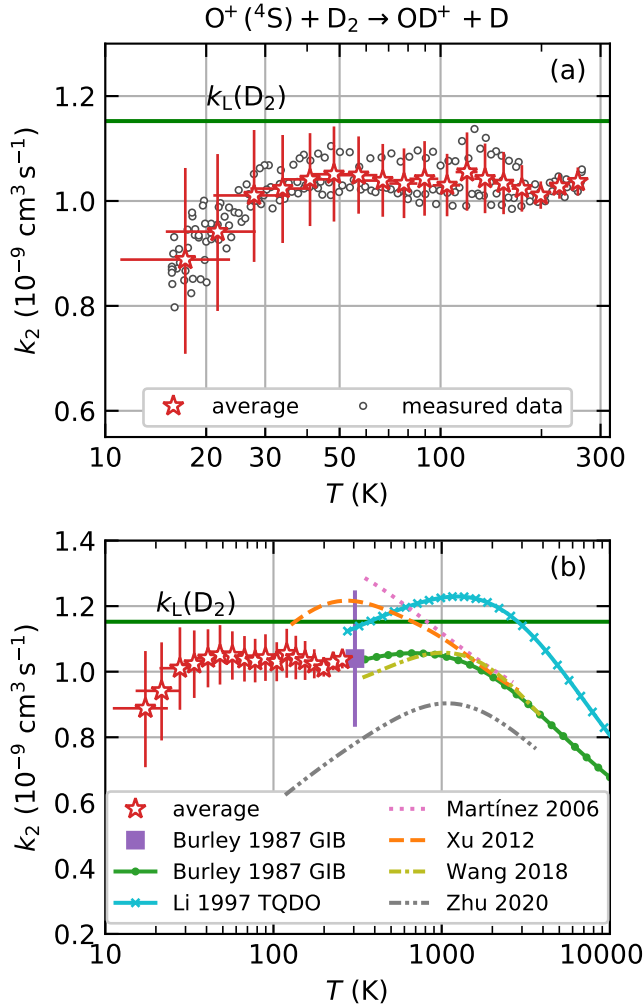


Figure 4. Temperature dependence of the reaction rate coefficient  $k_2$  of the reaction of  $O^+(^4S)$  ions with  $D_2$ . The horizontal straight lines indicate the value of the Langevin collisional rate coefficient  $k_L(D_2)$  for the reaction (2). Panel (a) – The present measured data are indicated by smaller points and the binned and averaged data are indicated by stars. Panel (b) – The comparison of the temperature dependence of  $k_2$  measured in the present study with the values from previous experiments: TQDO<sup>4</sup> and GIB<sup>38</sup>. See text for details on the calculation of thermal reaction rate coefficients from experimental cross-sections/. The value of the reaction rate coefficient reported by Burley, Ervin, and Armentrout<sup>38</sup> for 300 K is plotted as a full square. Included are also dependencies recalculated from the theoretical cross-sections obtained by Martínez *et al.*<sup>39</sup>, Xu *et al.*<sup>40</sup>, Wang *et al.*<sup>41</sup>, and by Zhu *et al.*<sup>42</sup>.

(stars) is compared with the results of previous experimental and theoretical studies. The experimental data obtained in the triple-quadrupole double-octopole experiment (TQDO) by Li *et al.*<sup>4</sup> and in the guided ion beam mass spectrometer experiment (GIB) by Burley, Ervin, and Armentrout<sup>38</sup> are also shown. To our knowledge, there are no previous experimental data for temperatures (or equivalent collision energies) below 200 K. For comparison with our data, we calculated the ther-

mal reaction rate coefficients from the published experimental cross-sections assuming Maxwell velocity distribution. In these calculations, the low- and high-energy parts of the cross-section unavailable in the experimental data of Li *et al.*<sup>4</sup> and Burley, Ervin, and Armentrout<sup>38</sup> were approximated by Langevin cross-section. If the low- and high-energy parts of the cross-section were extrapolated by a constant value, the resulting reaction rate coefficients plotted in the lower panel of Figure 4 would differ by at most 10 % in the displayed temperature range. The temperature dependencies of the reaction rate coefficients obtained by the thermal averaging of the published theoretical dependencies of the cross-sections on the collision energies are also included in the lower panel of Figure 4 (see Refs. 39, 41–43). Martínez *et al.*<sup>39</sup> calculated the cross-sections for the reaction of  $O^+(^4S)$  ions with  $D_2$  and HD for collisional energies as low as 1 meV. They note that their results show strong oscillations below 0.02 eV possibly due to a problem in the description of the reaction probability at very low energies. Therefore, we included their cross-sections only for energies higher than 0.02 eV in the thermal averaging procedure described above.

The results of various experimental and theoretical studies are in good mutual agreement for higher collisional energies corresponding to temperatures over 200 K. As mentioned earlier, there are no experimental data available for temperatures below 200 K. The value of the reaction rate coefficient for the reaction (2) obtained in the present study at 300 K is within errors of the measurement in excellent agreement with previous experimental studies of Burley, Ervin, and Armentrout<sup>38</sup> and Li *et al.*<sup>4</sup>. The calculations of Zhu *et al.*<sup>42</sup> hint at a decrease of the reaction rate coefficient at low collisional energies. However, below 300 K, the measured reaction rate coefficients are almost independent of temperature and approximately 15 % lower than the Langevin reaction rate coefficient.

## B. The reaction of $O^+(^4S)$ ions with HD

The reaction of  $O^+(^4S)$  ion with HD molecule has two reaction channels, (3a) and (3b) with products  $OH^+$  and  $OD^+$ , respectively. The corresponding reaction rate coefficients are denoted as  $k_{3a}$  and  $k_{3b}$  and the overall reaction rate coefficient is  $k_3 = (k_{3a} + k_{3b})$ . The temperature-dependent product branching ratios for the  $OH^+$  and  $OD^+$  products are defined as  $BR_{OH^+} = k_{3a}/(k_{3a} + k_{3b})$  and  $BR_{OD^+} = k_{3b}/(k_{3a} + k_{3b})$ , respectively. Note that in some publications, these quantities are named as fractions (denoted as  $f(OH^+)$  and  $f(OD^+)$ , see e.g. Ref. 44). In some publications, the branching ratio is defined as the fraction of the cross-section for channel  $OH^+$  and the sum of the cross-sections for both channels, in this case, the branching ratio is a function of collisional energy.

To obtain  $k_{3a}$  and  $k_{3b}$ , the time evolution of the numbers of primary  $O^+$  ions and  $OH^+$  and  $OD^+$  product ions in the trap has to be measured. This is complicated by subsequent reactions of the product ions with HD leading to the formation of  $H_2O^+$ ,  $HDO^+$ , and  $D_2O^+$  ions. In addition, the primary product ion  $OD^+$  and the secondary product ion  $H_2O^+$  have equal masses (18 Da) and they cannot be distinguished by the mass

spectrometer of the ion detection system. In the data analysis, we consider the balance equations for  $O^+$ ,  $OH^+$ ,  $OD^+$ , and  $H_2O^+$  ions in the ion trap.

In Figure 5, we show examples of measured time evolutions of the normalized numbers of the reactant  $O^+$  ions and of the primary and the secondary product ions. In the experiments, low electron energy ( $E_e = 37$  eV) was used in the ion source to minimize the population of the excited  $O^+(^2D, ^2P)$  ions in the ion trap (see the discussion in section II C). The data shown in the panels (a) and (b) of Figure 5 were measured during short and long trapping time, respectively, using a relatively low number density of HD. At the short trapping time, the decrease of the number of  $O^+$  ions in the ion trap is relatively small and the production of  $OH^+$  and  $OD^+$  ions is distinguishable from the formation of the secondary products. On the other hand, the decrease of the normalized number of  $O^+$  ions at long trapping times is mono-exponential through nearly three orders in magnitude, indicating that the  $O^+$  loss rate is not influenced by relaxation in the trap. The data corresponding to the secondary products of mass 19 and 20 Da are in Figure 5 labeled as  $A_{19}^+$  and  $A_{20}^+$ , respectively. The measured time evolution of the normalized number of  $H^+$  ions produced in the reaction of the excited metastable  $O^+(^2D, ^2P)$  ions with HD is also plotted in Figure 5. A similar evolution of the number of  $D^+$  ions was also observed and monitored, but it is not included in Figure 5. The small normalized number of  $H^+$  ions indicates the low fraction of the metastable  $O^+(^2D, ^2P)$  ions in the trap and their negligible influence on the measured value of the reaction rate coefficient  $k_3$  and on the measured branching ratios  $BR_{OH^+}$  and  $BR_{OD^+}$ .

To confirm the binary character of the reaction of  $O^+(^4S)$  with HD, we measured the dependence of the loss rate  $r_3$  on number density [HD] at several temperatures. Examples of such dependencies measured at 62 K and 98 K are shown in Figure 6. The linearity of both dependencies indicates that the loss of  $O^+(^4S)$  is controlled by a binary reaction with HD.

The reaction of  $O^+(^4S)$  ions with HD in the ion trap was studied at temperatures from 15 K up to 235 K. The reaction rate coefficient  $k_3$  and the branching ratios  $BR_{OH^+}$  and  $BR_{OD^+}$  were obtained by fitting the measured time evolutions of the number of  $O^+$  and  $OH^+$  ions together with the sum of numbers of  $OD^+$  and  $H_2O^+$  ions in the ion trap. The resulting reaction rate coefficients are shown in Figure 7. The measured data (small points) and the binned and averaged values of  $k_3$  (stars) are plotted in panel (a) of Figure 7. In panel (b) of Figure 7, the present results are compared with the temperature dependencies of  $k_3$  calculated by thermal averaging over Maxwell distribution from the cross-sections obtained in GIB experiments by Sunderlin and Armentrout<sup>2</sup> and Buley, Ervin, and Armentrout<sup>38</sup> and from the theoretical cross-sections<sup>39,40,45,46</sup>. Note that the calculations by Xu *et al.*<sup>40</sup> use the same potential energy surface as those by Martínez *et al.*<sup>39</sup> but include Coriolis coupling effects.

The branching ratios  $BR_{OH^+}$  and  $BR_{OD^+}$  were obtained from measurements at low density of HD in the ion trap ( $\approx 10^{10}$  cm<sup>-3</sup>) and short storage times (see Figure 5), where the time evolution of numbers of  $OH^+$  and  $OD^+$  product ions in the ion trap is dominated by their production in reactions

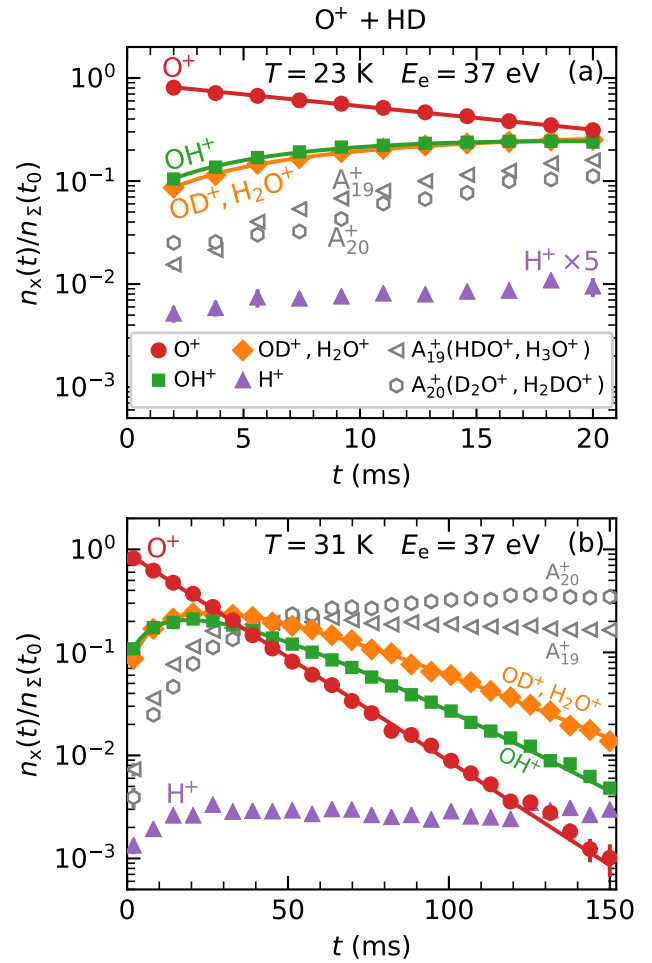


Figure 5. Measured time evolutions of the normalized numbers ( $n_X(t)/n_{\Sigma}(t_0)$ ) of the indicated ions after the injection of  $O^+$  ions into the ion trap filled with He buffer gas and HD reactant gas. Panel (a) – short trapping time – the time evolutions of the normalized numbers of the primary  $O^+$  ions and of the produced ions measured at  $T = 23$  K,  $[HD] = 4.1 \times 10^{10}$  cm<sup>-3</sup>,  $[He] = 3.9 \times 10^{13}$  cm<sup>-3</sup>, and  $E_e = 37$  eV. The values plotted for  $H^+$  ions are multiplied by a factor of 5. The lines indicate the results from the fits (see the explanation in the text). The obtained reaction rate coefficient is  $k_3(23 \text{ K}) = (1.28 \pm 0.31) \times 10^{-9}$  cm<sup>3</sup> s<sup>-1</sup>. Panel (b) – long trapping time – the time evolutions of the normalized numbers of the primary  $O^+$  ions and of the produced ions measured at  $T = 31$  K,  $[HD] = 3.0 \times 10^{10}$  cm<sup>-3</sup>,  $[He] = 3.4 \times 10^{13}$  cm<sup>-3</sup> and  $E_e = 37$  eV. The obtained reaction rate coefficient is  $k_3(31 \text{ K}) = (1.54 \pm 0.35) \times 10^{-9}$  cm<sup>3</sup> s<sup>-1</sup>.

(3a) and (3a), respectively. The resulting values of  $BR_{OH^+}$  are shown in Figure 8. The  $H_2$  impurity in HD gas, declared by the gas supplier to be less than 3 % would change the value of  $BR_{OH^+}$  by at most 1.5% (depending on the actual amount of  $H_2$  in the used HD gas) and this uncertainty is included in the overall error of the  $BR_{OH^+}$  values plotted in Figure 8. The values obtained in the present study are compared with data from previous experiments (GIB<sup>2,38</sup> and VT-SIFDT<sup>44</sup>) and with calculated (theoretical) temper-

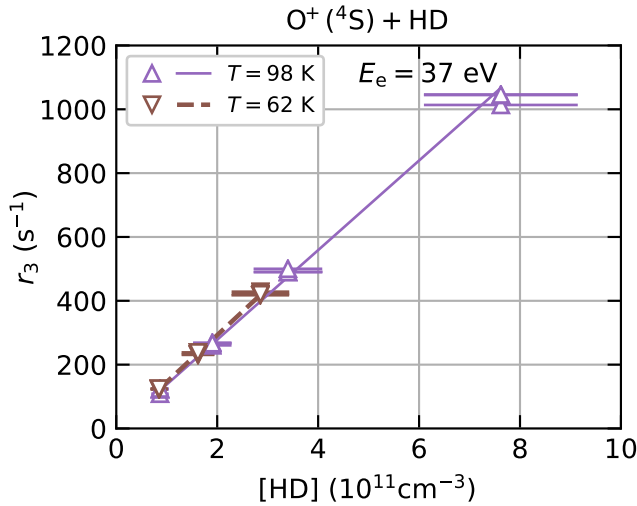


Figure 6. Measured dependencies of the reaction rate  $r_3$  of  $O^+(^4S)$  ions on the number density  $[HD]$  at collisional temperatures  $T = 62$  K and  $T = 98$  K (with  $[He] = 2.23 \times 10^{13} \text{ cm}^{-3}$  and  $[He] = 1.78 \times 10^{13} \text{ cm}^{-3}$  respectively). The displayed error bars include the systematic uncertainty of reactant number density determination that is estimated as 20 % of the measured value. The reaction rate coefficients obtained as the slope of the plotted dependencies are  $k_3(62 \text{ K}) = (1.47 \pm 0.23) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  and  $k_3(98 \text{ K}) = (1.40 \pm 0.14) \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$ .

ature dependencies<sup>39,40,45,46</sup>. As can be seen from figure 8, the present values are in good agreement with previous experimental data, which were obtained at temperatures/collision energies above  $\approx 100$  K. To our best knowledge, there are no previous experimental data for temperatures below 100 K.

#### IV. DISCUSSION AND SUMMARY

##### A. The role of metastable $O^+(^2D)$ and $O^+(^2P)$ ions

By using low energy electrons ( $E_e < 50$  eV) in the ion source (SIS) we obtained a dominant population of the ground-state ions  $O^+(^4S)$  and only very low population ( $< 5$  %) of excited metastable  $O^+(^2D, ^2P)$  ions. The fraction of  $O^+(^2D, ^2P)$  ions in the ion trap was monitored by measuring the production of  $D^+$  and  $H^+$  ions in the reaction of  $O^+(^2D, ^2P)$  ions with  $D_2$  and  $HD$ , respectively (see Figure 1 and Figure 5). The obtained dependence of the population of  $O^+(^2D, ^2P)$  on  $E_e$  is consistent with results<sup>1</sup> obtained from the probing of the population of  $O^+(^2D, ^2P)$  ions in the ion trap by the reaction with  $N_2$ . Because the studied reactions of  $O^+(^4S)$  ions with  $D_2$  and  $HD$  are fast, a small relative number of the metastable  $O^+(^2D, ^2P)$  ions ( $< 5$  %) will not influence the determination of the reaction rate coefficients, by more than  $\pm 10$  % of a value of the Langevin collisional rate coefficient. Based on the determination of the relative population of  $O^+(^2D, ^2P)$  ions in the ion trap, we concluded that the reaction rate coefficients obtained from the measured decays of numbers of  $O^+$  ions in the ion trap are given predominantly

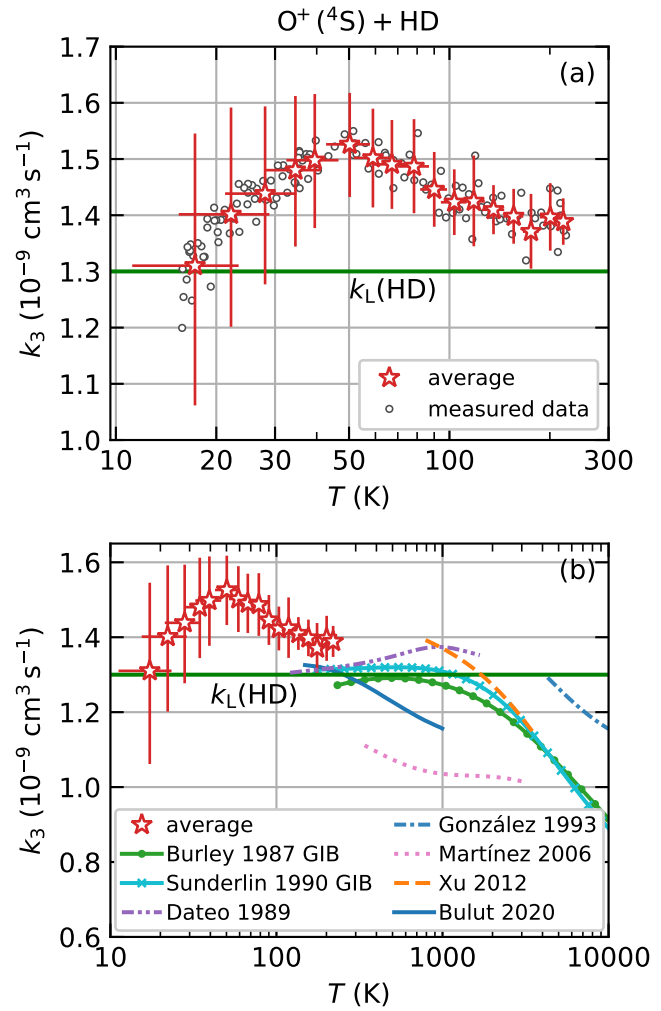


Figure 7. Temperature dependence of the rate coefficient  $k_3$  of the reaction of  $O^+(^4S)$  ions with  $HD$  molecule. Panel (a) – the present measured data are indicated with smaller points. The stars indicate the binned and averaged values. Panel (b) – the comparison of the temperature dependence of  $k_3$  obtained in the present study with the values evaluated from previous GIB experiments by Sunderlin and Armentrout<sup>2</sup> and Burley, Ervin, and Armentrout<sup>38</sup> and from the theoretical cross-sections calculated by Dateo and Clary<sup>45</sup>, González *et al.*<sup>46</sup>, Martínez *et al.*<sup>39</sup>, and Xu *et al.*<sup>40</sup>. Viggiano *et al.*<sup>44</sup> reported value of  $1.2 \times 10^{-9} \text{ cm}^3 \text{ s}^{-1}$  with 40% uncertainty at 93, 300, and 509 K. For details on the evaluation of reaction rate coefficients from theoretical and experimental cross-sections see section III A. The horizontal straight line ( $k_L(HD)$ ) in both panels indicates the value of the Langevin collisional rate coefficient for the reaction (3).

by the reaction of ground-state ions  $O^+(^4S)$ . The presented branching ratios are given by the ratios of the corresponding reaction rate coefficients and, as such, are also pertaining to the reaction of ground state  $O^+(^4S)$  ions with  $HD$ .

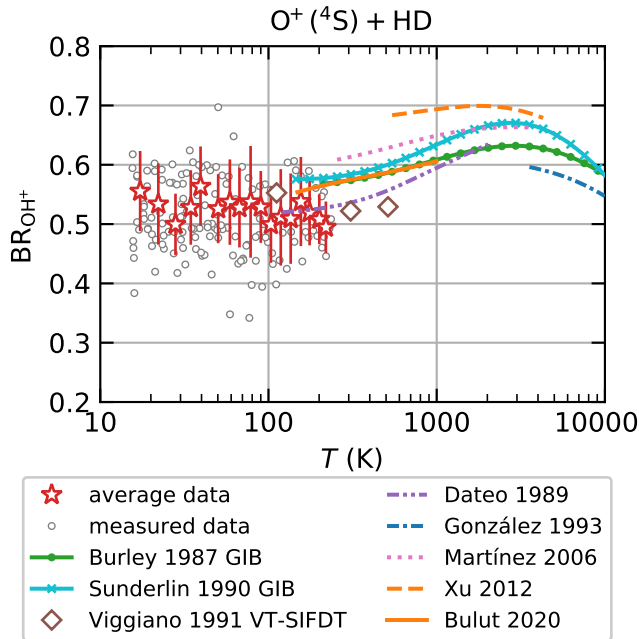


Figure 8. Product branching ratio  $BR_{OH^+} = k_{3a}/k_3$  for the production of  $OH^+$  ions in the reaction of  $O^+(^4S)$  with HD. The measured data are indicated with the smaller points; the stars indicate the binned and averaged values. The present data are compared with experimental data obtained in GIB (Burley, Ervin, and Armentrout<sup>38</sup>, Sunderlin and Armentrout<sup>2</sup>) and VT-SIFDT (Viggiano *et al.*<sup>44</sup>) experiments and with calculated (theoretical) temperature dependencies (Dateo and Clary<sup>45</sup>, González *et al.*<sup>46</sup>, Martínez *et al.*<sup>39</sup>, Xu *et al.*<sup>40</sup>). The rate coefficients  $k_{3a}$  and  $k_3$  were obtained from the published experimental and theoretical cross-sections by thermal averaging over Maxwell distribution. For details, see section III A. Only the low energy values obtained by Viggiano *et al.*<sup>44</sup> at each particular gas temperature are plotted.

### B. The temperature dependencies of measured reaction rate coefficients

To summarize the obtained results and to show the observed effect of isotope exchange, the temperature dependencies of the rate coefficients of the reactions of  $O^+(^4S)$  with HD, and with normal  $D_2$  are plotted in Figure 9. The temperature dependence of the rate coefficient of the reactions of  $O^+(^4S)$  with normal  $H_2$  measured in our previous study<sup>1</sup> is also included in the figure. The dependencies are plotted for temperature from 15 up to 300 K. We can see that the reaction rate coefficients are nearly constant with values close to the values of the corresponding Langevin collisional rate coefficients. The differences between the reaction rate coefficients and corresponding Langevin collisional rate coefficients are within the absolute accuracy of the experimental data.

We can see a small decrease in the rate coefficients towards low temperature, but this is within the relative accuracy of experimental data (as indicated by the error bars). The low value of  $k_2$  can be coupled to the lower value of the collisional rate coefficient  $k_L(D_2)$ . The values obtained for  $k_3$  are slightly higher than  $k_L(HD)$  but again, the difference is within

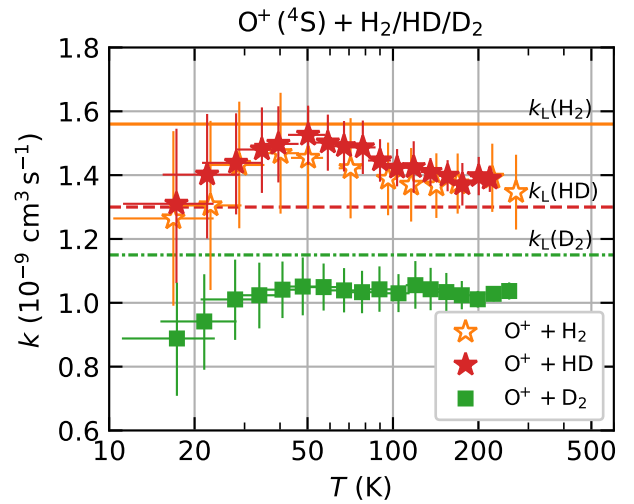


Figure 9. Measured temperature dependencies of the rate coefficients of reactions of ground state  $O^+(^4S)$  ions with HD, normal  $D_2$ , and normal  $H_2$ . The horizontal straight lines show the values of indicated Langevin collisional rate coefficients. The data for  $H_2$  are adopted from our previous study<sup>1</sup>.

the accuracy of the measured data. The comparison of the present results with those obtained at 300 K in previous studies shows very good agreement (see Figure 4 and Figure 7 above and Figure 5 in Ref. 1). We cannot make a comparison at lower temperatures because the present study is the first systematic experimental study of all three reactions at temperatures below 100 K. The temperature dependence of the measured branching ratio  $BR_{OH^+}$  for production of  $OH^+$  in the reaction of  $O^+(^4S)$  with HD is also in good agreement with previous values measured at temperatures down to  $\approx 105$  K (see Refs. 2 and 44) and at 300 K (Ref. 38).

### V. CONCLUSION AND OUTLOOK

We have studied the reactions of  $O^+$  ions in the electronic ground state ( $^4S$ ) with  $D_2$  and HD using the cryogenic 22-pole RF ion trap and we measured the temperature dependencies of the reaction rate coefficients at temperatures from 300 K down to 15 K. For the reaction of  $O^+(^4S)$  ions with HD, we also measured the isotopic branching ratio for the production of  $OH^+$  and  $OD^+$  ions.

The rate coefficients measured for the reaction of  $O^+(^4S)$  ions with HD and  $D_2$  are, considering the experimental errors, independent of temperature in the covered temperature range and close to the corresponding Langevin collisional reaction rate coefficients. We previously obtained similar results<sup>1</sup> for the reaction of  $O^+(^4S)$  ion with  $H_2$ . This is the first study giving reaction rate coefficients for the titular reactions for temperatures below 100 K. The present data are in good agreement with previous experimental and theoretical studies in the overlapping temperature ranges.

The obtained isotopic branching ratio for the formation of  $OH^+$  and  $OD^+$  ions in the reaction of  $O^+(^4S)$  with HD agrees

very well with the results previous experimental studies performed at higher collisional energies and also with the results recent calculations. No substantial temperature dependence of the branching ratio was observed in the covered temperature range.

We hope that our results will help to improve the models of interstellar chemistry and could serve as a benchmark for theoretical calculations.

## ACKNOWLEDGMENTS

We thank the Technische Universität Chemnitz and the Deutsche Forschungsgemeinschaft for lending the 22-pole trap instrument to the Charles University. We thank prof. Dieter Gerlich for discussion and helpful suggestions. We thank prof. M. González and R. Martínez for providing us with tabulated cross-sections from Ref. 39. This work was partly supported by the Czech Science Foundation (GACR 20-22000S), and by the Charles University (project Nr. GAUK1162920).

....

## REFERENCES

- <sup>1</sup>A. Kovalenko, T. D. Tran, S. Rednyk, Š. Roučka, P. Dohnal, R. Plašil, Dieter Gerlich, and J. Glošik, "OH + Formation in the Low-temperature O + (<sup>4</sup>S) + H<sub>2</sub> Reaction," *Astrophys. J.* **856**, 100 (2018).
- <sup>2</sup>L. S. Sunderlin and P. B. Armentrout, "Temperature dependence of the reaction of O<sup>+</sup> with HD," *Chem. Phys. Lett.* **167**, 188–192 (1990).
- <sup>3</sup>R. T. Wiedmann, R. G. Tonkyn, M. G. White, K. Wang, and V. McKoy, "Rotationally resolved threshold photoelectron spectra of OH and OD," *J. Chem. Phys.* **97**, 768–772 (1992).
- <sup>4</sup>X. Li, Y. L. Huang, G. D. Flesch, and N. Ng C Y, C., "Absolute state-selected total cross sections for the ion–molecule reactions O+(<sup>4</sup>S,2D,2P)+H<sub>2</sub>(D<sub>2</sub>)," *J. Chem. Phys.* **106**, 564–571 (1997).
- <sup>5</sup>M. W. Chase and National Institute of Standards and Technology (U.S.), *NIST-JANAF Thermochemical Tables* (American Chemical Society; American Institute of Physics for the National Institute of Standards and Technology, [Washington, D.C.]; Woodbury, N.Y., 1998).
- <sup>6</sup>J. E. Sansonetti and W. C. Martin, "Handbook of Basic Atomic Spectroscopic Data," *J. Phys. Chem. Ref. Data* **34**, 1559–2259 (2005).
- <sup>7</sup>J. Liu, E. J. Salumbides, U. Hollenstein, J. C. J. Koelemeij, K. S. E. Eikema, W. Ubachs, and F. Merkt, "Determination of the ionization and dissociation energies of the hydrogen molecule," *J. Chem. Phys.* **130**, 174306 (2009).
- <sup>8</sup>T. D. Tran, S. Rednyk, A. Kovalenko, Š. Roučka, P. Dohnal, R. Plašil, Dieter Gerlich, and J. Glošik, "Formation of H<sub>2</sub>O<sup>+</sup> and H<sub>3</sub>O<sup>+</sup> Cations in Reactions of OH + and H<sub>2</sub>O + with H<sub>2</sub>: Experimental Studies of the Reaction Rate Coefficients from T = 15 to 300 K," *Astrophys. J.* **854**, 25 (2018).
- <sup>9</sup>S. S. Kumar, F. Grussie, Y. V. Suleimanov, H. Guo, and H. Kreckel, "Low temperature rates for key steps of interstellar gas-phase water formation," *Sci. Adv.* **4**, eaar3417 (2018).
- <sup>10</sup>M. J. Jensen, R. C. Bilodeau, C. P. Safvan, K. Seiersen, L. H. Andersen, H. B. Pedersen, and O. Heber, "Dissociative Recombination of H<sub>3</sub>O<sup>+</sup>, HD<sub>2</sub>O<sup>+</sup>, and D<sub>3</sub>O<sup>+</sup>," *ApJ* **543**, 764 (2000).
- <sup>11</sup>N. Indriolo, D. A. Neufeld, M. Gerin, P. Schilke, A. O. Benz, B. Winkel, K. M. Menten, E. T. Chambers, J. H. Black, S. Bruderer, E. Falgarone, B. Godard, J. R. Goicoechea, H. Gupta, D. C. Lis, V. Ossenkopf, C. M. Persson, P. Sonnentrucker, F. F. S. van der Tak, E. F. van Dishoeck, M. G. Wolfire, and F. Wyrowski, "Herschel Survey of Galactic OH<sup>+</sup>, H<sub>2</sub>O<sup>+</sup>, and H<sub>3</sub>O<sup>+</sup>: Probing the Molecular Hydrogen Fraction and Cosmic-Ray Ionization Rate," *Astrophys. J.* **800**, 40 (2015).
- <sup>12</sup>T. Albertsson, D. A. Semenov, A. I. Vasyunin, T. Henning, and E. Herbst, "New Extended Deuterium Fractionation Model: Assessment at Dense ISM Conditions and Sensitivity Analysis," *Astrophys. J. Suppl. Ser.* **207**, 27 (2013).
- <sup>13</sup>B. Parise, F. Du, F.-C. Liu, A. Belloche, H. Wiesemeyer, R. Güsten, K. M. Menten, H.-W. Hübers, and B. Klein, "Detection of OD towards the low-mass protostar IRAS 16293-2422," *Astron. Astrophys.* **542**, L5 (2012).
- <sup>14</sup>J. K. Jørgensen and E. F. van Dishoeck, "THE HDO/H<sub>2</sub>O RATIO IN GAS IN THE INNER REGIONS OF A LOW-MASS PROTOSTAR," *Astrophys. J.* **725**, L172–L175 (2010).
- <sup>15</sup>D. Gerlich and S. Horning, "Experimental investigation of radiative association processes as related to interstellar chemistry," *Chem. Rev.* **92**, 1509–1539 (1992).
- <sup>16</sup>D. Gerlich, "Ion-neutral collisions in a 22-pole trap at very low energies," *Phys. Scr.* **1995**, 256 (1995).
- <sup>17</sup>D. Gerlich, G. Borodi, A. Luca, C. Mogo, and M. A. Smith, "Reactions between Cold CH<sub>x</sub><sup>+</sup> and Slow H and H<sub>2</sub>," *Z. Für Phys. Chem.* **225**, 475–492 (2011).
- <sup>18</sup>D. Gerlich, P. Jusko, Š. Roučka, I. Zymak, R. Plašil, and J. Glošik, "Ion Trap Studies of H<sup>+</sup> + H → H<sub>2</sub> + e<sup>-</sup> between 10 and 135 K," *Astrophys. J.* **749**, 22 (2012).
- <sup>19</sup>D. Gerlich, R. Plašil, I. Zymak, M. Hejduk, P. Jusko, D. Mulin, and J. Glošik, "State Specific Stabilization of H<sup>+</sup> + H<sub>2</sub>(j) Collision Complexes," *J. Phys. Chem. A* **117**, 10068–10075 (2013).
- <sup>20</sup>I. Zymak, P. Jusko, Š. Roučka, R. Plašil, P. Rubovič, D. Gerlich, and J. Glošik, "Ternary association of H<sup>+</sup> ion with H<sub>2</sub> at 11 K, experimental study," *Eur. Phys. J. Appl. Phys.* **56**, 24010 (2011).
- <sup>21</sup>I. Zymak, M. Hejduk, D. Mulin, R. Plašil, J. Glošik, and D. Gerlich, "Low-Temperature Ion Trap Studies of N+(<sup>3</sup>P<sub>1/2</sub>) + H<sub>2</sub>(j) → NH<sup>+</sup> + H," *Astrophys. J.* **768**, 86 (2013).
- <sup>22</sup>R. Plašil, I. Zymak, P. Jusko, D. Mulin, D. Gerlich, and J. Glošik, "Stabilization of H<sup>+</sup>+H<sub>2</sub> collision complexes between 11 and 28K," *Phil. Trans. R. Soc. A* **370**, 5066–5073 (2012).
- <sup>23</sup>D. Mulin, Š. Roučka, P. Jusko, I. Zymak, R. Plašil, D. Gerlich, R. Wester, and J. Glošik, "H/D exchange in reactions of OH<sup>-</sup> with D<sub>2</sub> and of OD<sup>-</sup> with H<sub>2</sub> at low temperatures," *Phys. Chem. Chem. Phys.* **17**, 8732–8739 (2015).
- <sup>24</sup>R. Plašil, Š. Roučka, A. Kovalenko, T. D. Tran, S. Rednyk, P. Dohnal, D. Shapko, D. Gerlich, and J. Glošik, "Reaction of N<sup>+</sup> ion with H<sub>2</sub>, HD, and D<sub>2</sub> at low temperatures: Experimental study of the pathway to deuterated nitrogen-containing molecules in the interstellar medium," *Astron. Astrophys.* (2020), in preparation.
- <sup>25</sup>M. Hejduk, P. Dohnal, J. Varju, P. Rubovič, R. Plašil, and J. Glošik, "Nuclear spin state-resolved cavity ring-down spectroscopy diagnostics of a low-temperature <sup>9</sup>Brm H<sub>3</sub><sup>+</sup>-dominated plasma," *Plasma Sources Sci. Technol.* **21**, 024002 (2012).
- <sup>26</sup>Š. Roučka, S. Rednyk, A. Kovalenko, T. D. Tran, R. Plašil, Á. Kálosi, P. Dohnal, D. Gerlich, and J. Glošik, "Effect of rotational excitation of H<sub>2</sub> on isotopic exchange reaction with OD<sup>-</sup> at low temperatures," *Astron. Astrophys.* **615**, L6 (2018).
- <sup>27</sup>J. Glošik, P. Hlavenka, R. Plašil, F. Windisch, D. Gerlich, A. Wolf, and H. Kreckel, "Action spectroscopy of and D<sub>2</sub>H<sup>+</sup> using overtone excitation," *Phil. Trans. R. Soc. A* **364**, 2931–2942 (2006).
- <sup>28</sup>P. Jusko, O. Asvany, A.-C. Wallerstein, S. Brünken, and S. Schlemmer, "Two-Photon Rotational Action Spectroscopy of Cold OH<sup>-</sup> at 1 ppb Accuracy," *Phys. Rev. Lett.* **112**, 253005 (2014).
- <sup>29</sup>D. Hauser, S. Lee, F. Carelli, S. Spieler, O. Lakhmanskaya, E. S. Endres, S. S. Kumar, F. Gianturco, and R. Wester, "Rotational state-changing cold collisions of hydroxyl ions with helium," *Nat. Phys.* **11**, 467–470 (2015).
- <sup>30</sup>E. S. Endres, G. Egger, S. Lee, O. Lakhmanskaya, M. Simpson, and R. Wester, "Incomplete rotational cooling in a 22-pole ion trap," *J. Mol. Spectrosc. Molecular Spectroscopy in Traps*, **332**, 134–138 (2017).
- <sup>31</sup>C. J. Zeippen, "Improved radiative transition probabilities for O II forbidden lines," *Astronomy and Astrophysics* **173**, 410–414 (1987).
- <sup>32</sup>M. Godefroid and C. F. Fischer, "MCHF-BP fine-structure splittings and transition rates for the ground configuration in the nitrogen sequence," *J. Phys. B: At. Mol. Phys.* **17**, 681 (1984).
- <sup>33</sup>J. Glošik, A. B. Rakshit, N. D. Twiddy, N. G. Adams, and D. Smith, "Measurement of the rates of reaction of the ground and metastable excited states of O<sub>2</sub><sup>+</sup>, NO<sup>+</sup> and O<sup>+</sup> with atmospheric gases at thermal energy," *J. Phys. B: At. Mol. Phys.* **11**, 3365 (1978).
- <sup>34</sup>D. Smith, N. G. Adams, and T. M. Miller, "A laboratory study of the reactions of N<sup>+</sup>, N<sub>2</sub><sup>+</sup>, N<sub>3</sub><sup>+</sup>, N<sub>4</sub><sup>+</sup>, O<sup>+</sup>, O<sub>2</sub><sup>+</sup>, and NO<sup>+</sup> ions with several

- molecules at 300 K,” *J. Chem. Phys.* **69**, 308–318 (1978).
- <sup>35</sup>R. Johnsen and M. A. Biondi, “Charge transfer coefficients for the O+(2D)+N<sub>2</sub> and O+(2D)+O<sub>2</sub> excited ion reactions at thermal energy,” *J. Chem. Phys.* **73**, 190–193 (1980).
- <sup>36</sup>P. M. Hierl, I. Dotan, J. V. Seeley, J. M. Van Doren, R. A. Morris, and A. A. Viggiano, “Rate constants for the reactions of O<sup>+</sup> with N<sub>2</sub> and O<sub>2</sub> as a function of temperature (300–1800 K),” *J. Chem. Phys.* **106**, 3540–3544 (1997).
- <sup>37</sup>J.-L. Le Garrec, S. Carles, T. Speck, J. B. A. Mitchell, B. R. Rowe, and E. E. Ferguson, “The ion–molecule reaction of O<sup>+</sup> with N<sub>2</sub> measured down to 23 K,” *Chem. Phys. Lett.* **372**, 485–488 (2003).
- <sup>38</sup>J. D. Burley, K. M. Ervin, and P. B. Armentrout, “Translational energy dependence of O+(<sup>4</sup>S) + H<sub>2</sub>(D<sub>2</sub>, HD) → OH+(OD<sup>+</sup>) + H(D) from thermal energies to 30 eV c.m.” *Int. J. Mass Spectrom. Ion Process.* **80**, 153–175 (1987).
- <sup>39</sup>R. Martínez, J. D. Sierra, S. K. Gray, and M. González, “Time dependent quantum dynamics study of the O<sup>+</sup>+H<sub>2</sub>(v=0,j=0)→OH<sup>+</sup>+H ion-molecule reaction and isotopic variants (D<sub>2</sub>,HD),” *J. Chem. Phys.* **125**, 164305 (2006), the cross section data were kindly provided by M. González and R. Martínez.
- <sup>40</sup>W. Xu, W. Li, S. Lv, H. Zhai, Z. Duan, and P. Zhang, “Coriolis Coupling Effects in O<sup>+</sup>(<sup>4</sup>S) + H<sub>2</sub>(X<sup>1</sup>Σ<sub>g</sub><sup>+</sup>) → OH<sup>+</sup>(X<sup>3</sup>Σ<sup>-</sup>) + H(<sup>2</sup>S) Reaction and Its Isotopic Variants: Exact Time-Dependent Quantum Scattering Study,” *J. Phys. Chem. A* **116**, 10882–10888 (2012).
- <sup>41</sup>X.-L. Wang, F. Gao, S.-B. Gao, L.-L. Zhang, Y.-Z. Song, and Q.-T. Meng, “Novel potential energy surface-based quantum dynamics of ion–molecule reaction O<sup>+</sup> + D<sub>2</sub> O<sup>+</sup> + D<sub>2</sub> → OD<sup>+</sup> + D,” *Chinese Phys. B* **27**, 043104 (2018).
- <sup>42</sup>Z. Zhu, L. Li, Q. Li, and B. Teng, “Dynamics studies of O<sup>+</sup> + D<sub>2</sub> reaction using the time-dependent wave packet method,” *Mol. Phys.* **118**, 1619855 (2020).
- <sup>43</sup>R. Martínez, J. Millán, and M. González, “Ab initio analytical potential energy surface and quasiclassical trajectory study of the O+(<sup>4</sup>S)+H<sub>2</sub>(X<sup>1</sup>Σ<sub>g</sub><sup>+</sup>)→OH+(X<sup>3</sup>Σ<sup>-</sup>)+H(2S) reaction and isotopic variants,” *J. Chem. Phys.* **120**, 4705–4714 (2004).
- <sup>44</sup>A. A. Viggiano, J. M. Van Doren, R. A. Morris, J. S. Williamson, P. L. Mundis, J. F. Paulson, and C. E. Dateo, “Rotational temperature dependence of the branching ratio for the reaction of O<sup>+</sup> with HD,” *J. Chem. Phys.* **95**, 8120–8123 (1991).
- <sup>45</sup>C. E. Dateo and D. C. Clary, “Isotopic branching ratio for the O<sup>+</sup>+HD reaction,” *J. Chem. Soc., Faraday Trans. 2* **85**, 1685–1696 (1989).
- <sup>46</sup>M. González, M. Gilibert, A. Aguilar, and R. Sayós, “A comparison between experimental and theoretical excitation functions for the O<sup>+</sup> + H<sub>2</sub>(<sup>4</sup>A<sup>+</sup>) system using trajectory calculations over a wide energy range,” *J. Chem. Phys.* **98**, 2927–2935 (1993).